

What is LINC-NIRVANA?

High Resolution
NIR Imager:
MCAO, Interferometry

λ : JHK

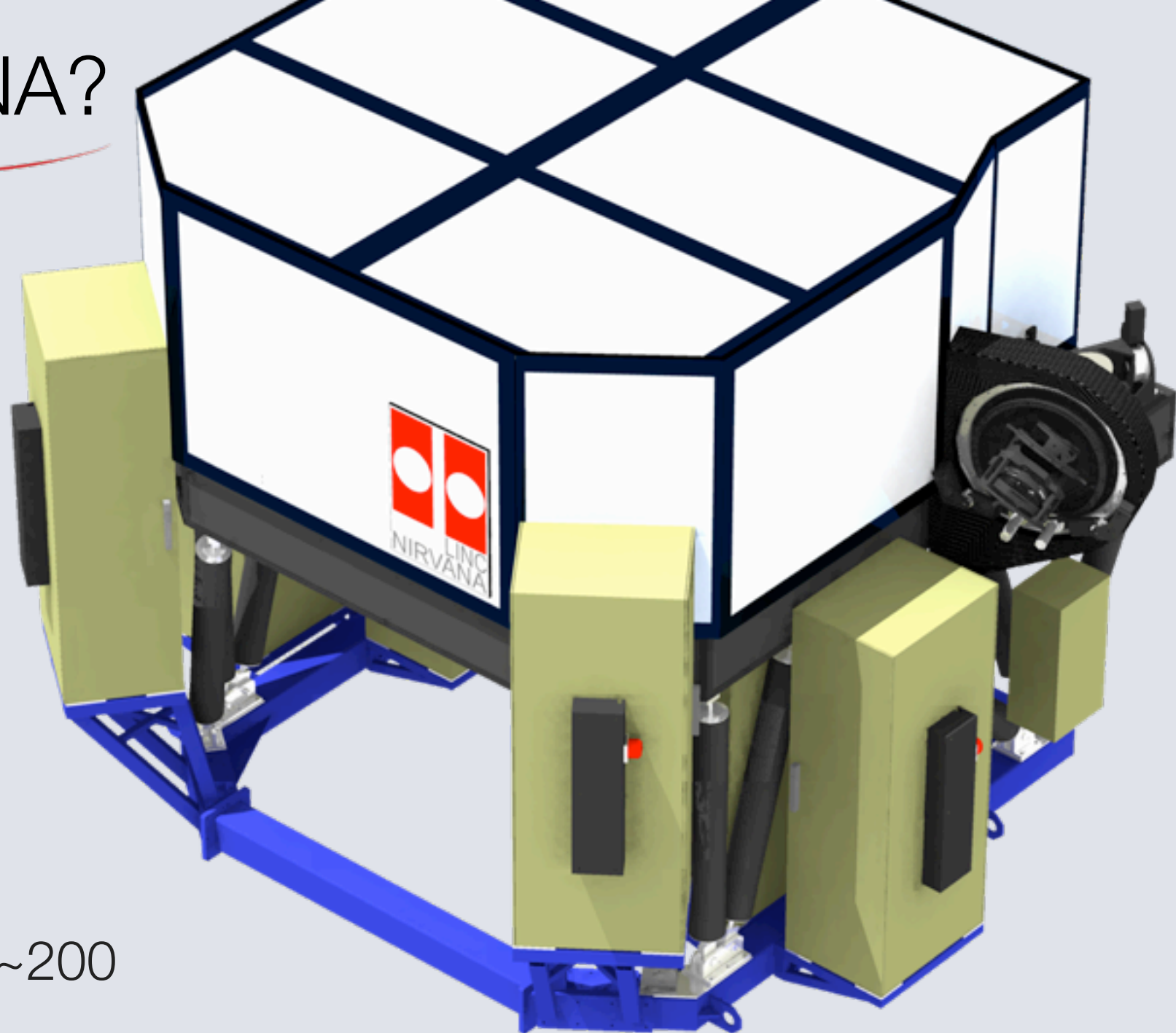
Pixel: 5.11 mas

$\lambda/23m$: 10 mas (J)

15 mas (H)

20 mas (K)

R: Broadband to R~200

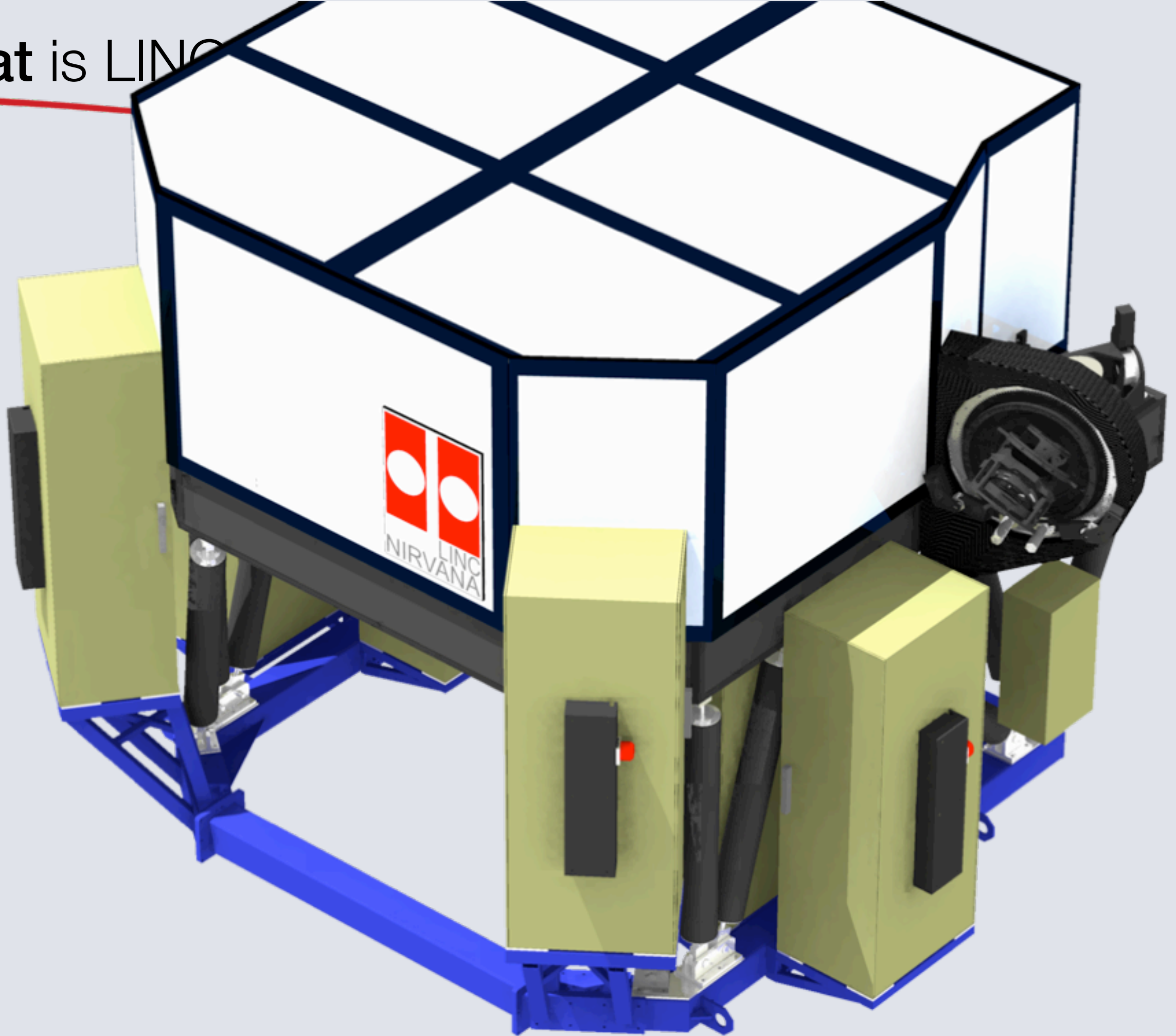


An International Project

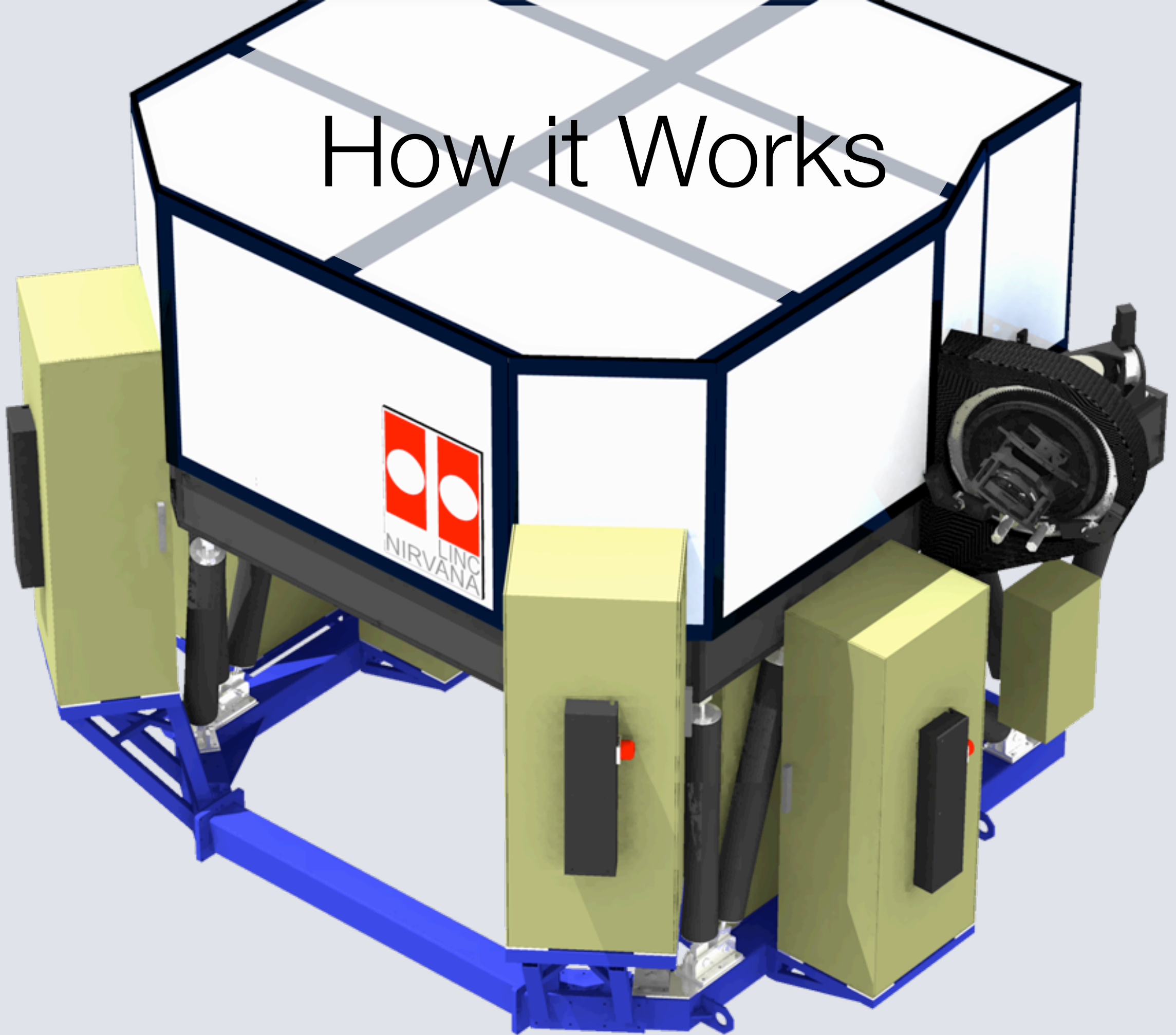


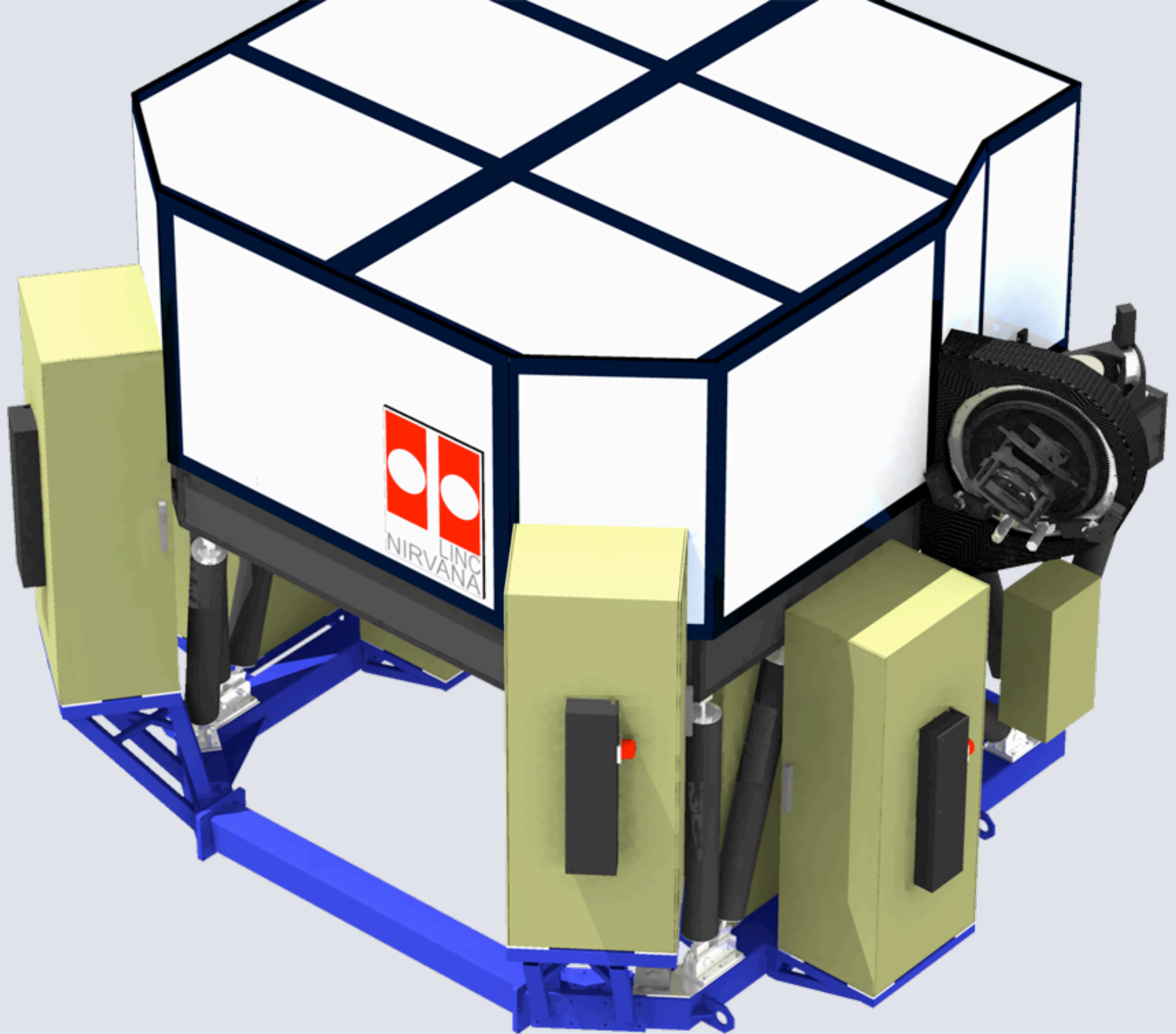
MPIA Heidelberg • INAF • University of Cologne • MPIfR Bonn

