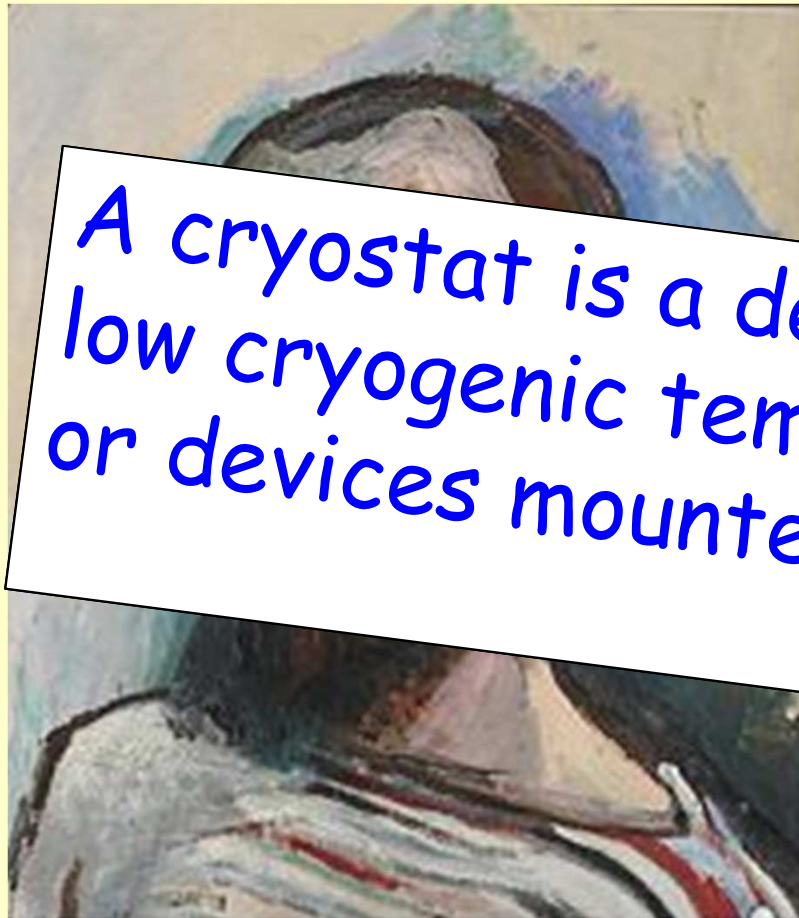


MATISSE cryostats



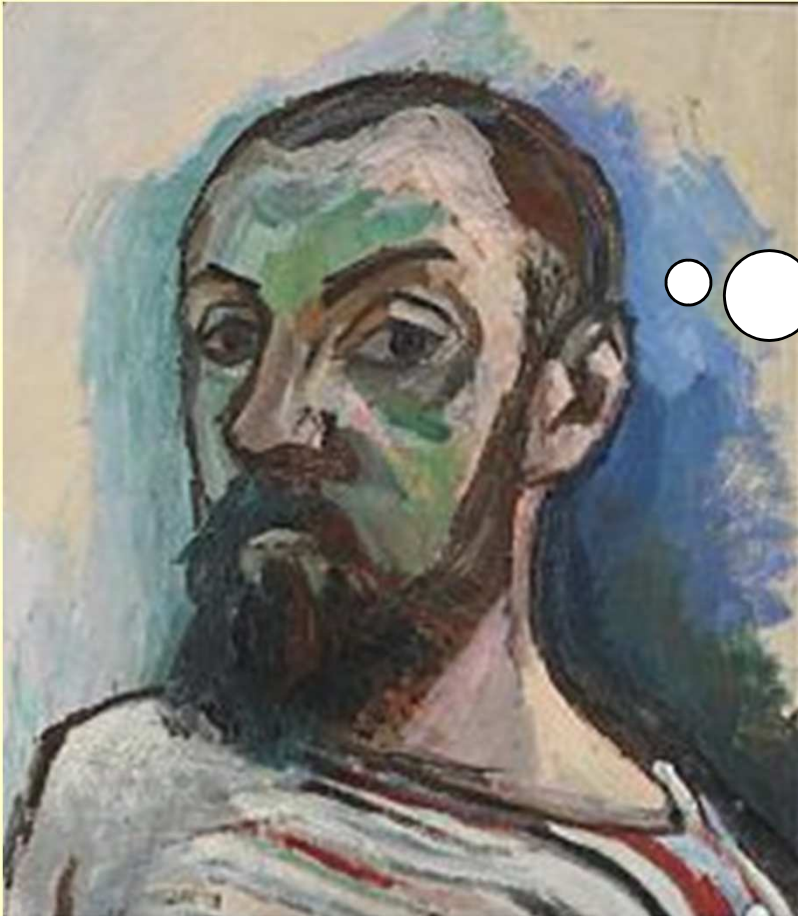
Cryostats?