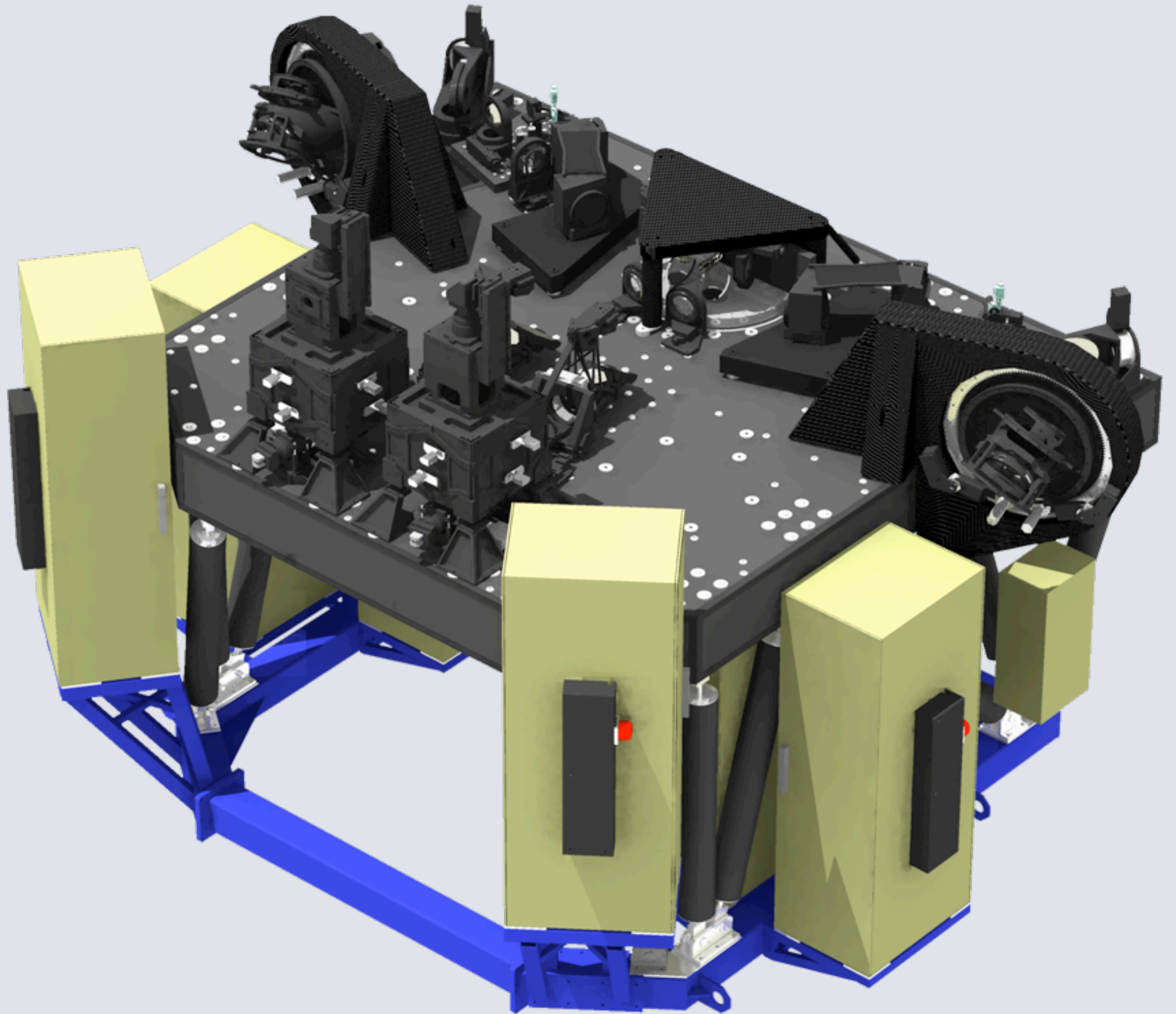
What is LINC

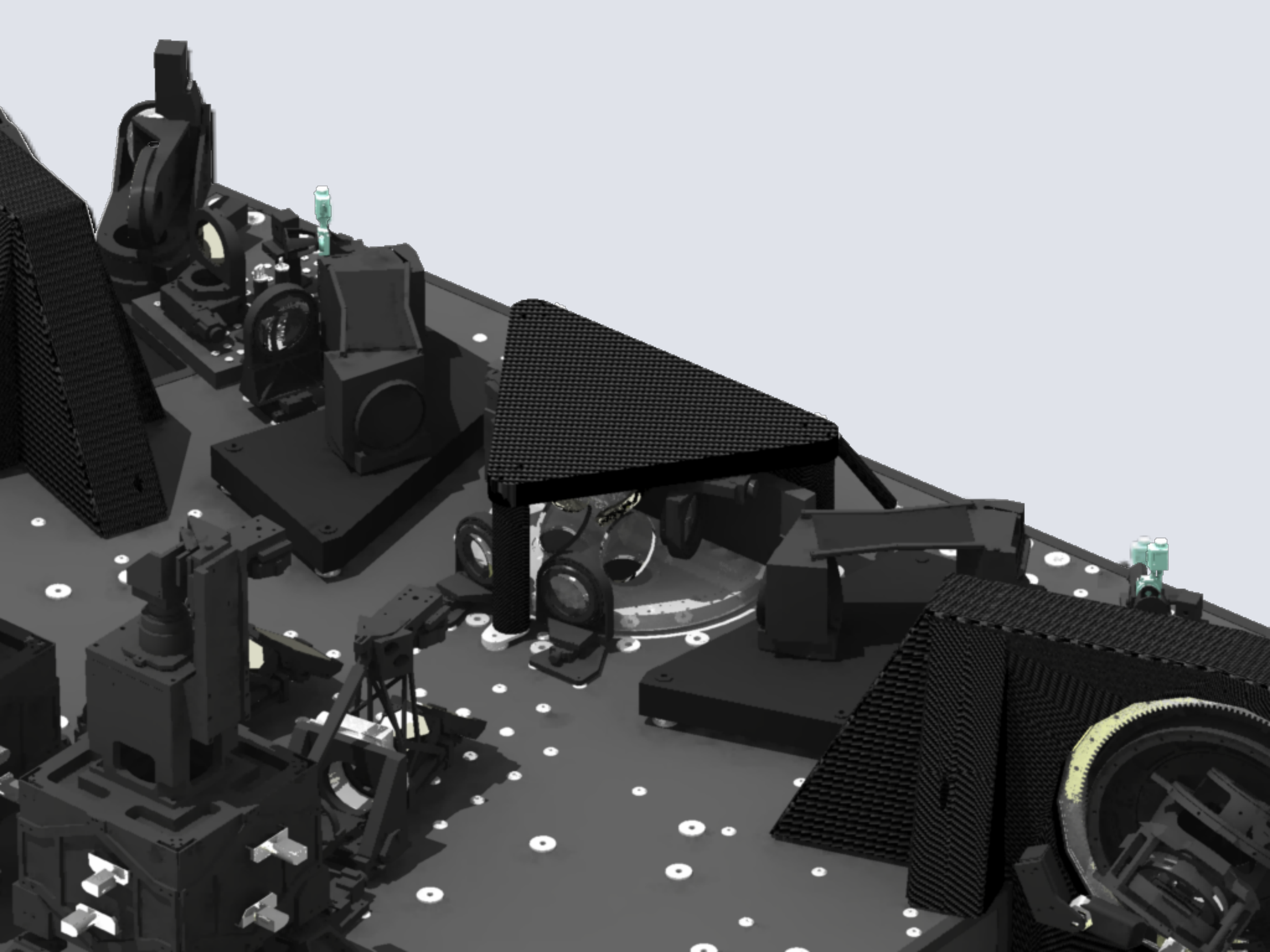


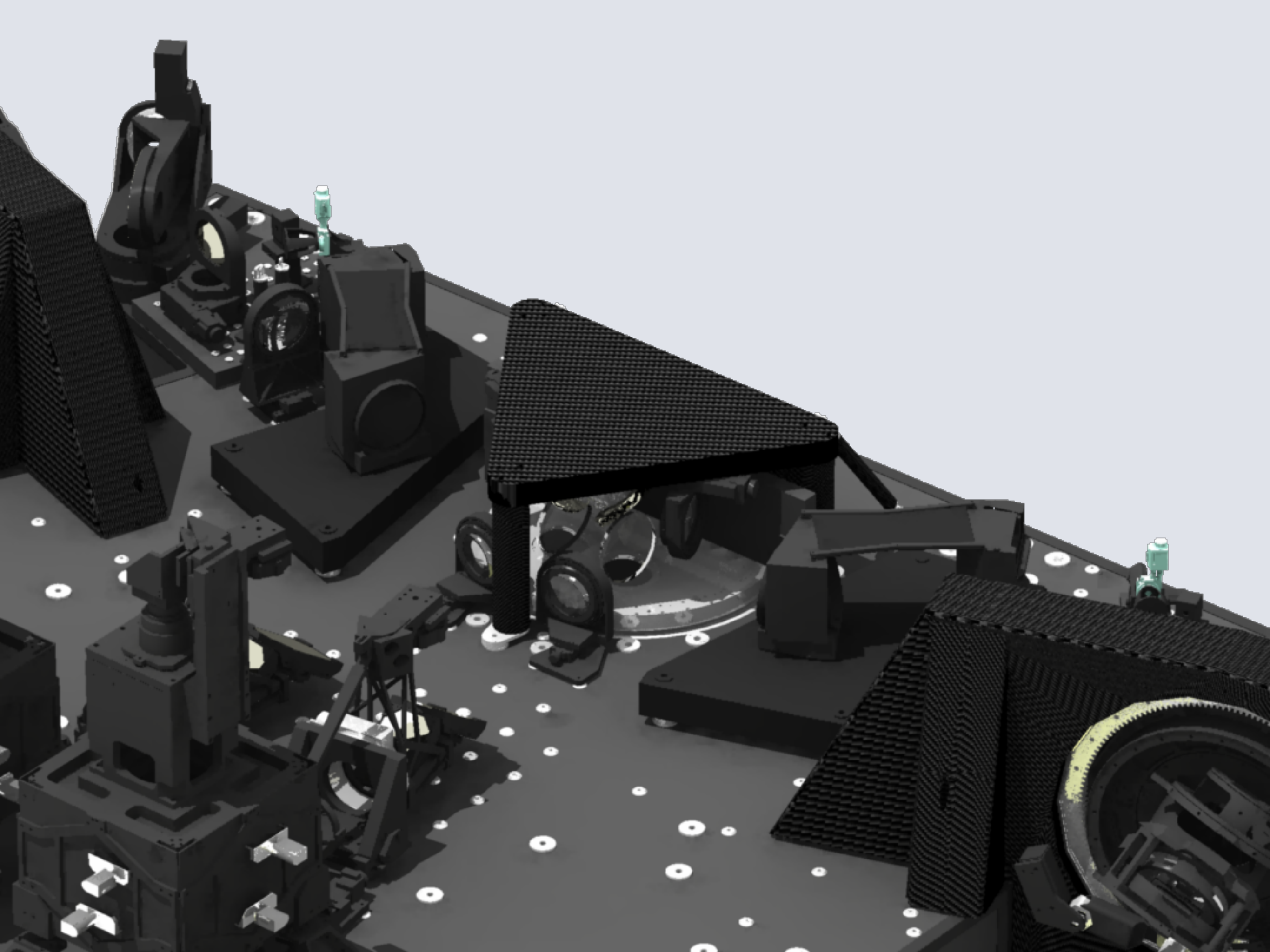
How it Works

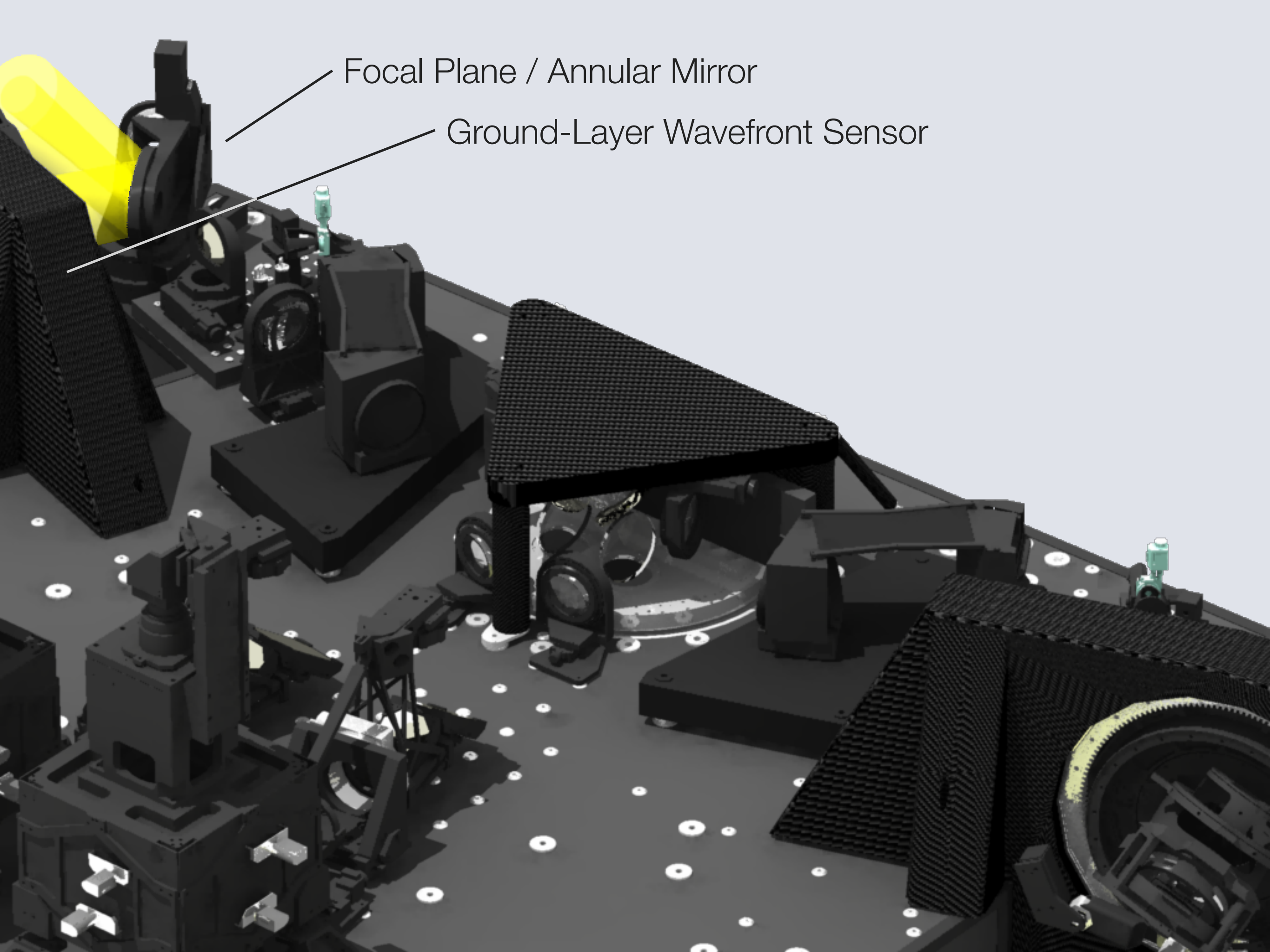






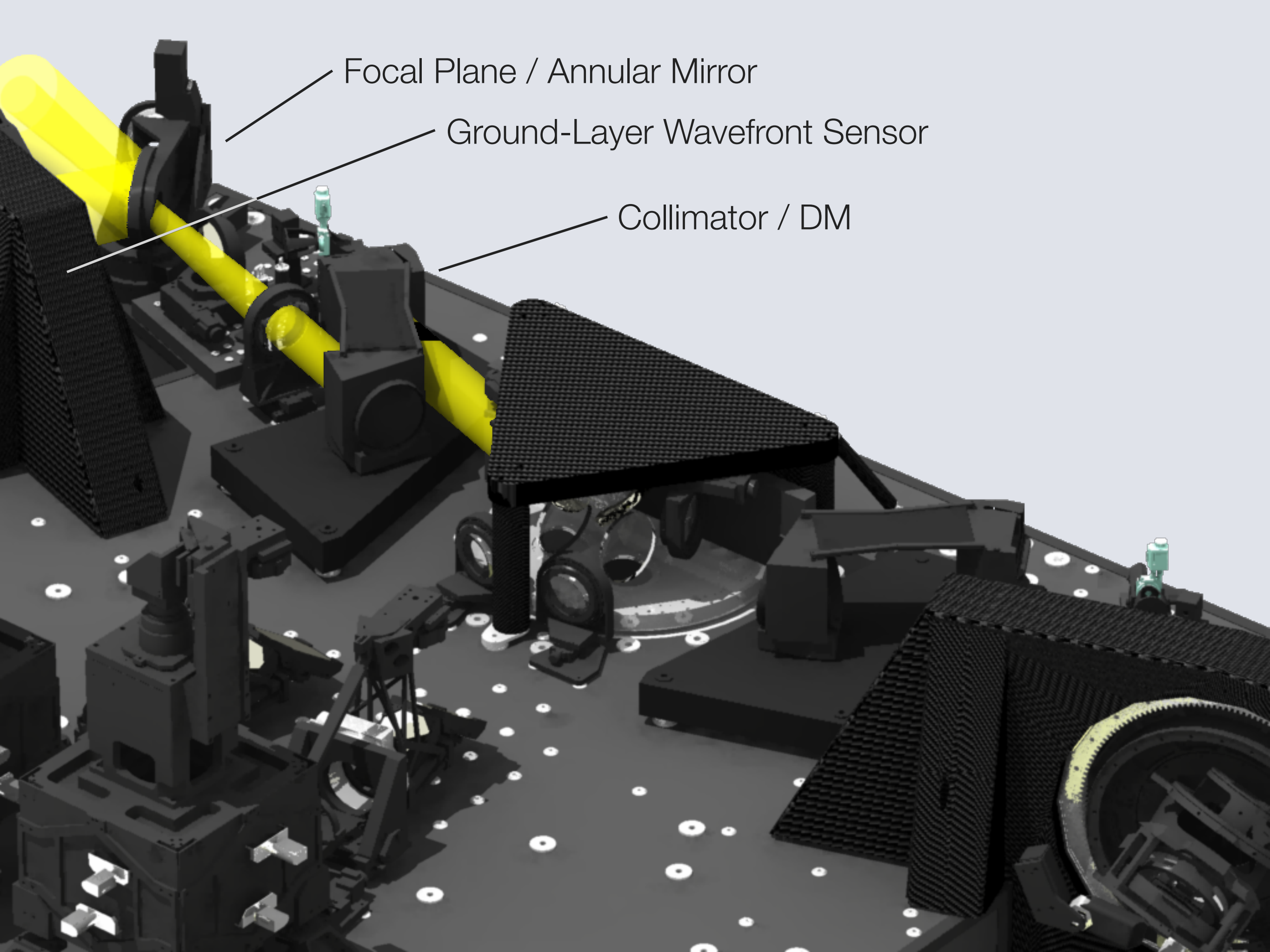






Focal Plane / Annular Mirror

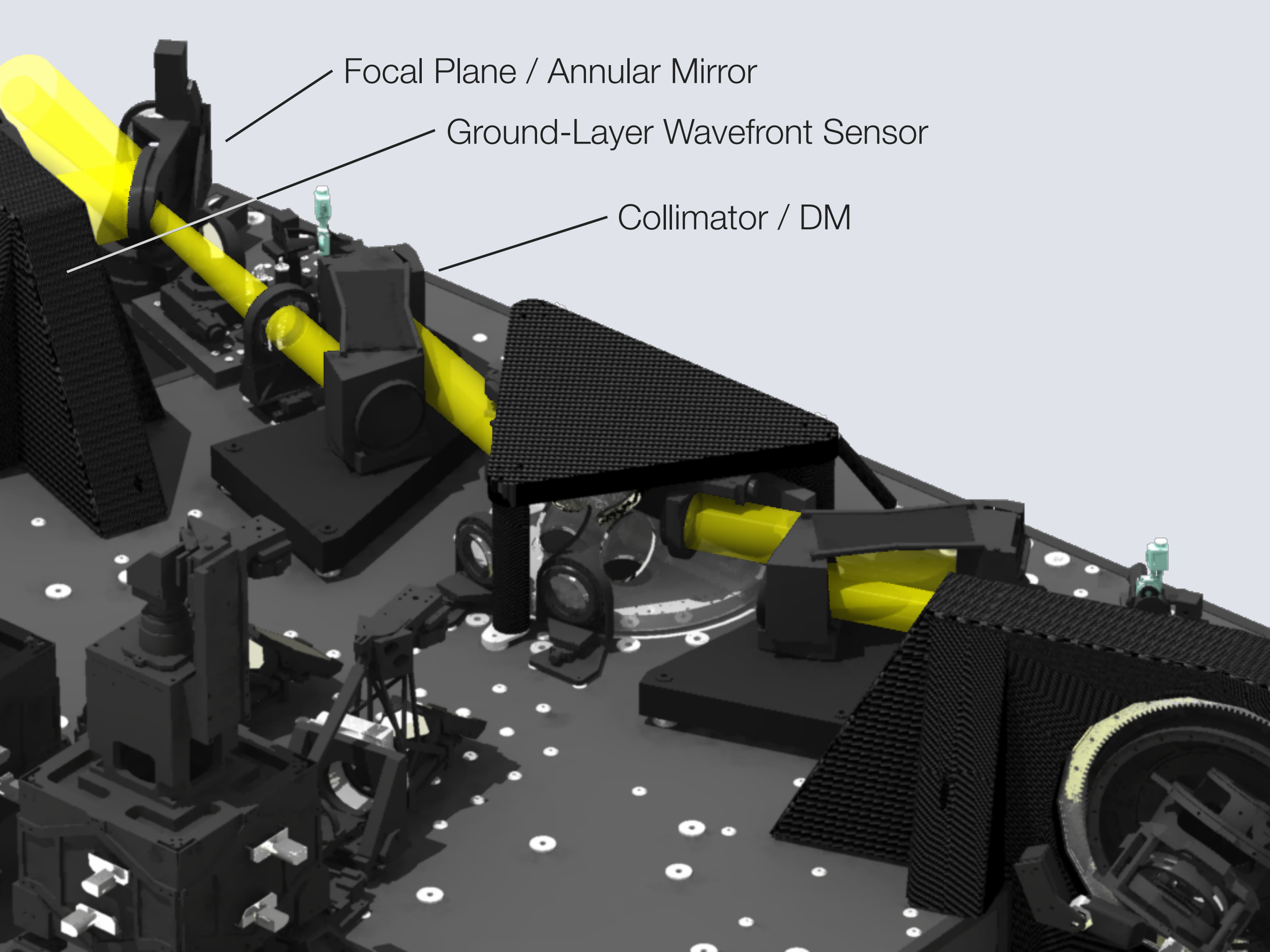
Ground-Layer Wavefront Sensor



Focal Plane / Annular Mirror

Ground-Layer Wavefront Sensor

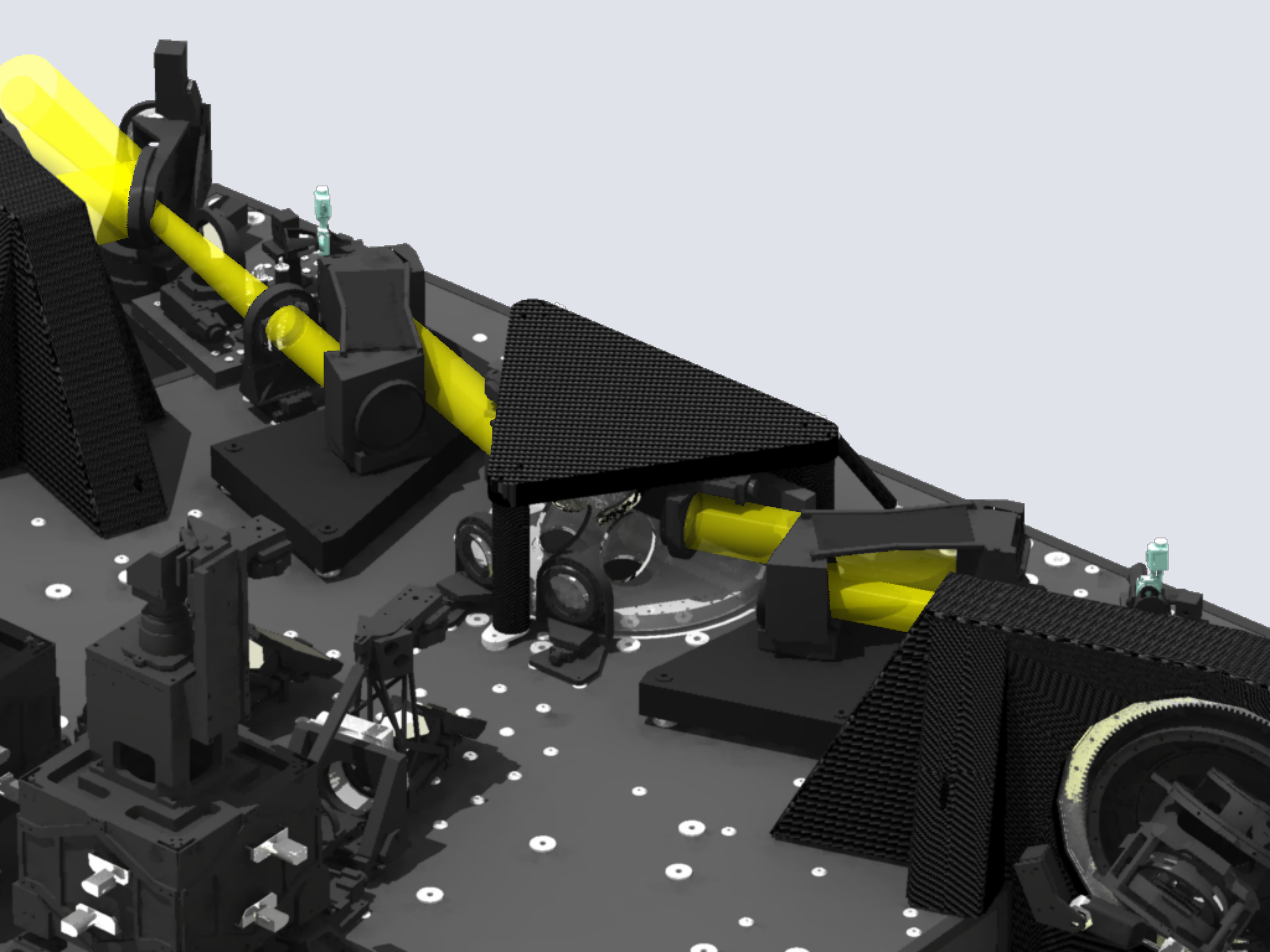
Collimator / DM

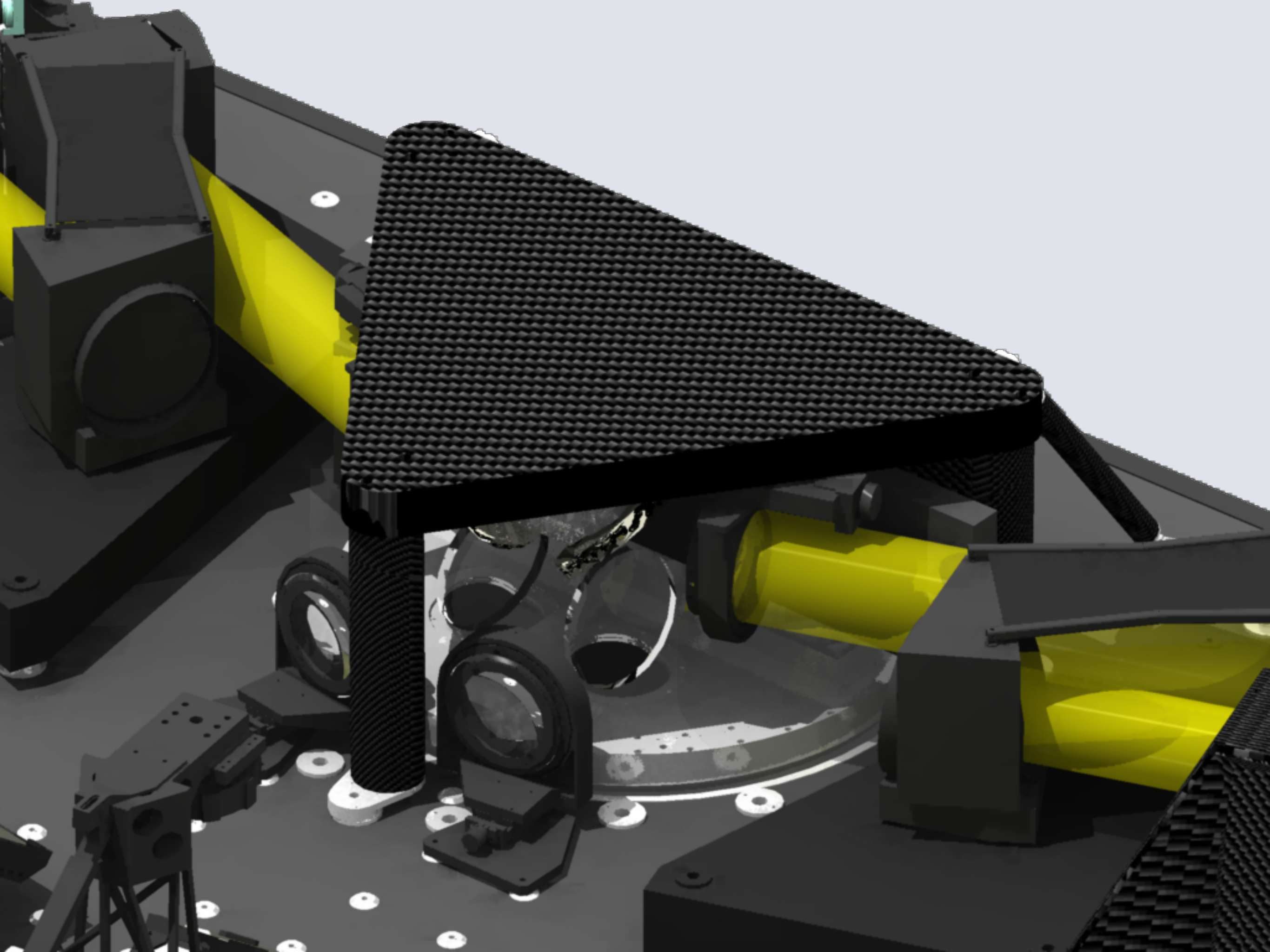


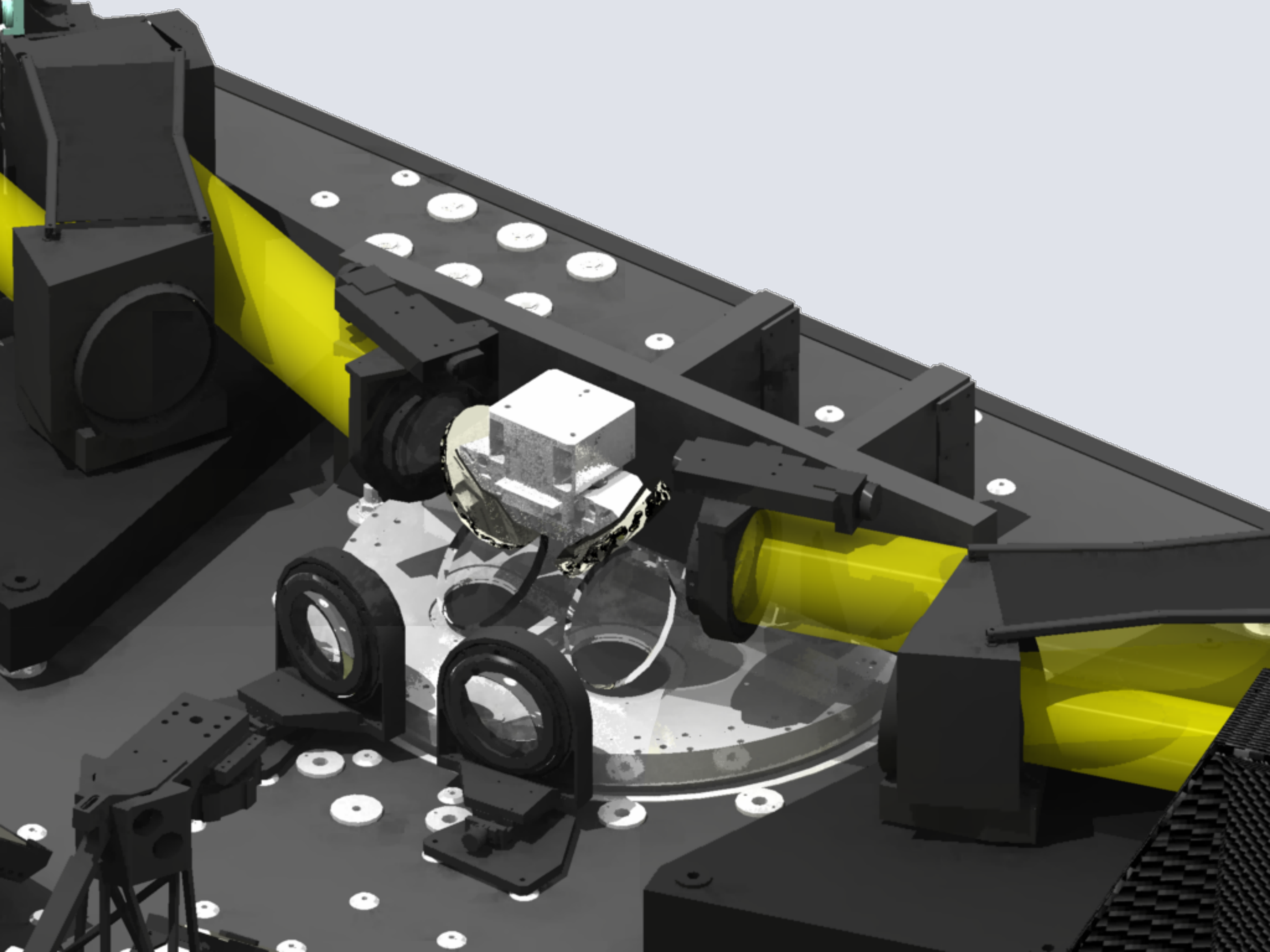
Focal Plane / Annular Mirror

Ground-Layer Wavefront Sensor

Collimator / DM



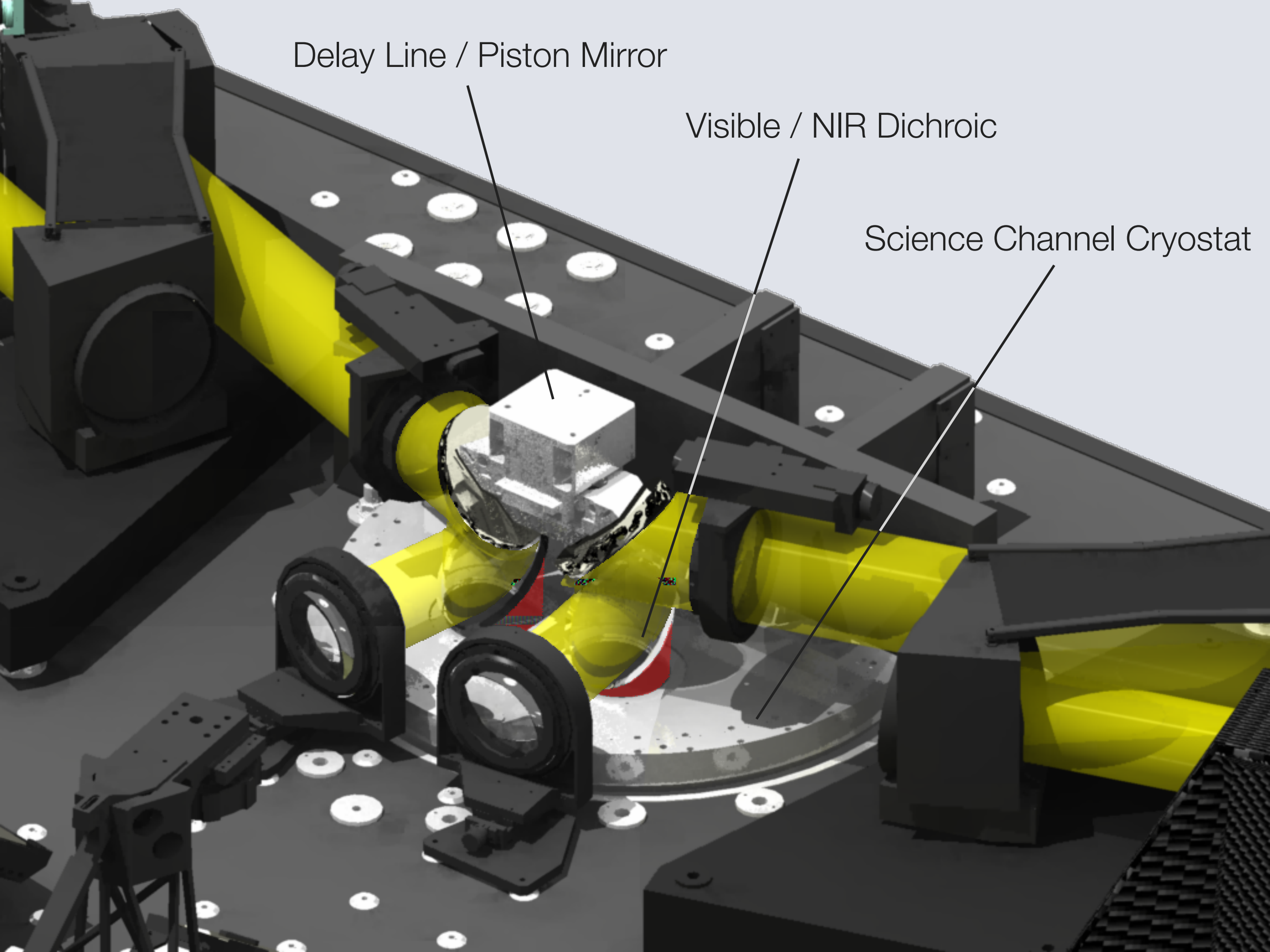


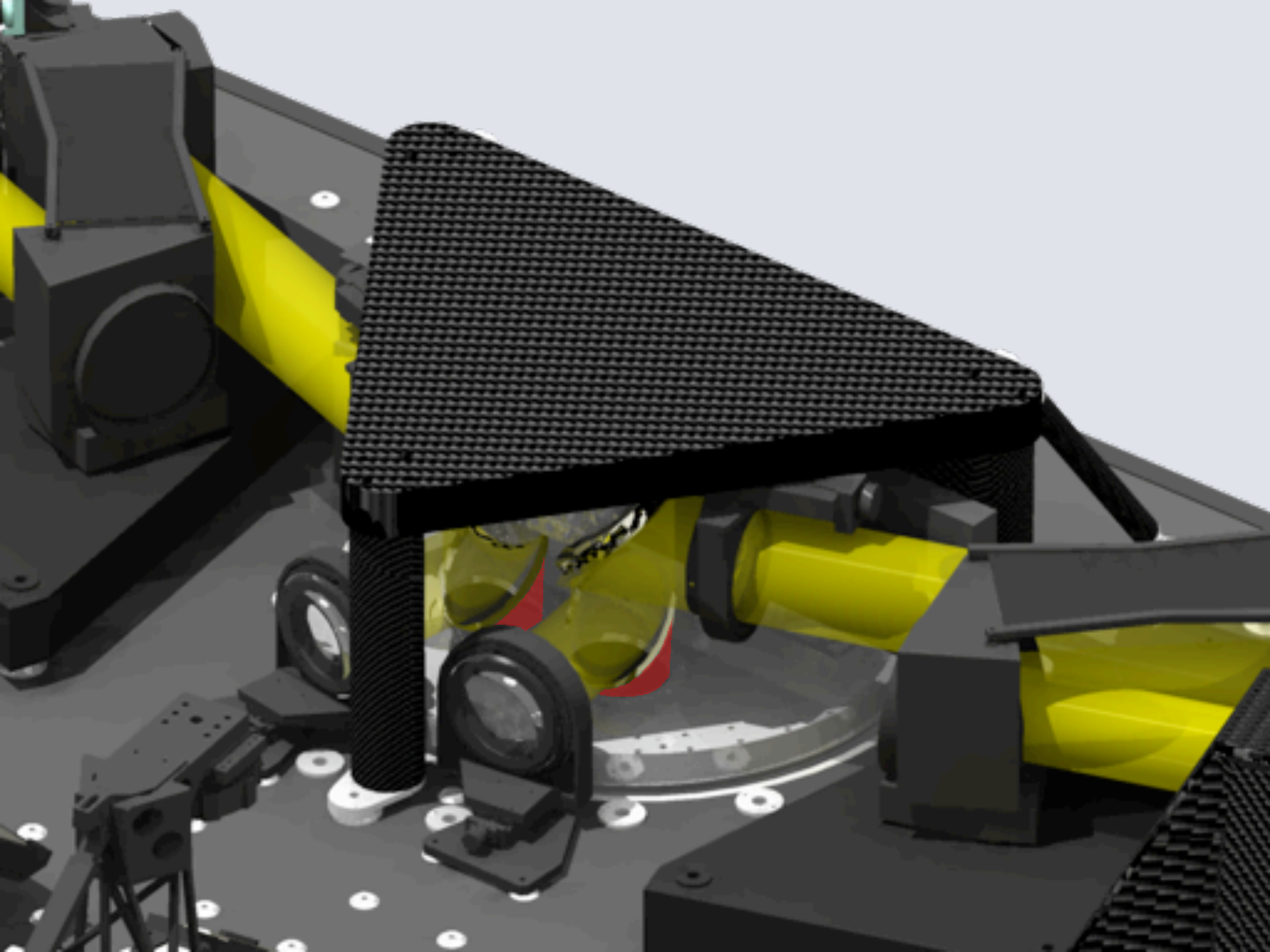


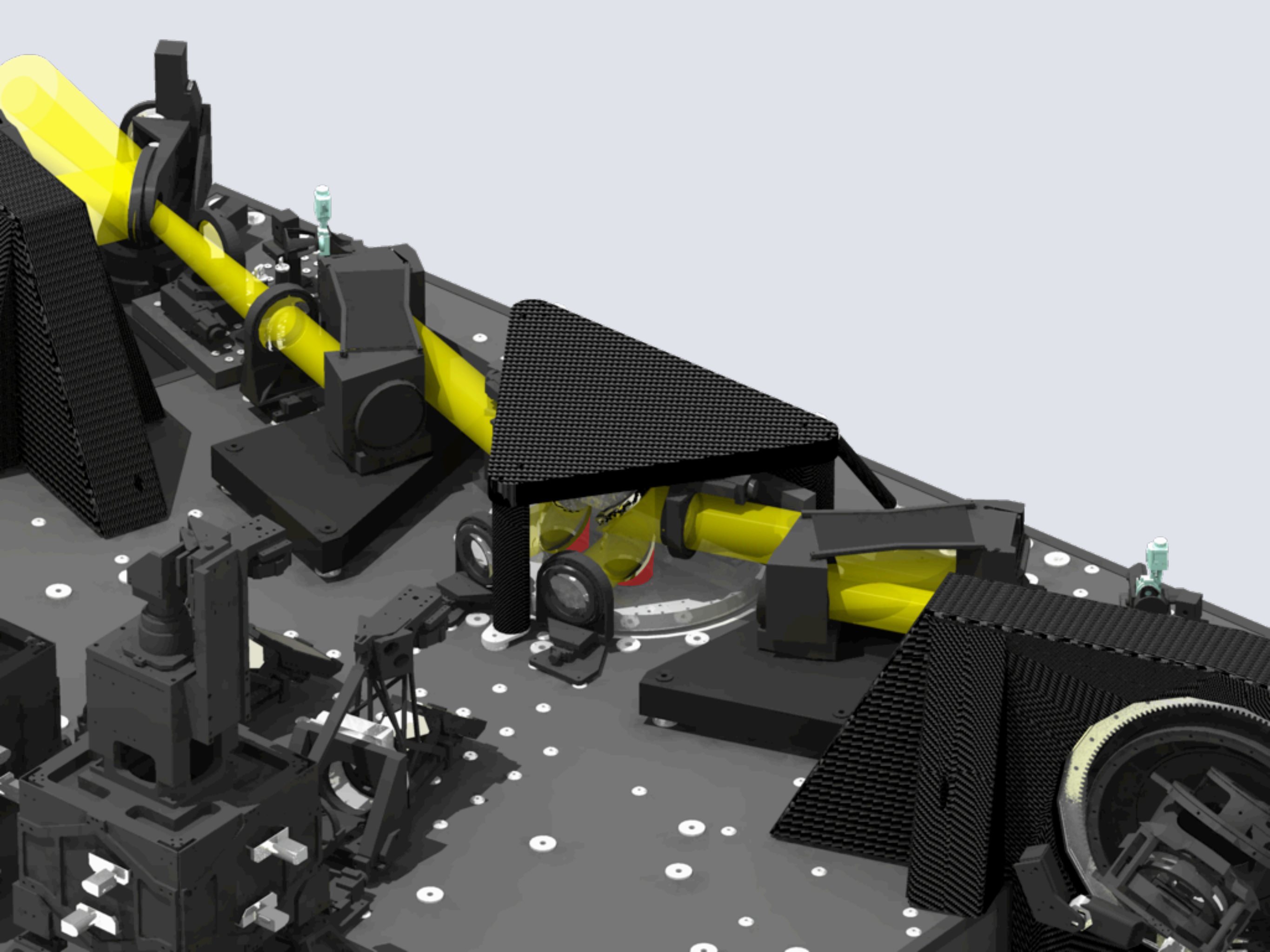
Delay Line / Piston Mirror

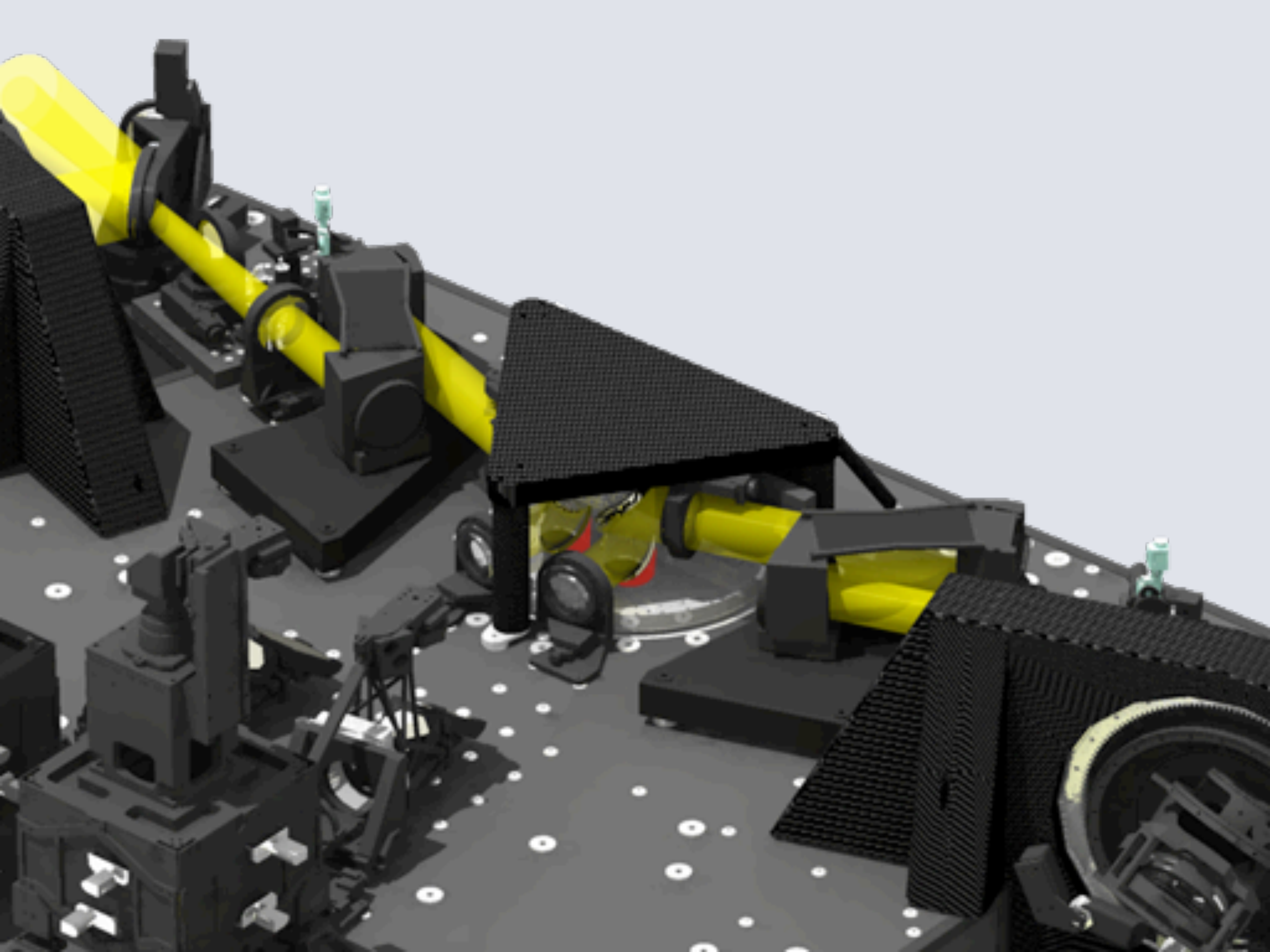
Visible / NIR Dichroic

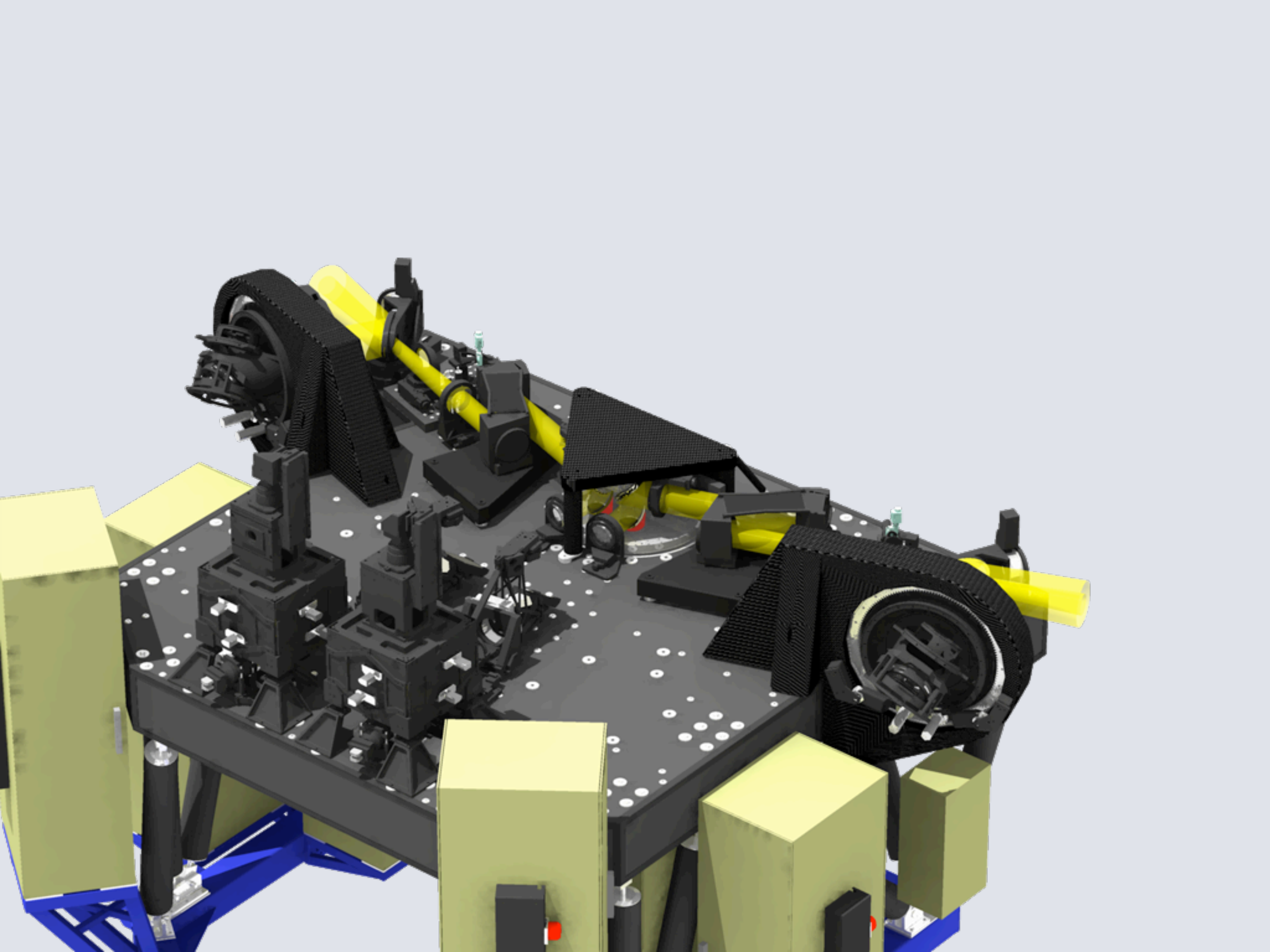
Science Channel Cryostat

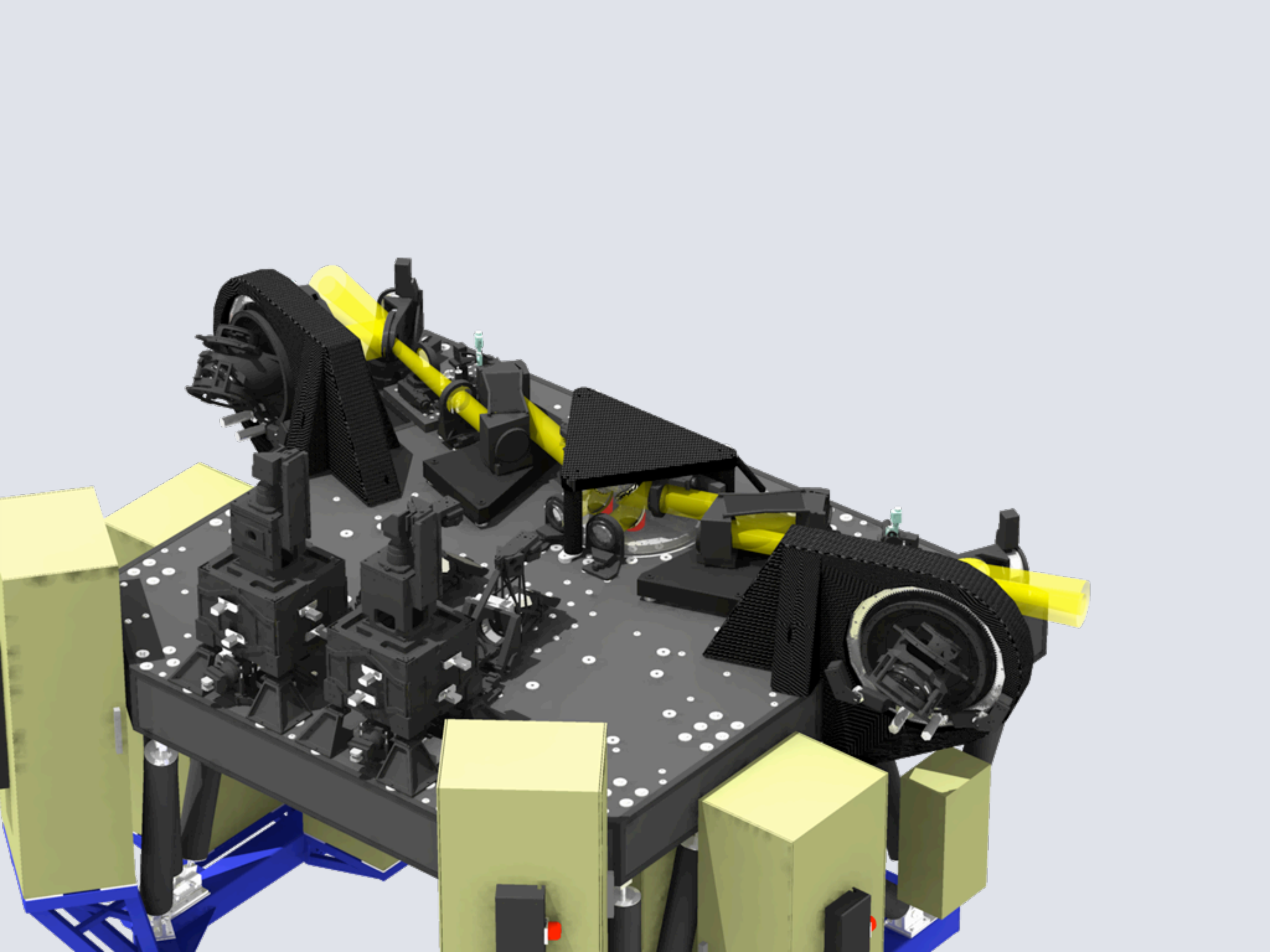




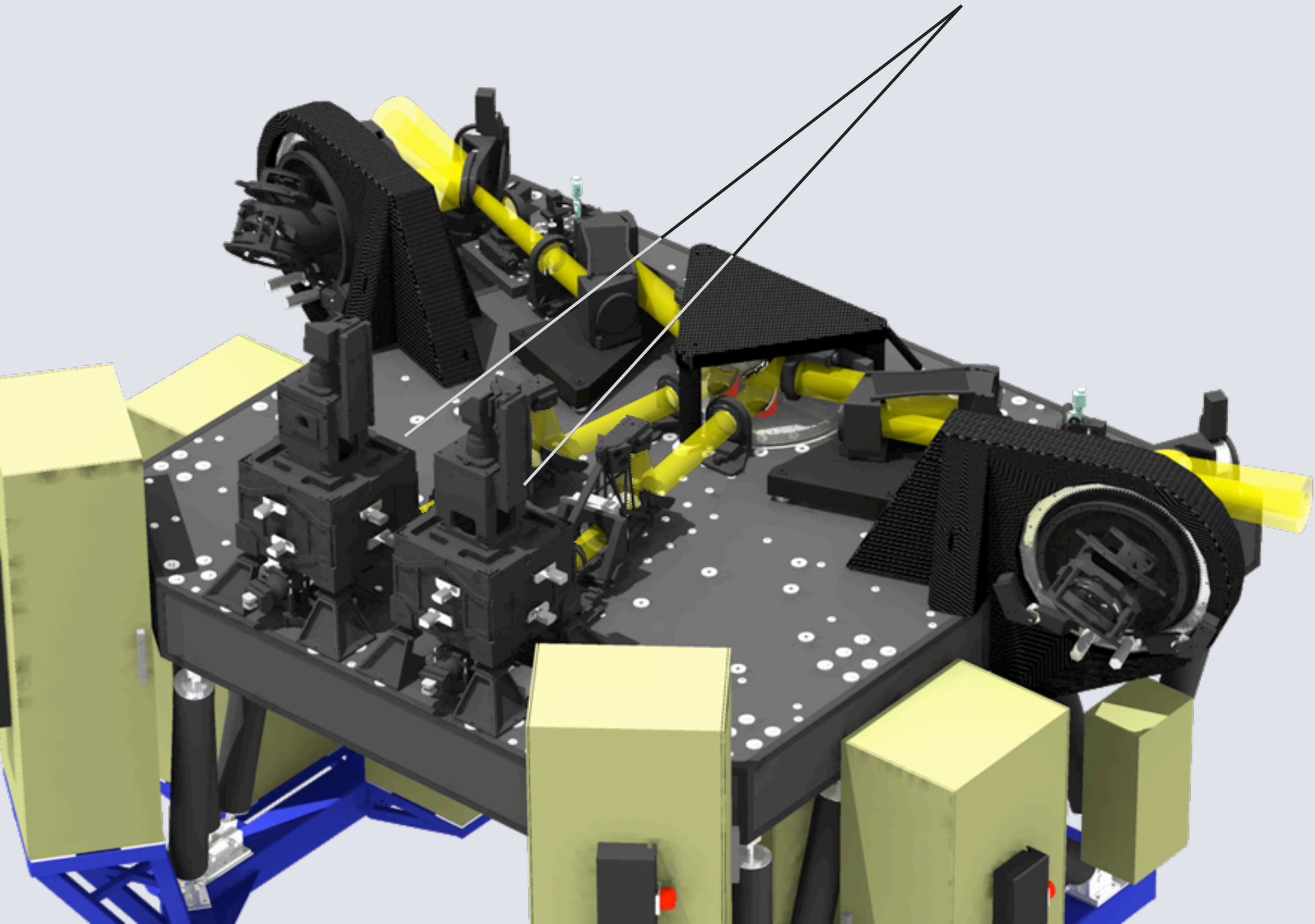


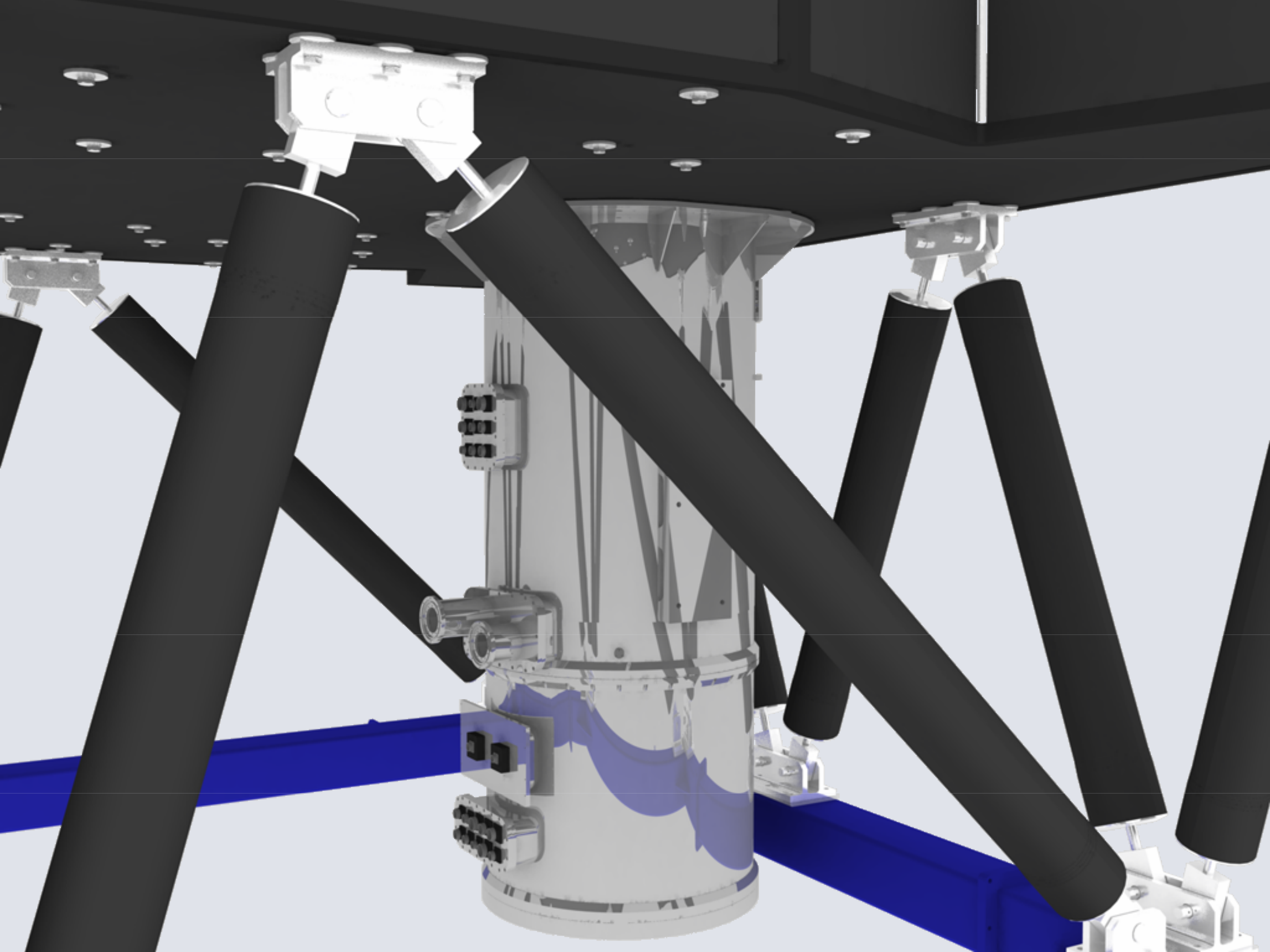






High-Layer Wavefront Sensors







LINC-NIRVANA Cryostat



LINC-NIRVANA Cryostat

Secondary

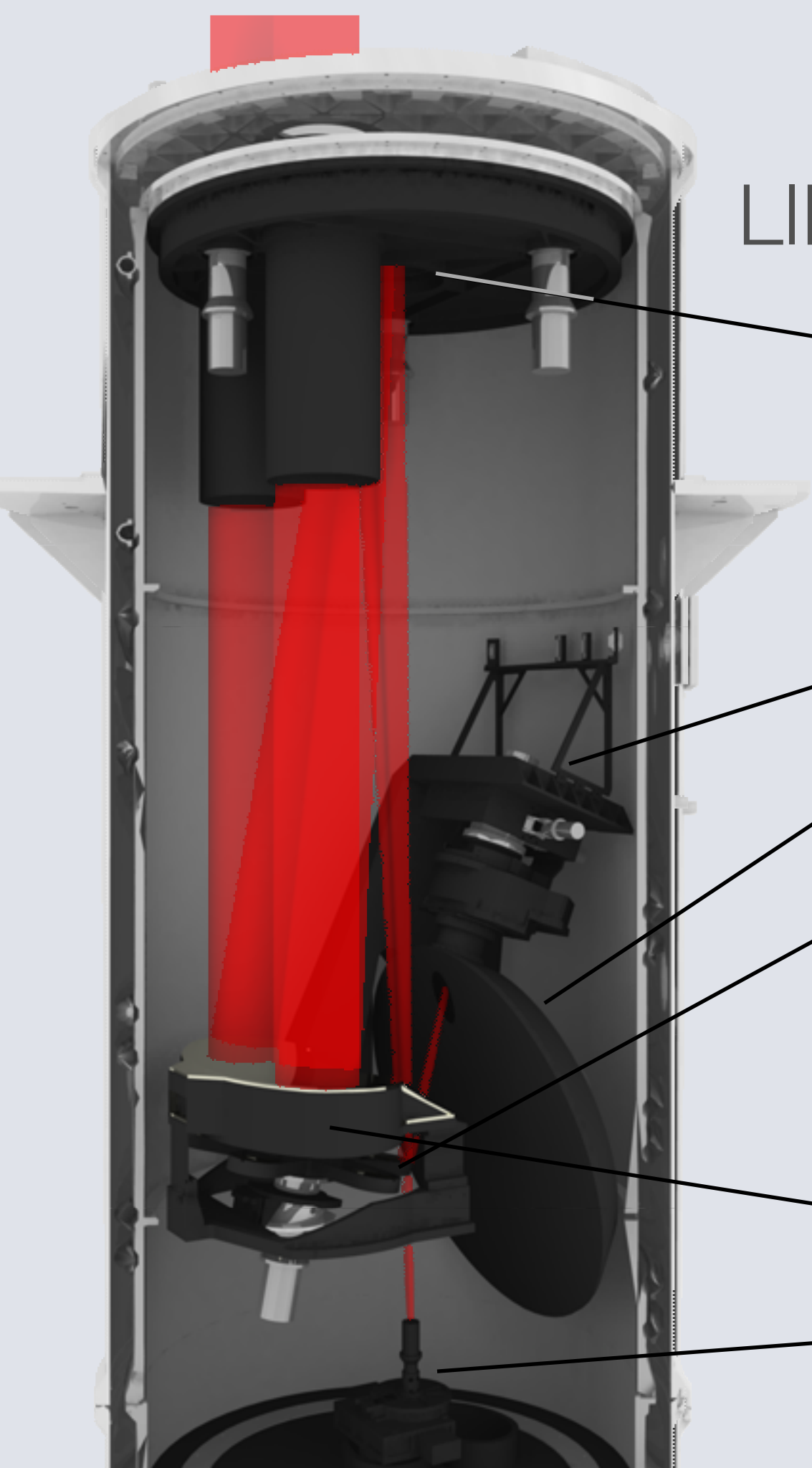
Science Detector / Derotator

Filter Wheel

IR-IR Dichroic

Primary

Fringe and Flexure Tracker



● About LINC-NIRVANA...

- What it is