What is a
A cryostat is a device used to maintain
low cryogenic temperatures of samples
or devices mounted within the cryostat.

wikipedia.org

Cryostat?

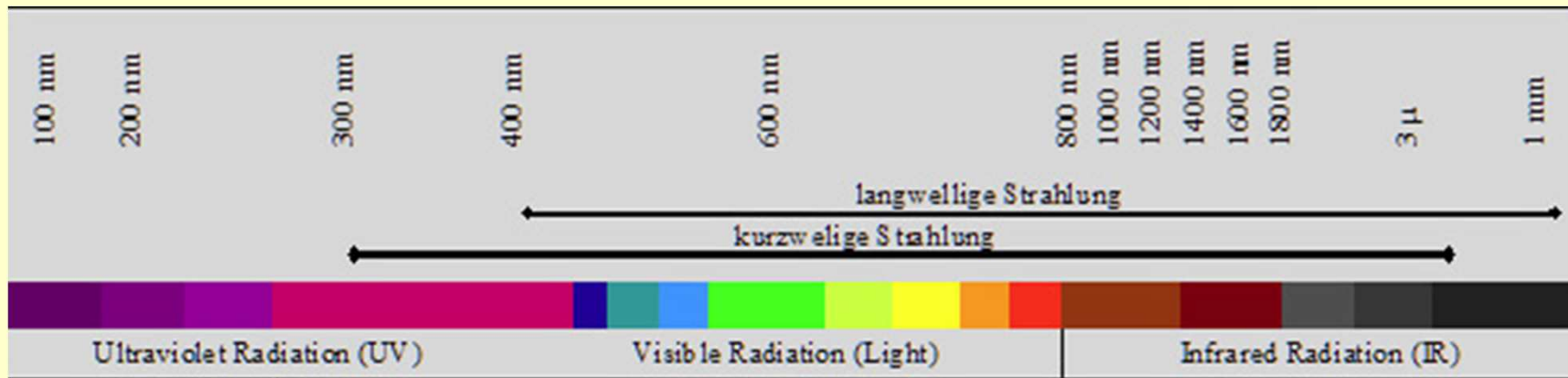


Why do we need
a cryostat?

Cryostat?

- Why do we need a cryostat?

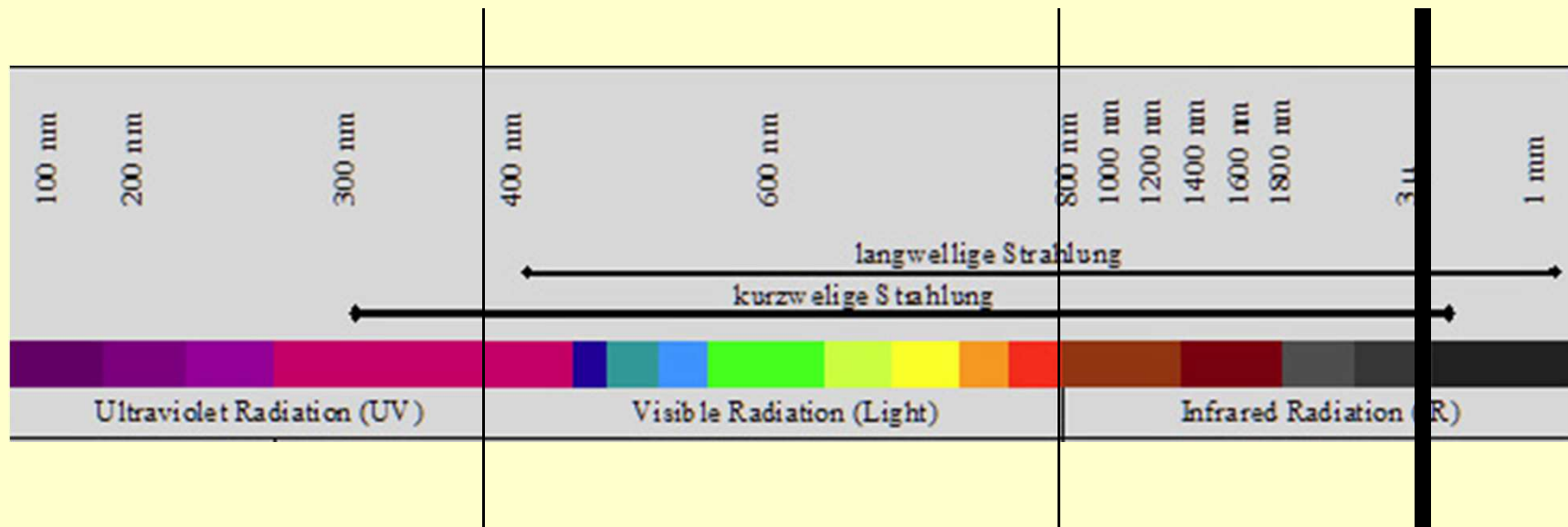
Spectrum from UV to IR



Cryostat?