- How it works

● Project Status...

- LINC-NIRVANA AIV (HD)
- First Light with Pathfinder (LBT)

● Implementation Plan...

- MCAO & Interferometry
- What's next...

LINC-NIRVANA AIV



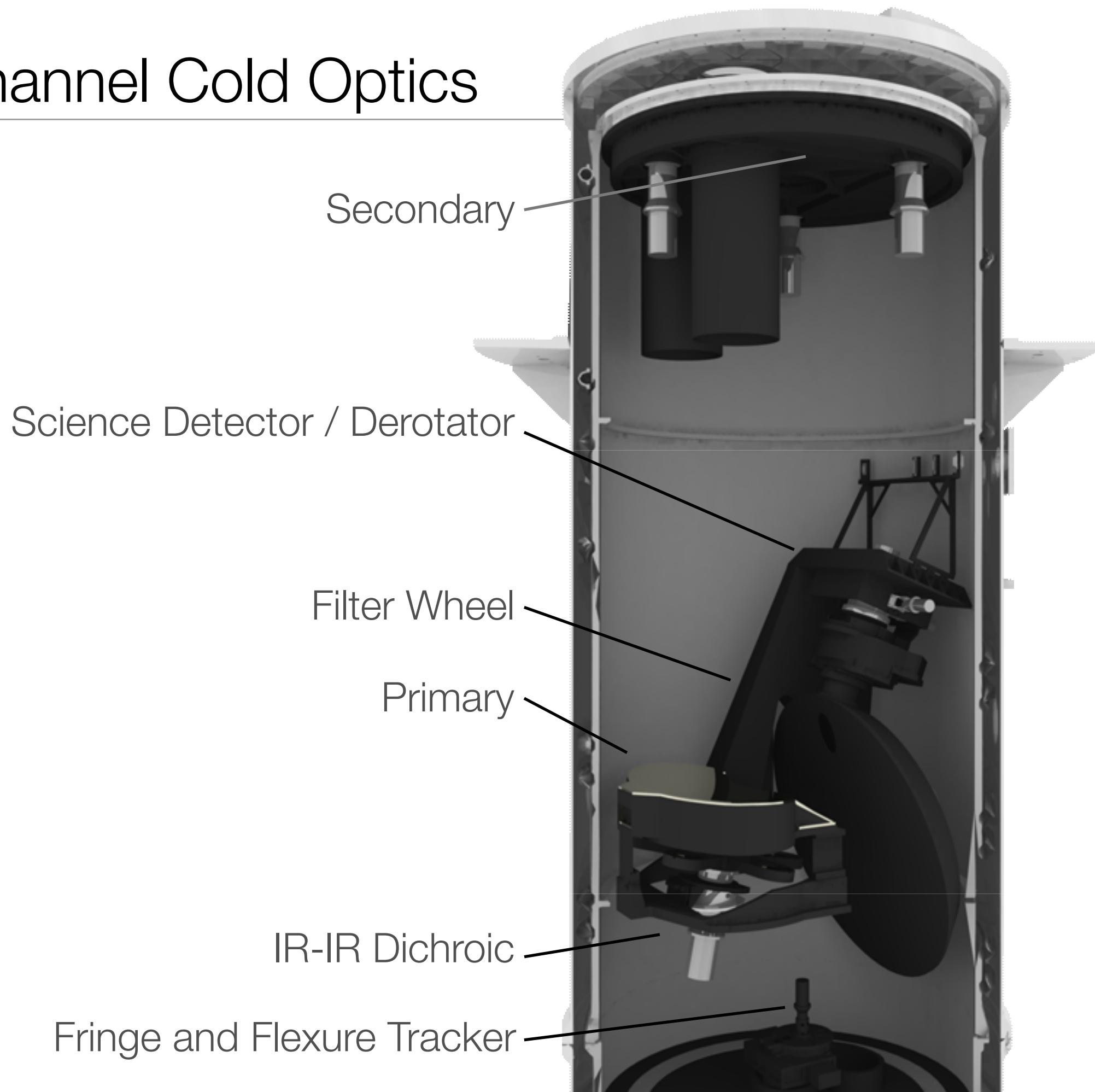
AstroTechTalk
26 September 2014
MPIA Heidelberg

- Ground-Layer WFS testing complete
- Science channel cold optics delivered, tested, accepted
- Fringe and flexure tracker progressing
- Piston mirror control optimization complete
- LBT rail flexure problem solved
- Ground, High-Layer wavefront sensors accepted
- OVMS complete and handed over
- Cold testing of cryomechanics, FFTS, etc.
- Final bench integration 1st arm complete...

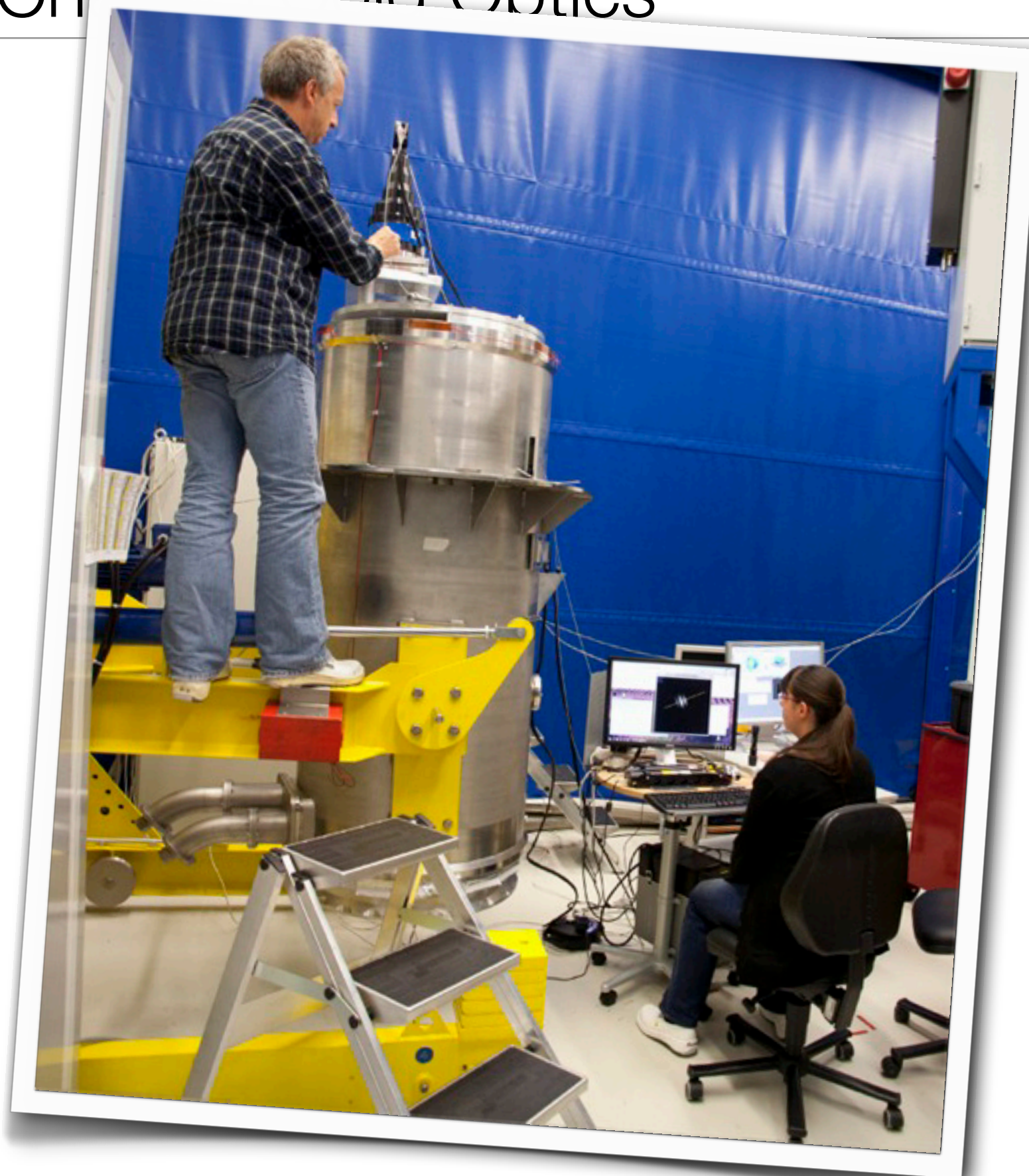
- Ground-Layer WFS testing complete
- Science channel cold optics delivered, tested, accepted
- Fringe and flexure tracker progressing
- Piston mirror control optimization complete
- LBT rail flexure problem solved
- Ground, High-Layer wavefront sensors accepted
- OVMS complete and handed over
- Cold testing of cryomechanics, FFTS, etc.
- Final bench integration 1st arm complete...

Science Channel Cold Optics

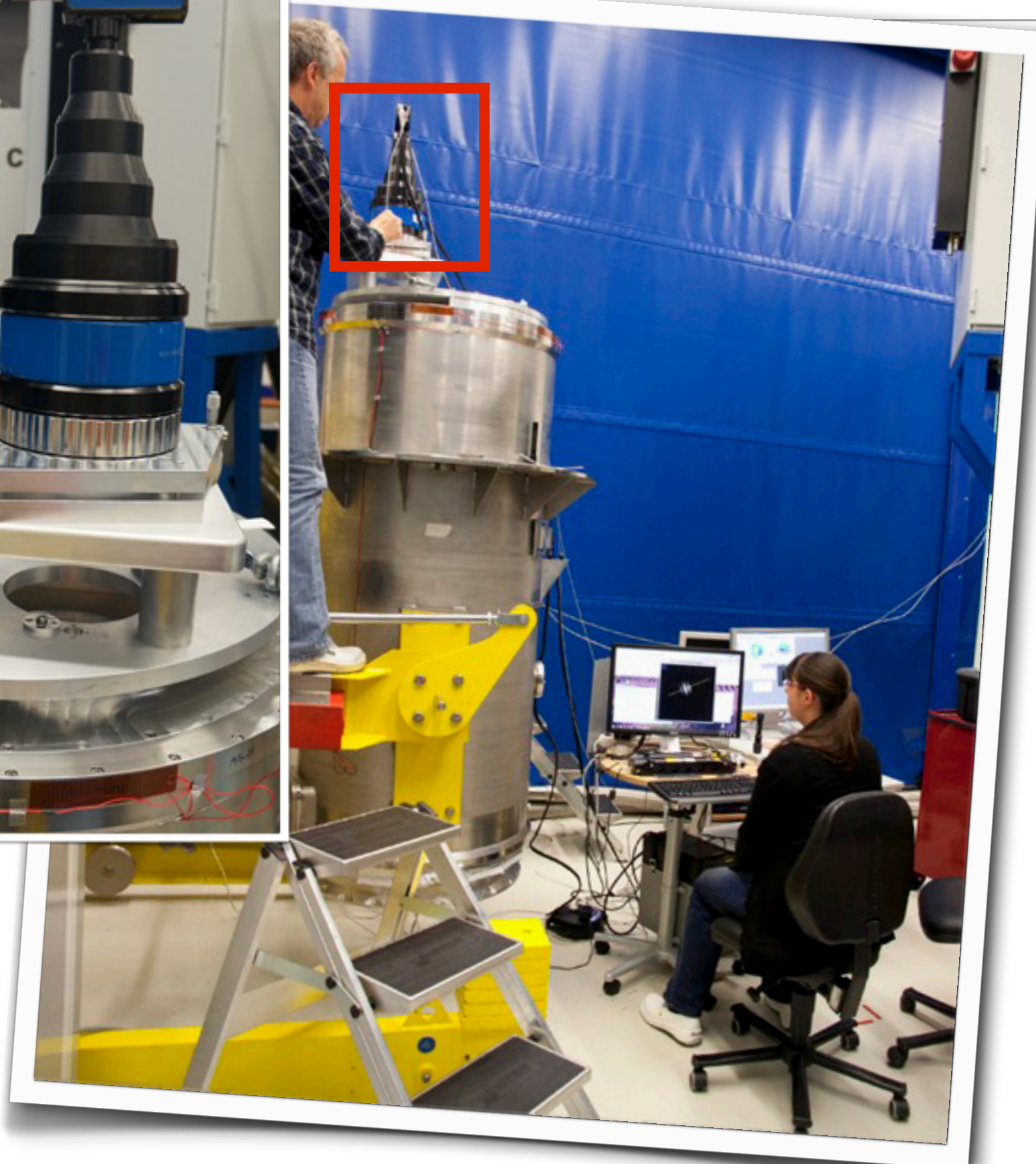
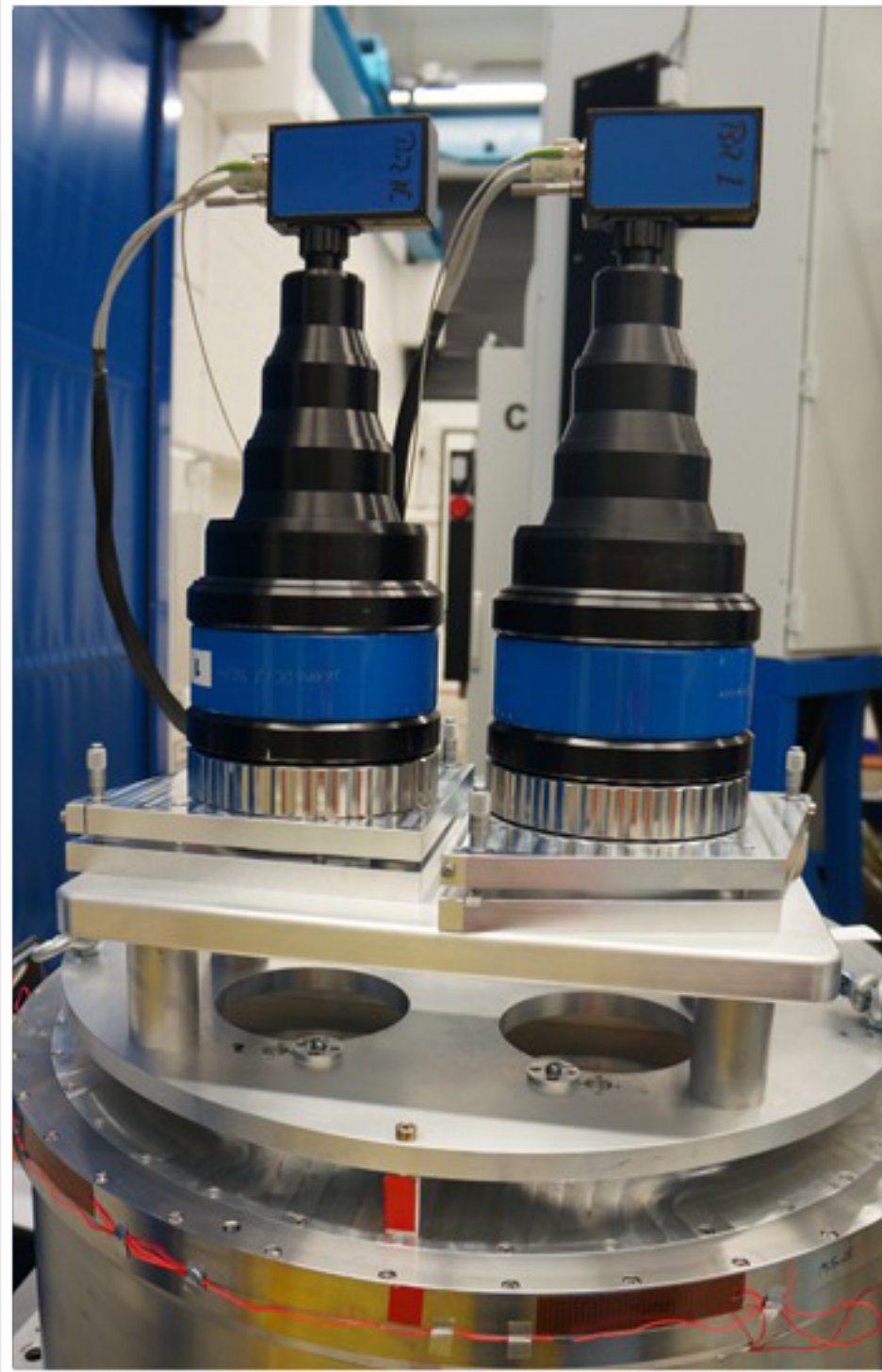
Science Channel Cold Optics



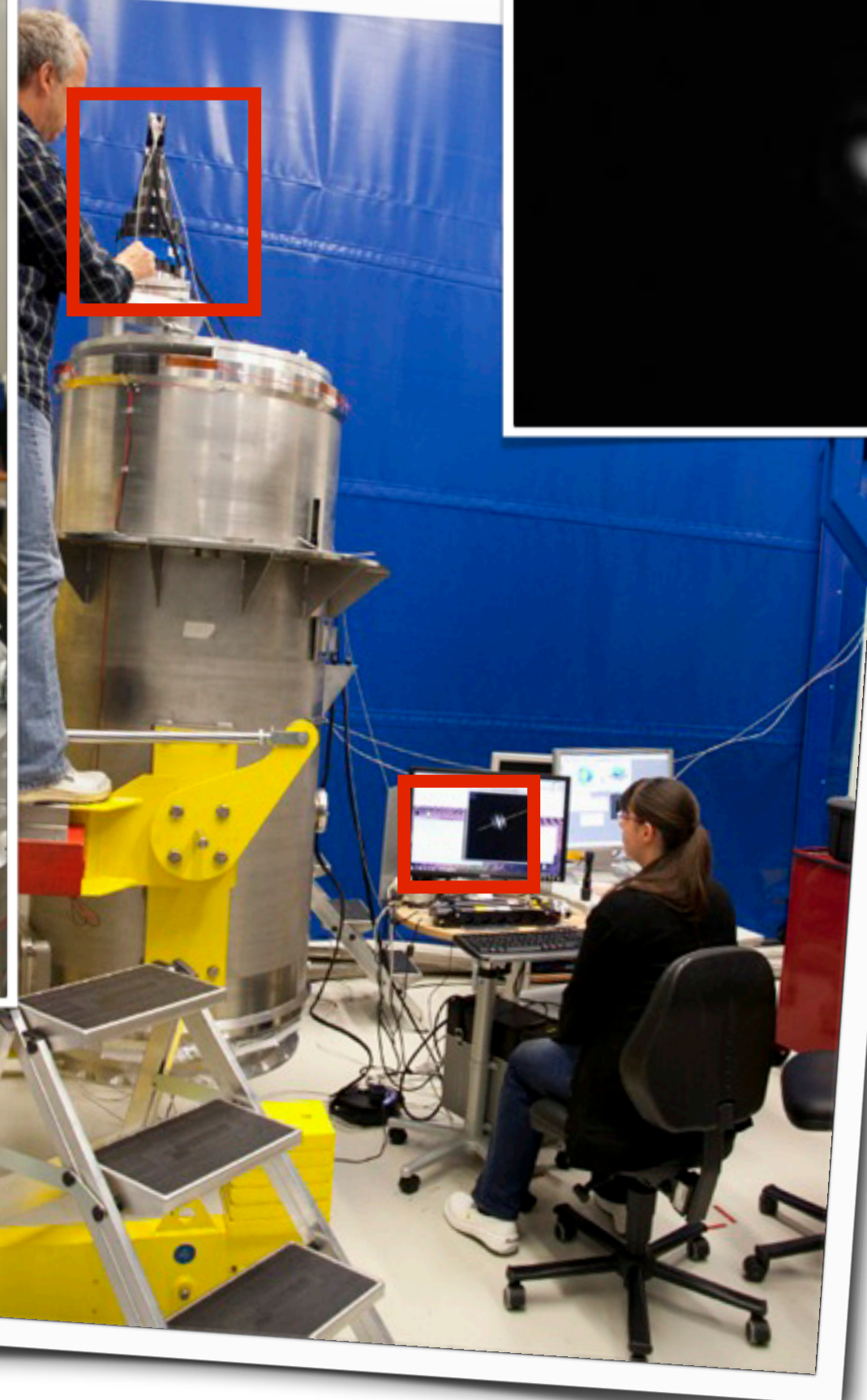
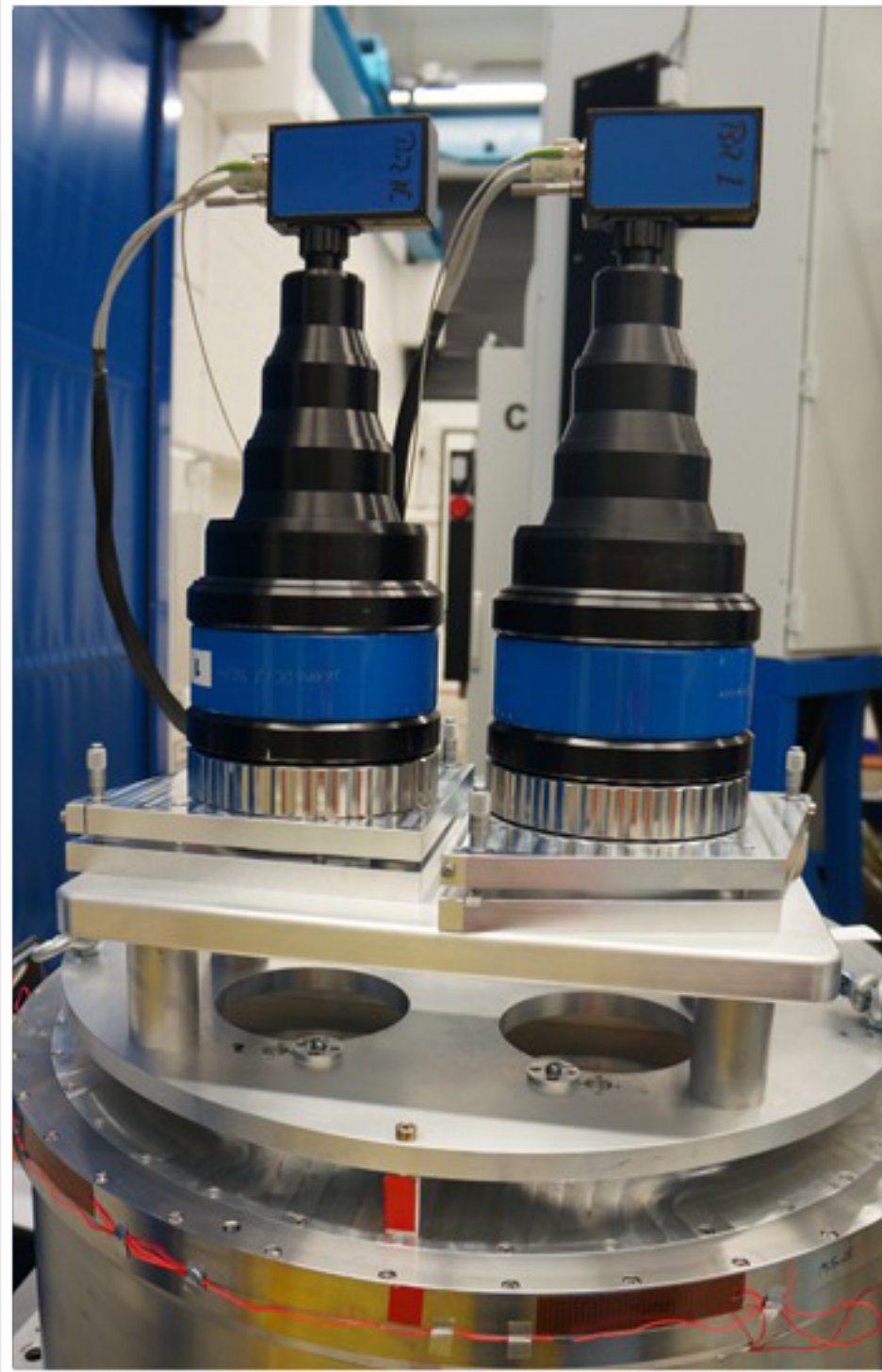
Science Channel Cold Optics



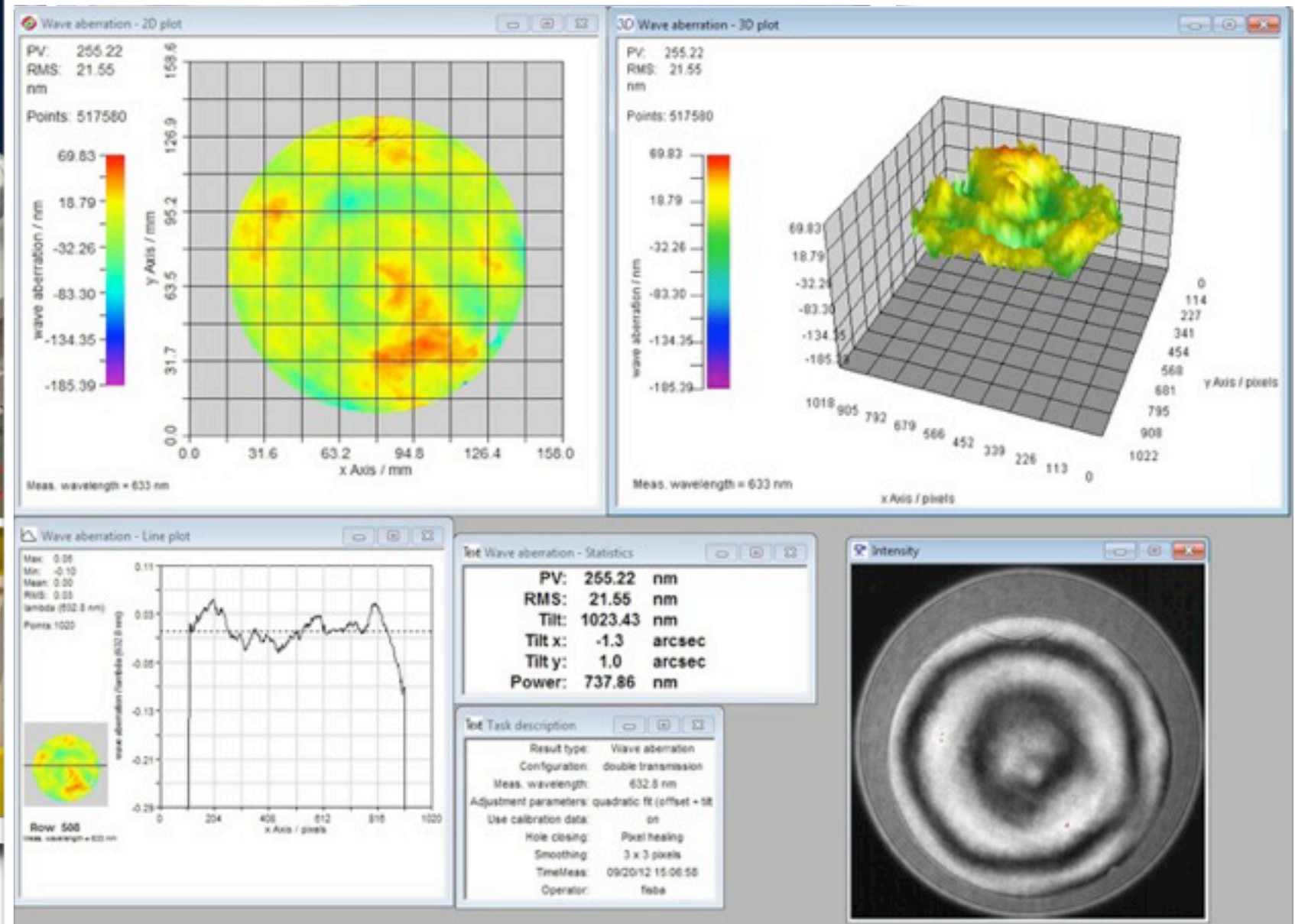
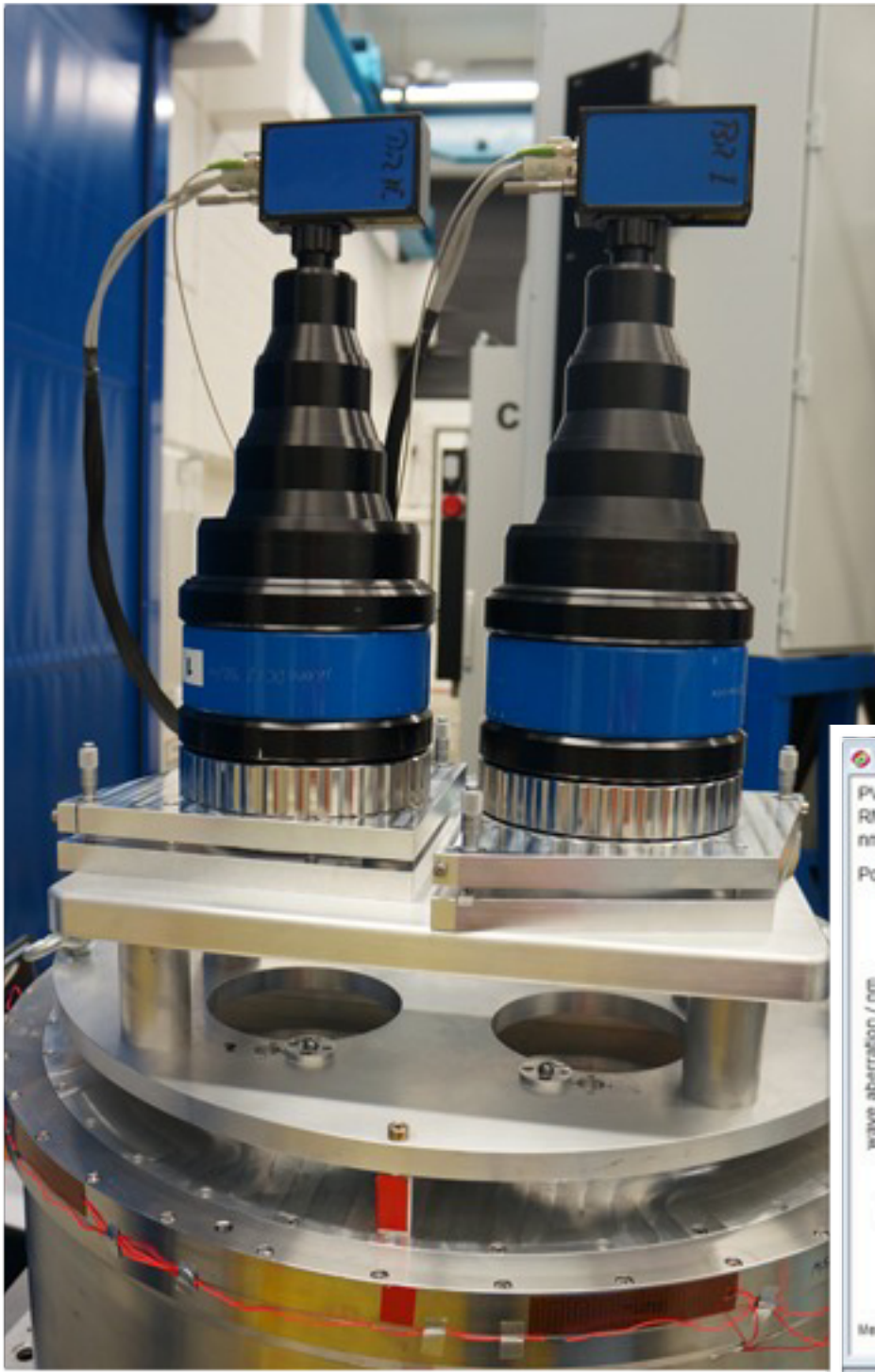
Cold Optics



Cold Optics



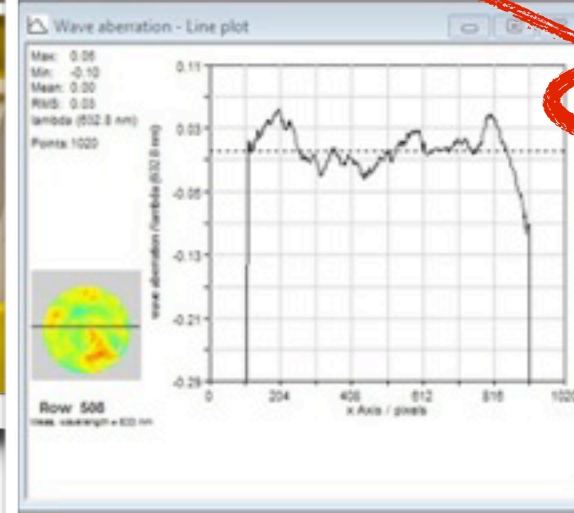
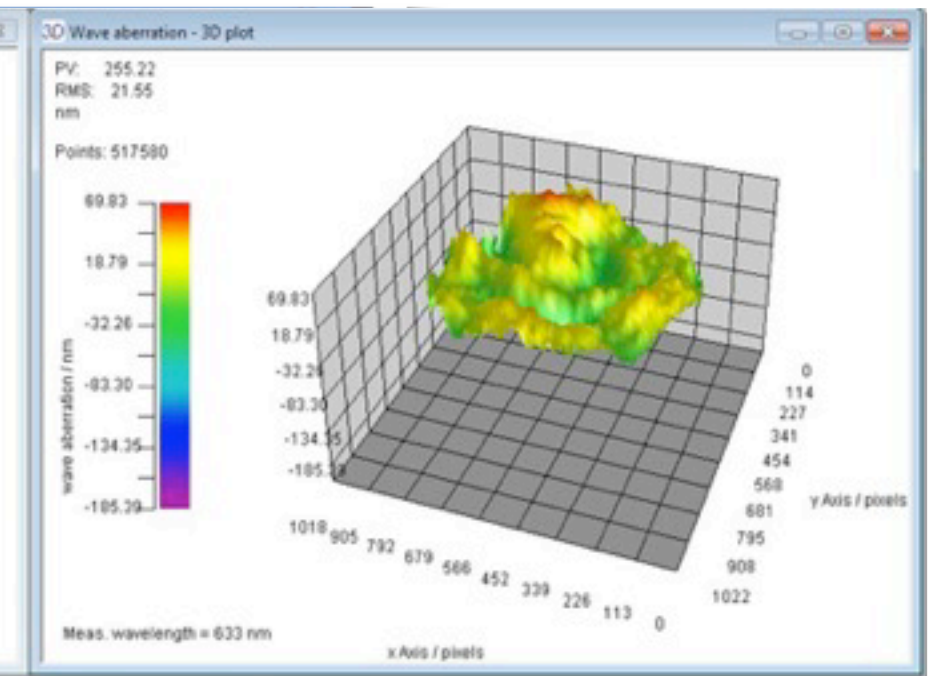
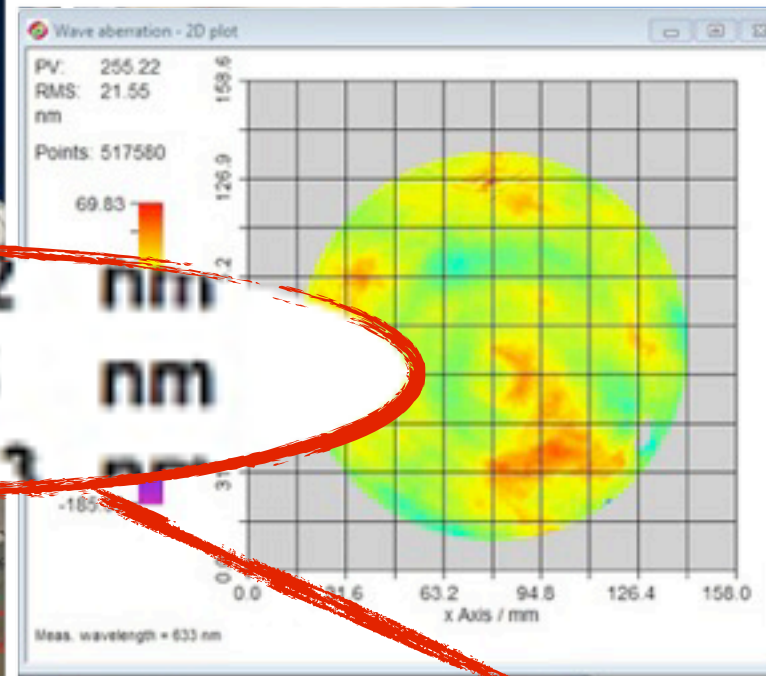
Cold Optics



Cold Optics

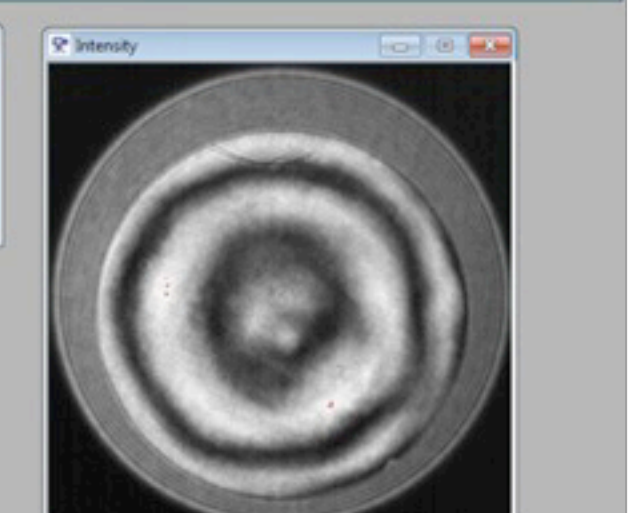


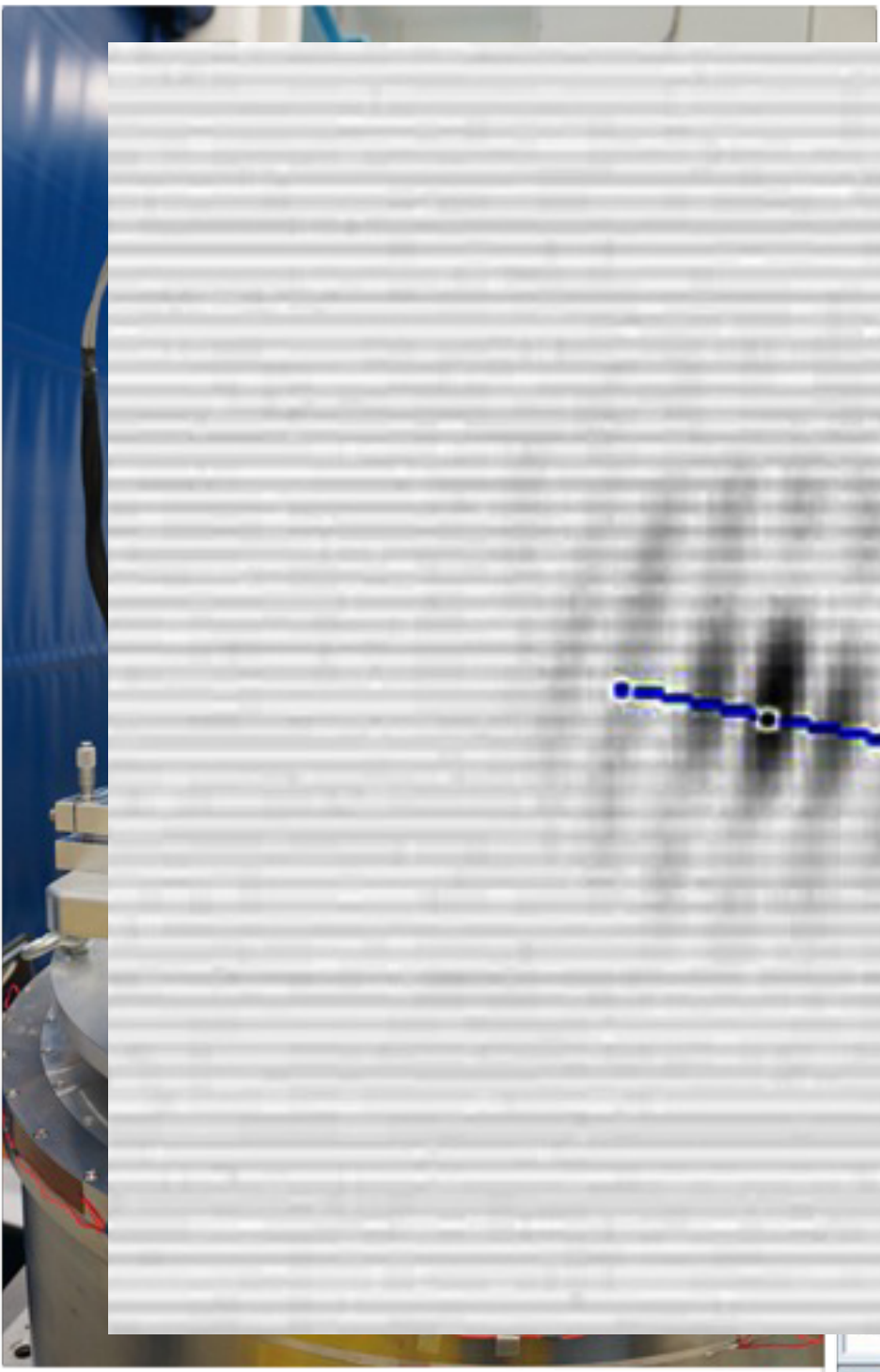
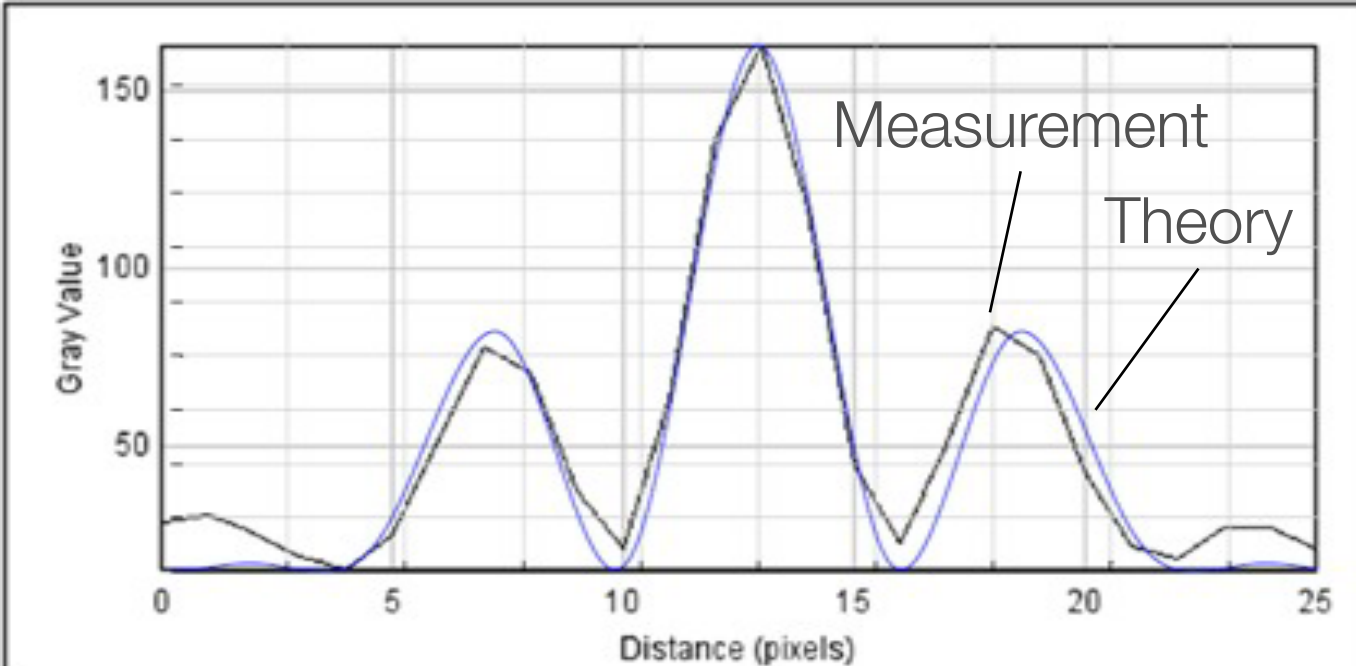
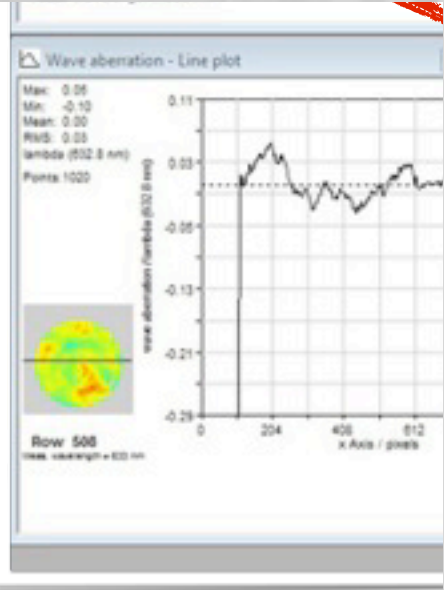
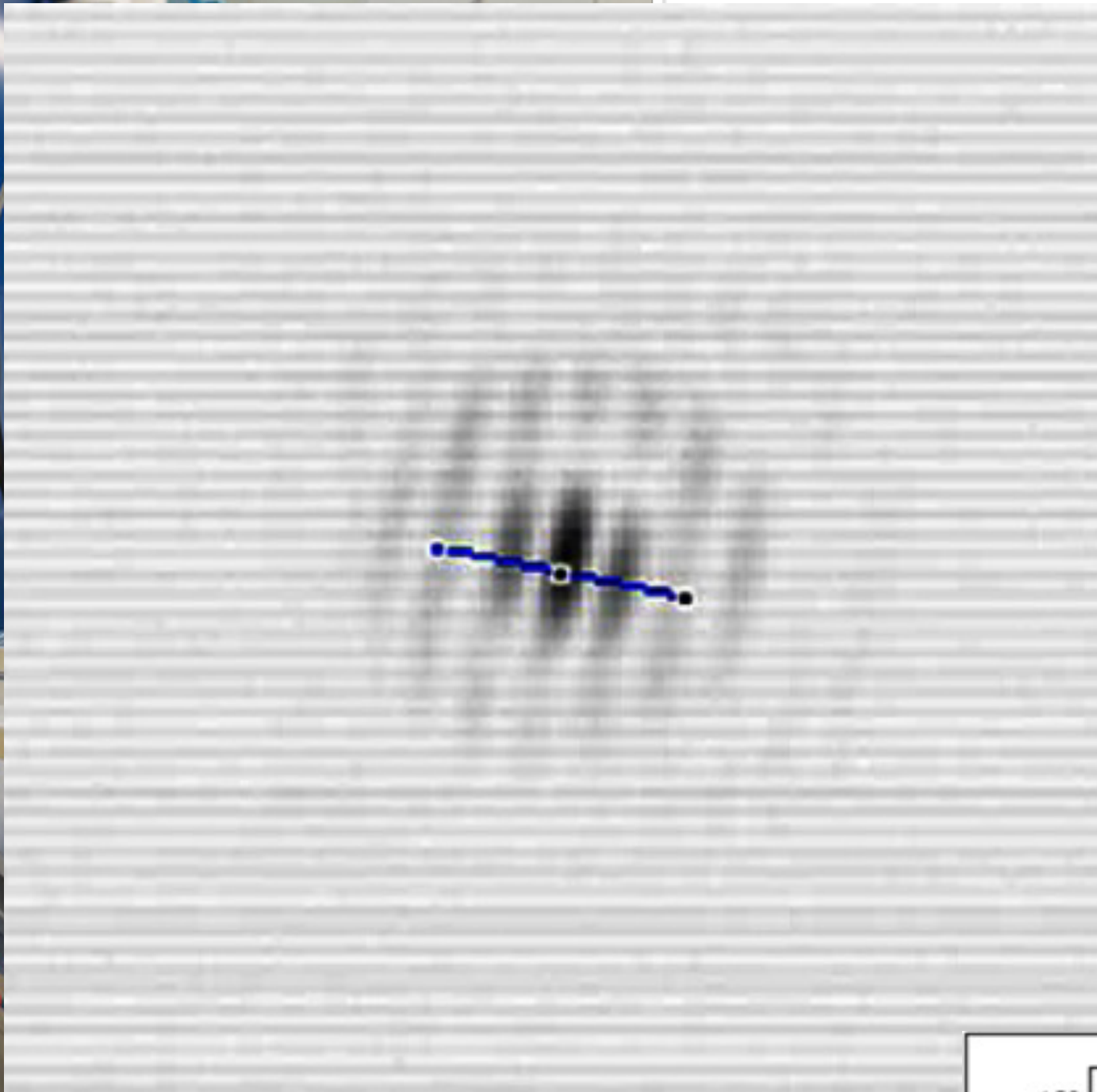
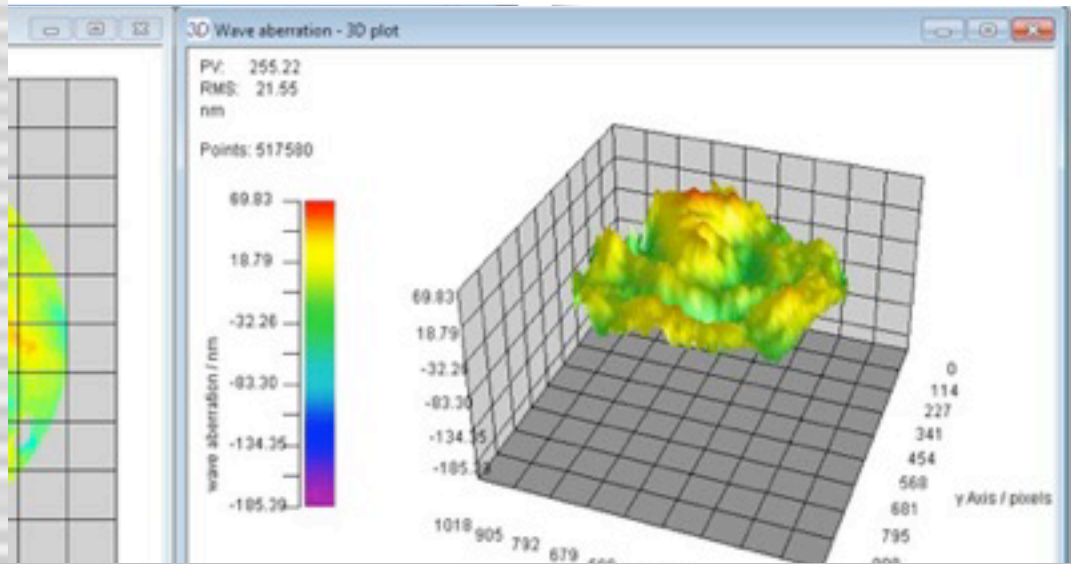
PV: 255.22 nm
RMS: 21.55 nm
Tilt: 1023.43 arcsec



PV: 255.22 nm
RMS: 21.55 nm
Tilt: 1023.43 arcsec
Tilt x: -1.5 arcsec
Tilt y: 1.0 arcsec
Power: 737.86 nm

Test Task description
Result type: Wave aberration
Configuration: double transmission
Meas. wavelength: 632.8 nm
Adjustment parameters: quadratic fit (offset = 18)
Use calibration data: on
Hole closing: Pixel healing
Smoothing: 3 x 3 pixels
Time/Meas: 09/20/12 15:06:58
Operator: feba





Wavefront Sensors 1

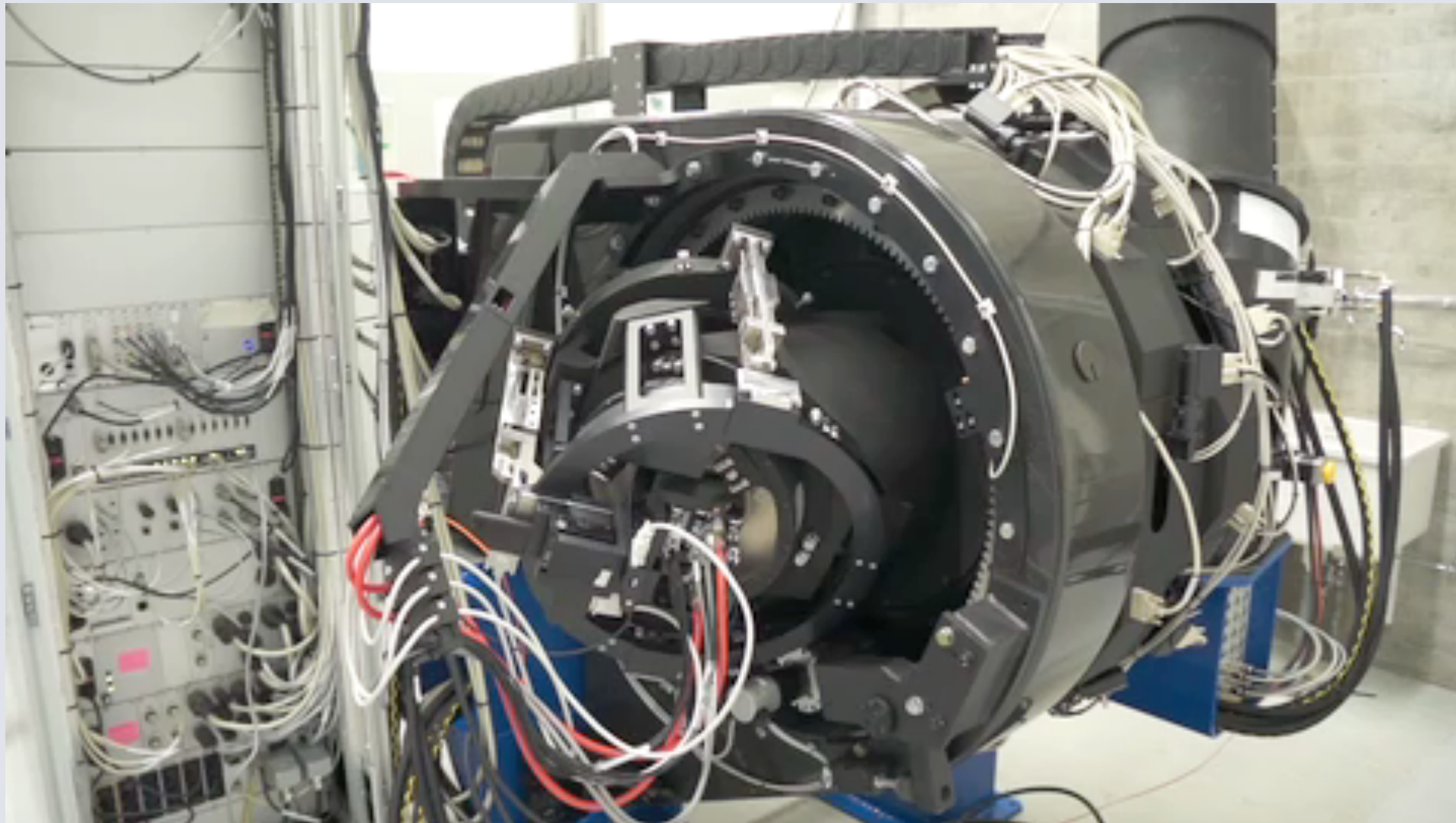


AstroTechTalk
26 September 2014
MPIA Heidelberg

Wavefront Sensors 1 Ground Layer



AstroTechTalk
26 September 2014
MPIA Heidelberg

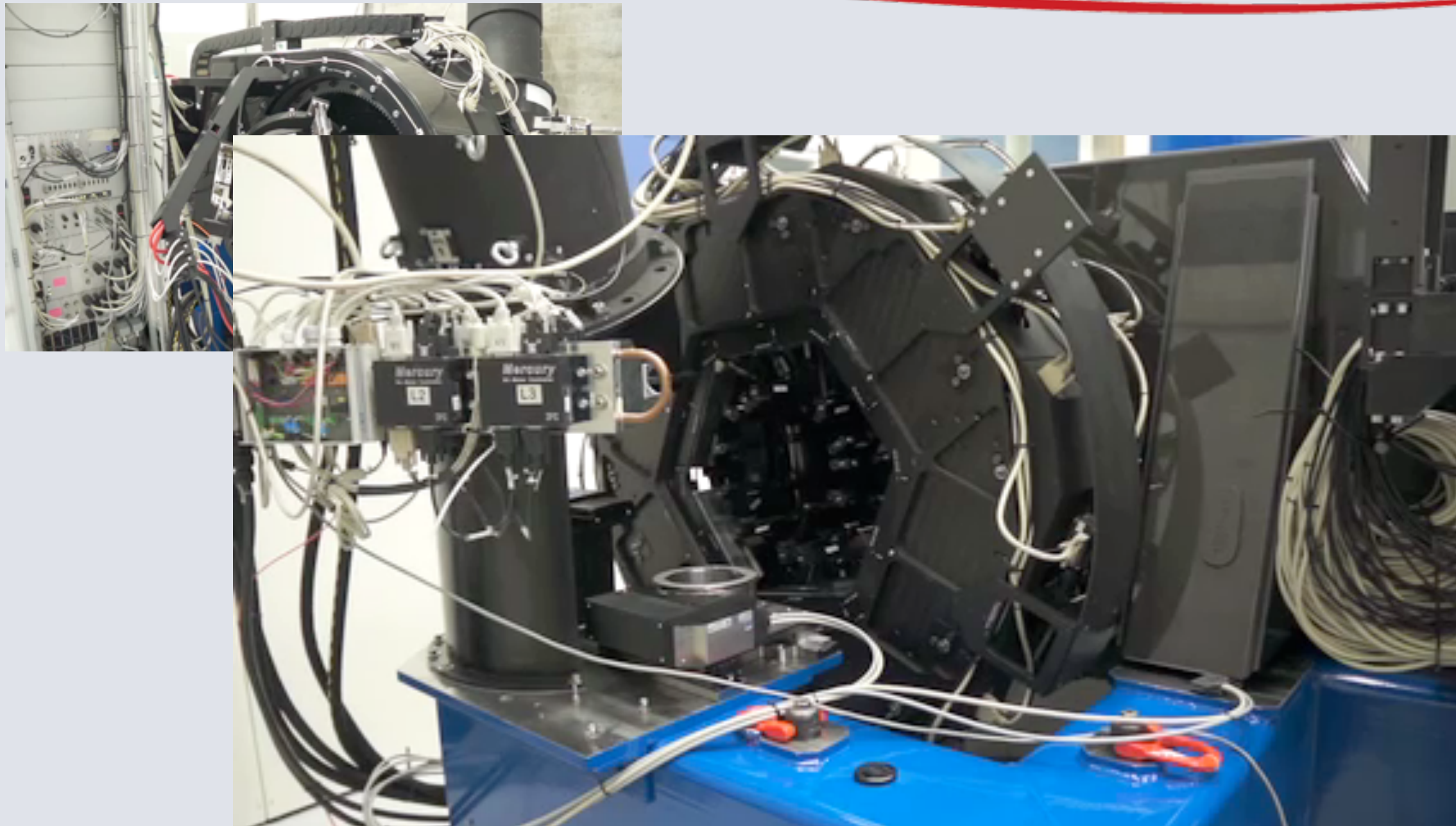


Sky Tracking

Wavefront Sensors 1 Ground Layer



AstroTechTalk
26 September 2014
MPIA Heidelberg



Sky Tracking

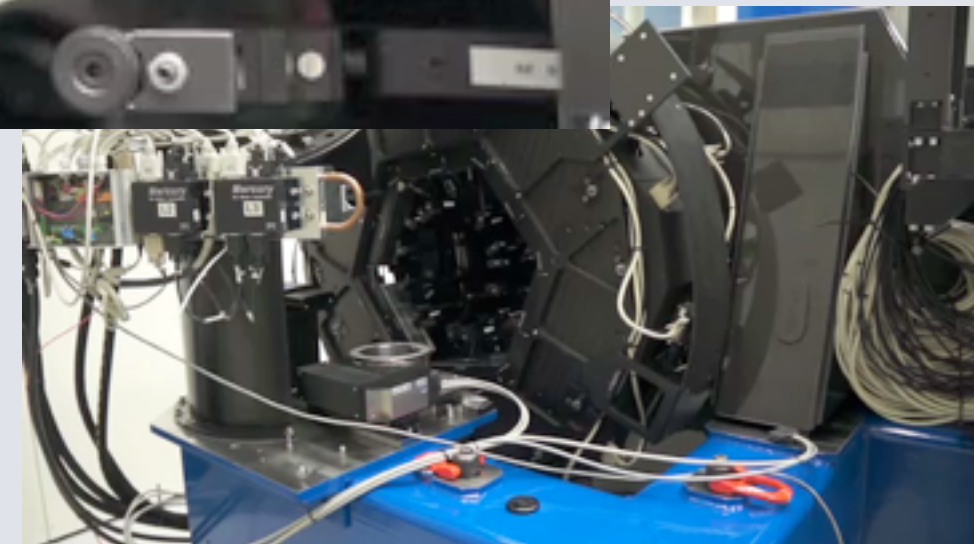
Wavefront Sensors 1 Ground Layer



AstroTechTalk
26 September 2014
MPIA Heidelberg



Star Finding



Wavefront Sensors 2



AstroTechTalk
26 September 2014
MPIA Heidelberg

Wavefront Sensors 2 High Layer



AstroTechTalk
26 September 2014
MPIA Heidelberg

The Operator GUI displays a schematic of the wavefront sensor layout and a control panel. The schematic shows eight sensors (SE01-SE08) arranged around a central focal plane (F) with a coordinate system (x, y). Each sensor is connected to a target (T01-T08). The control panel includes a table of coordinates, a 'Move to Target Tx' section with input fields and buttons, and a status bar at the bottom.

SE	X	Y
SE01	-48000.00	-23000.00
SE02	-48000.00	23000.00
SE03	23000.00	-48000.00
SE04	-23000.00	-48000.00
SE05	48000.00	23000.00
SE06	48000.00	-23000.00
SE07	-23000.00	48000.00
SE08	23000.00	48000.00

Move to Target Tx

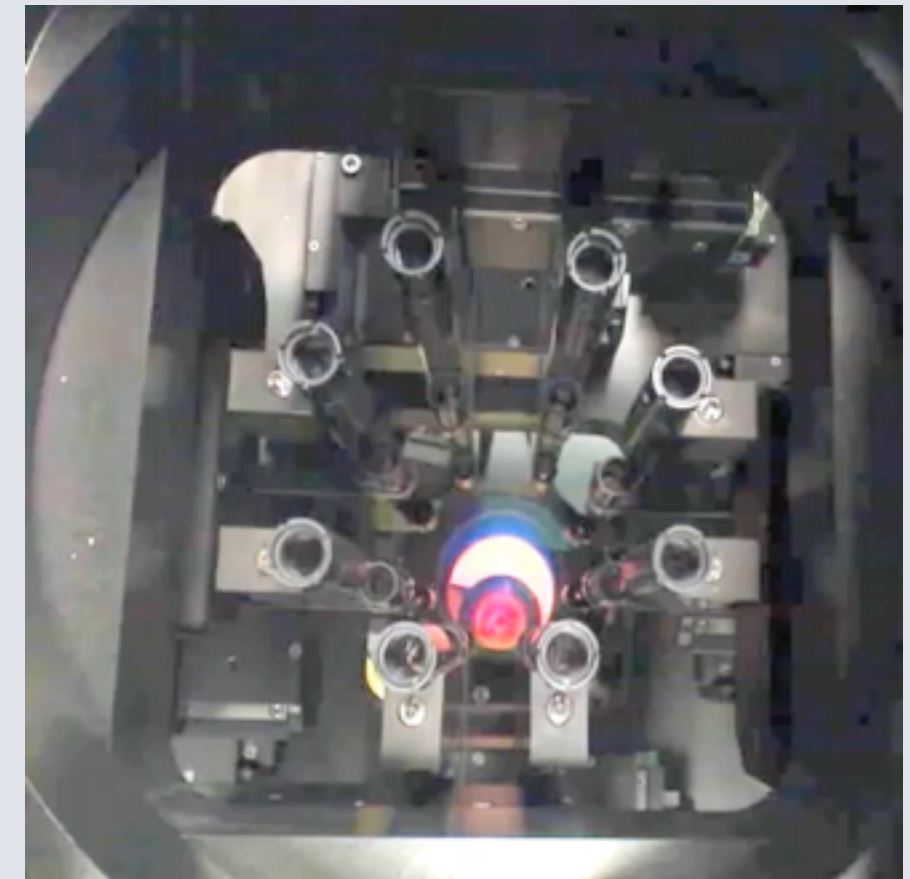
SE	Target	X	Y
SE01	T01	-35000	-13000
SE02	T02	-15000	13000
SE03	T03	13000	28000
SE04	T04	-13000	18000
SE05	T05	38000	8000
SE06	T06	28000	13000
SE07	T07	-13000	35000
SE08	T08	13000	10000

Target Name

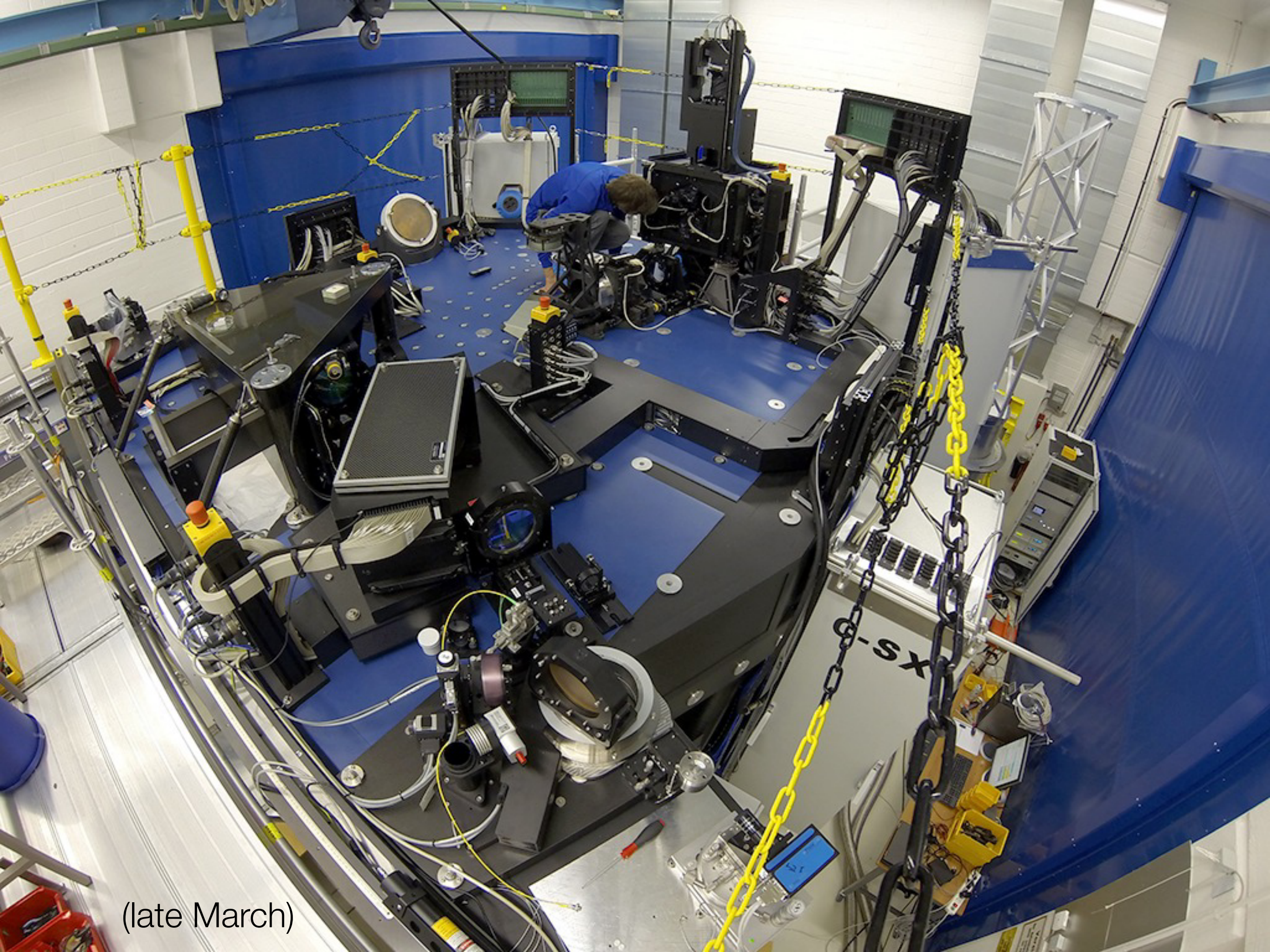
Move Absolute
Move Relative

READY

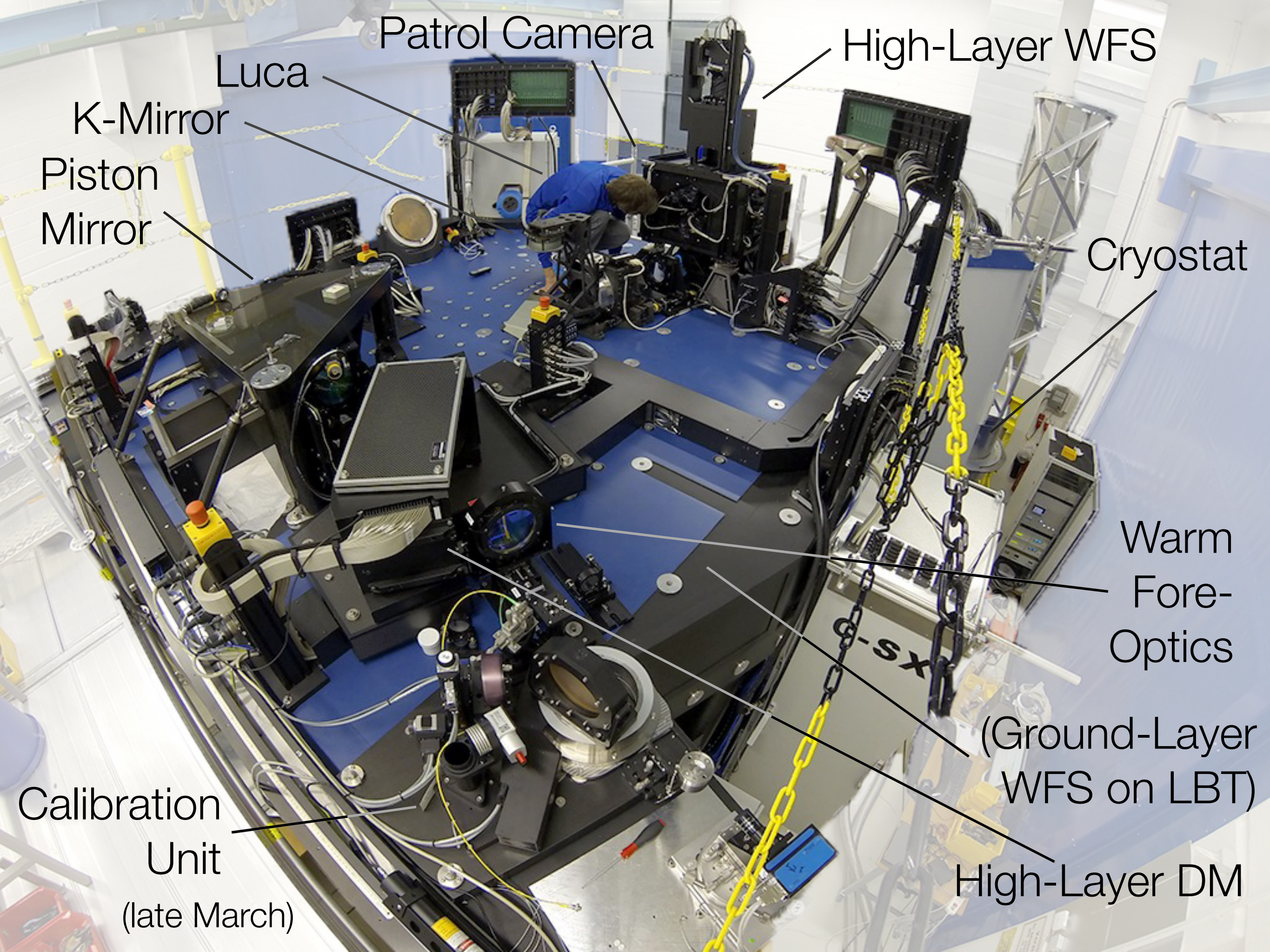
Operator GUI



Star Finding



(late March)



Patrol Camera

High-Layer WFS

Luca

K-Mirror

Piston Mirror

Mirror

Cryostat

Warm

Fore-

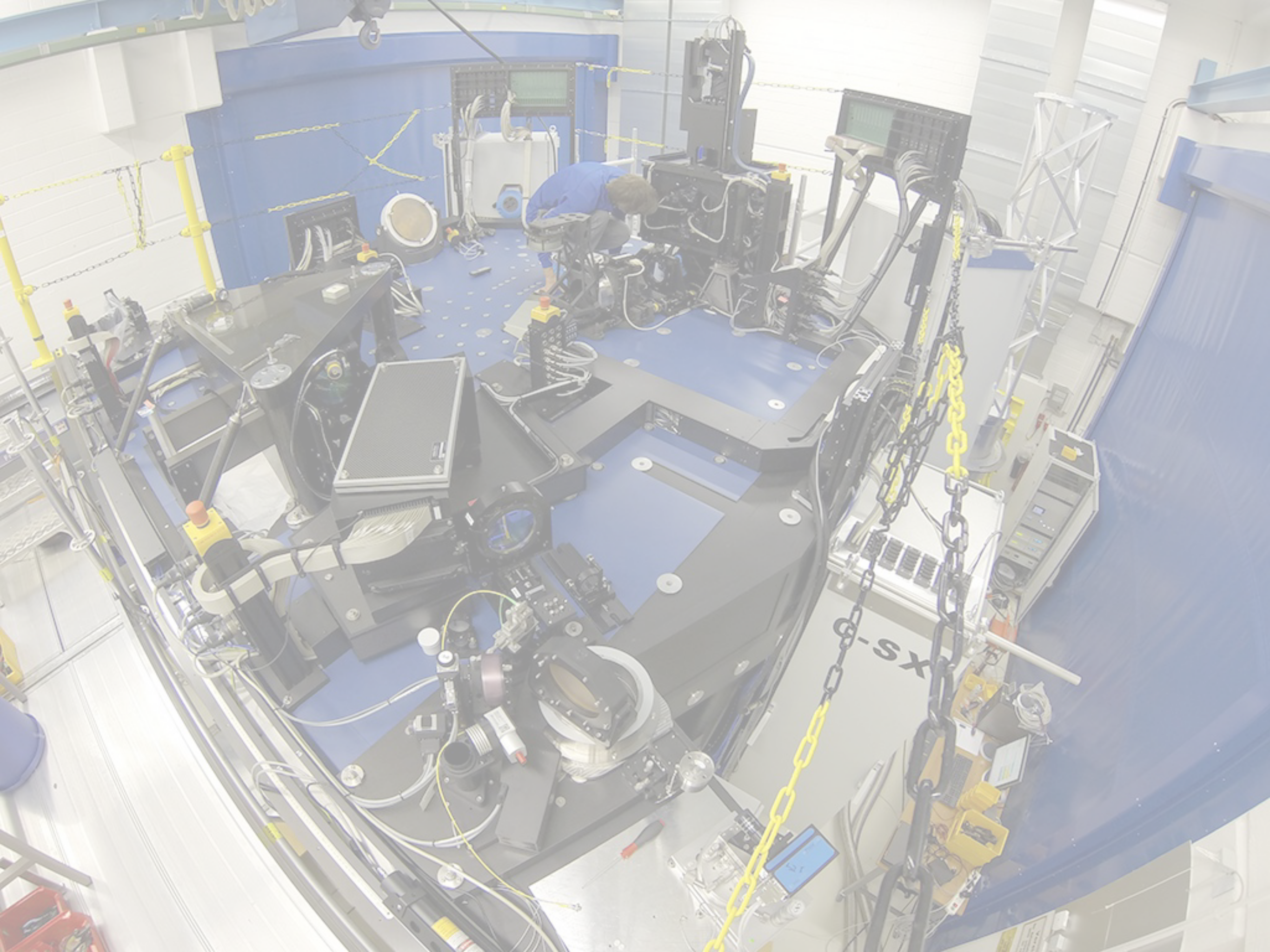
Optics

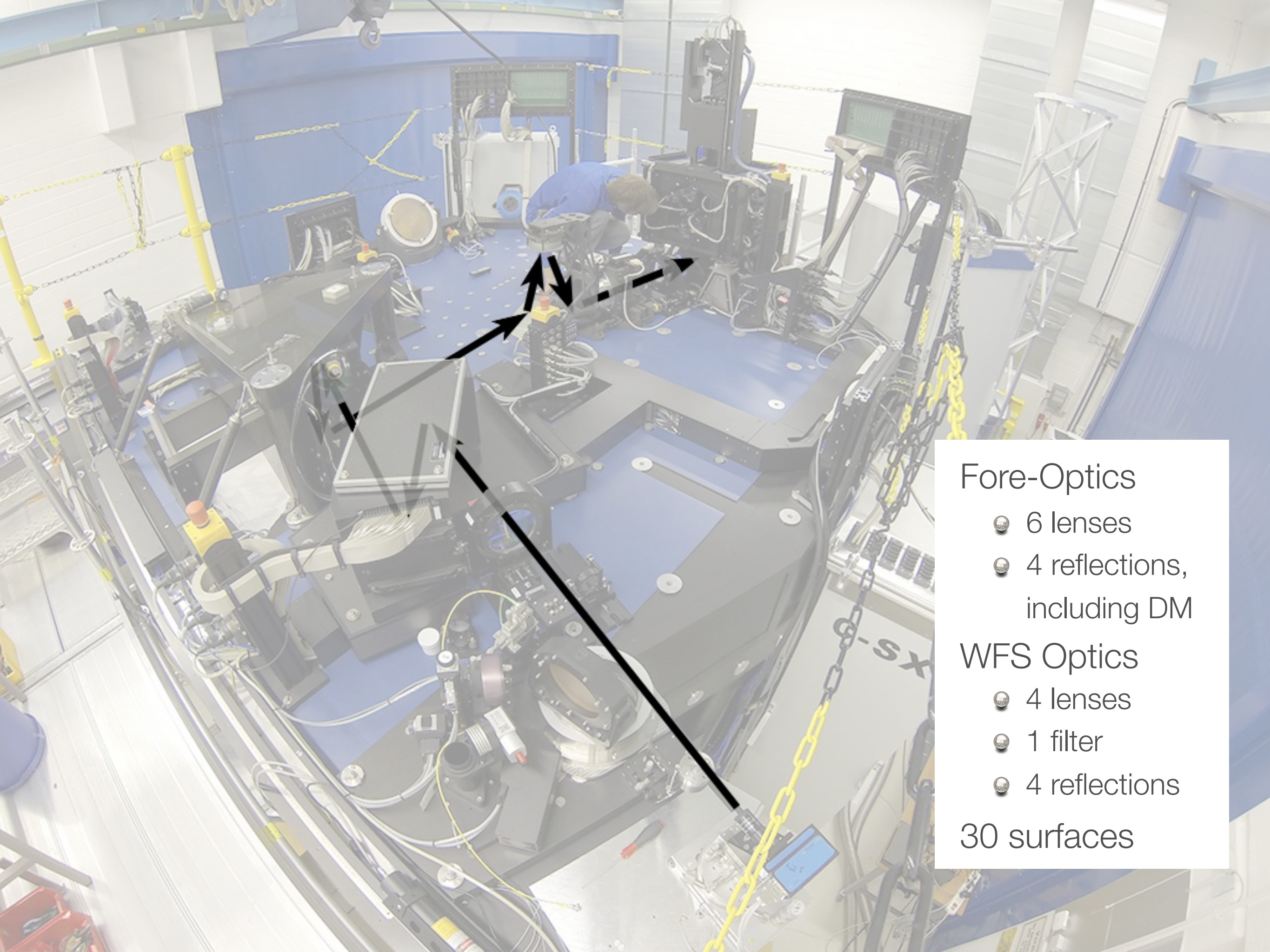
(Ground-Layer
WFS on LBT)

High-Layer DM

Calibration
Unit

(late March)





Fore-Optics

- 6 lenses
- 4 reflections, including DM

WFS Optics

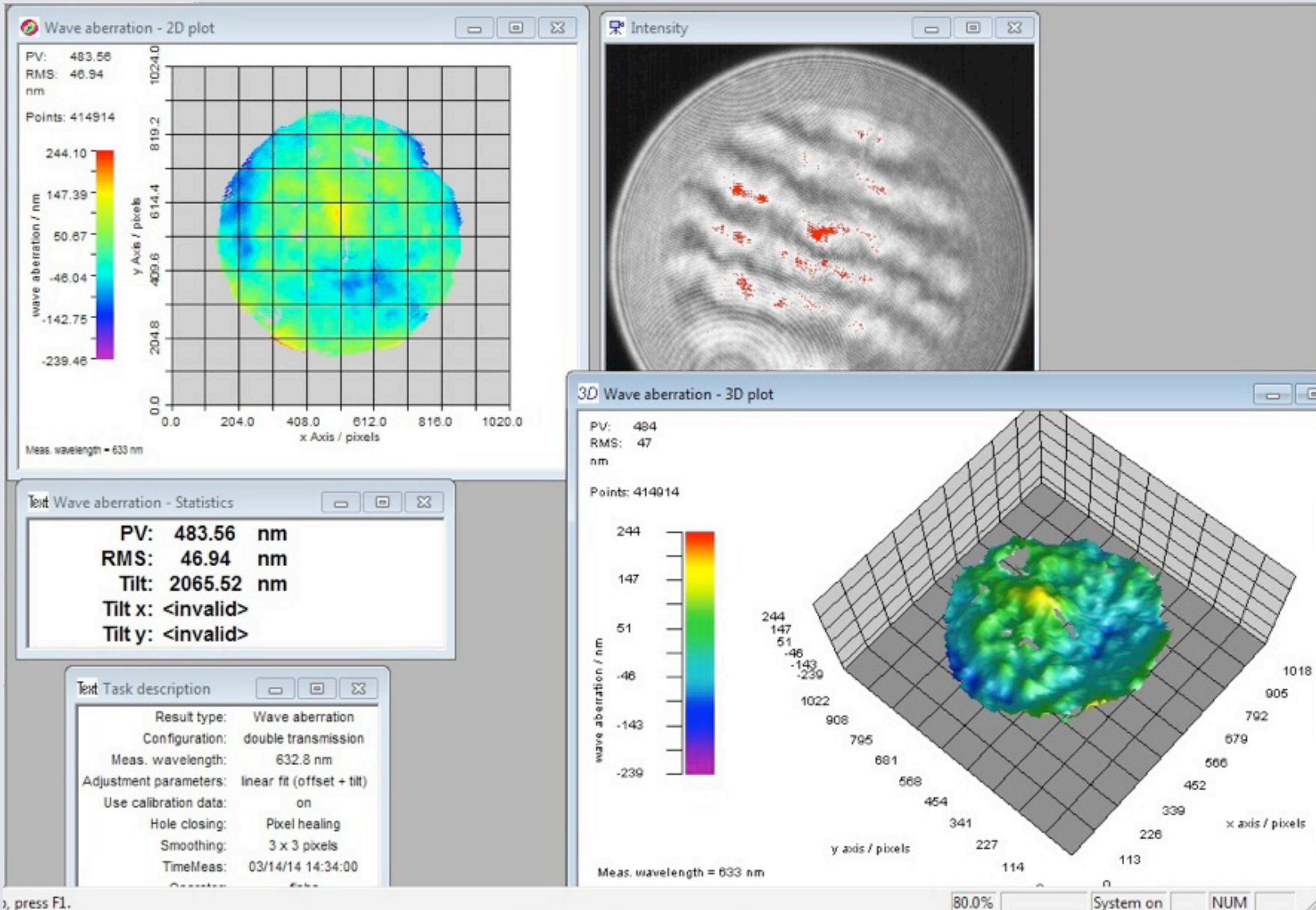
- 4 lenses
- 1 filter
- 4 reflections

30 surfaces

Delivered Wavefront Quality



AstroTechTalk
26 September 2014
MPIA Heidelberg



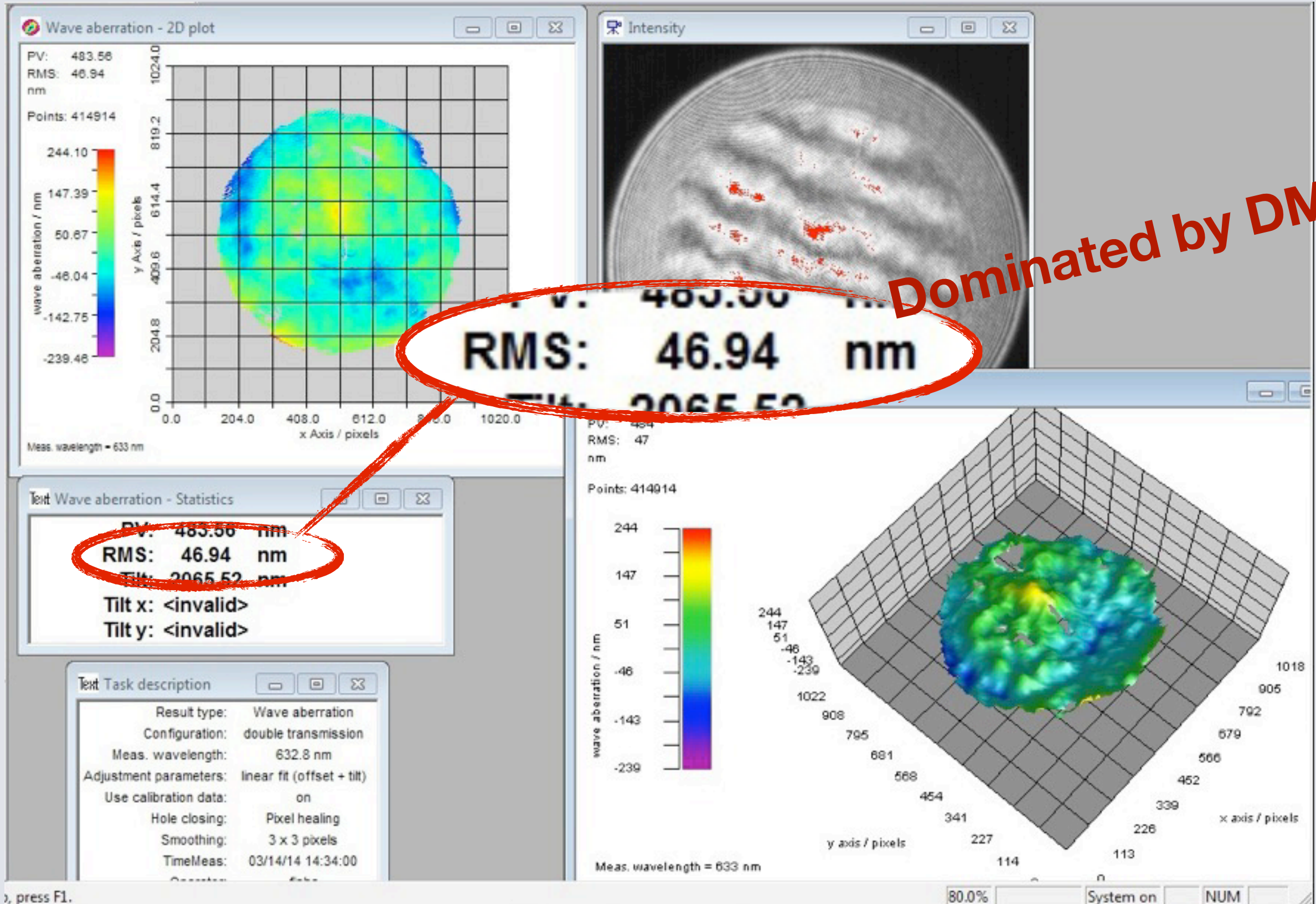
, press F1.

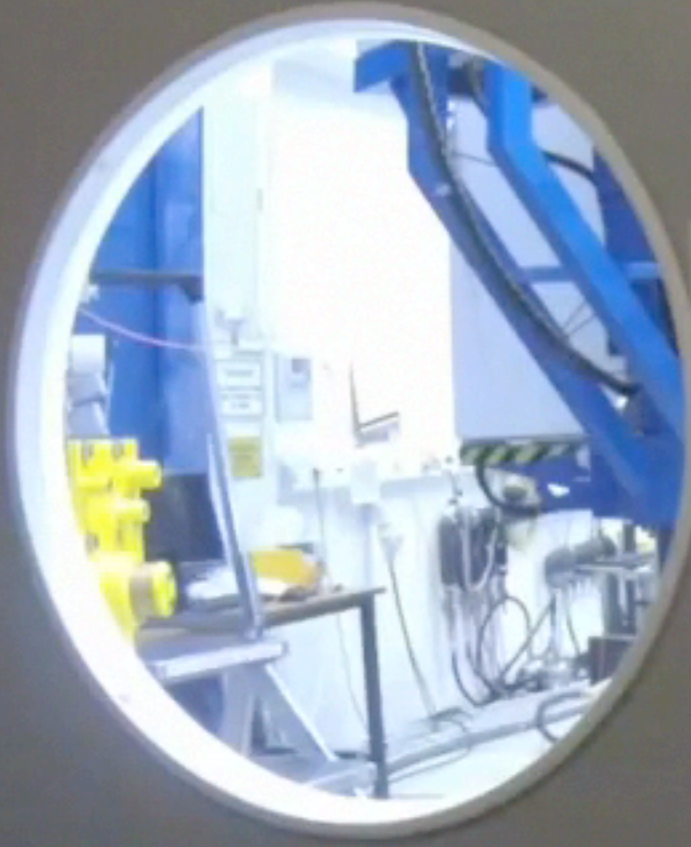
80.0%

System on

NUM

Delivered Wavefront Quality





Reinraum / Clean Room