- Why do we need a cryostat?

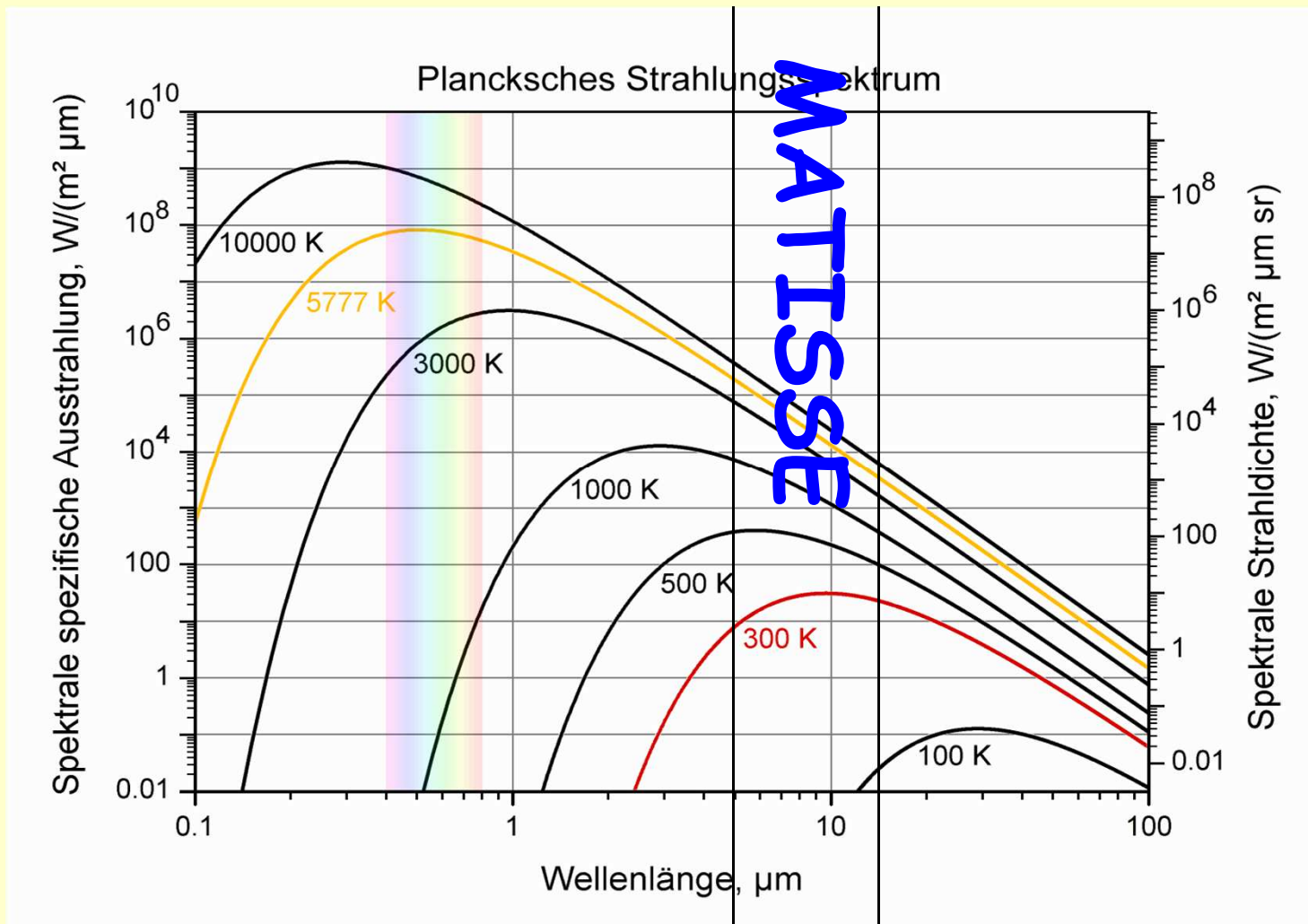
Spectrum from UV to NIR



MATISSE

Cryostat?

- Why do we need a cryostat?



Cryostat?

- Why do we need a cryostat?
 - IR detectors need cold environment
 - The longer the wavelength the colder the cryostat needs to be

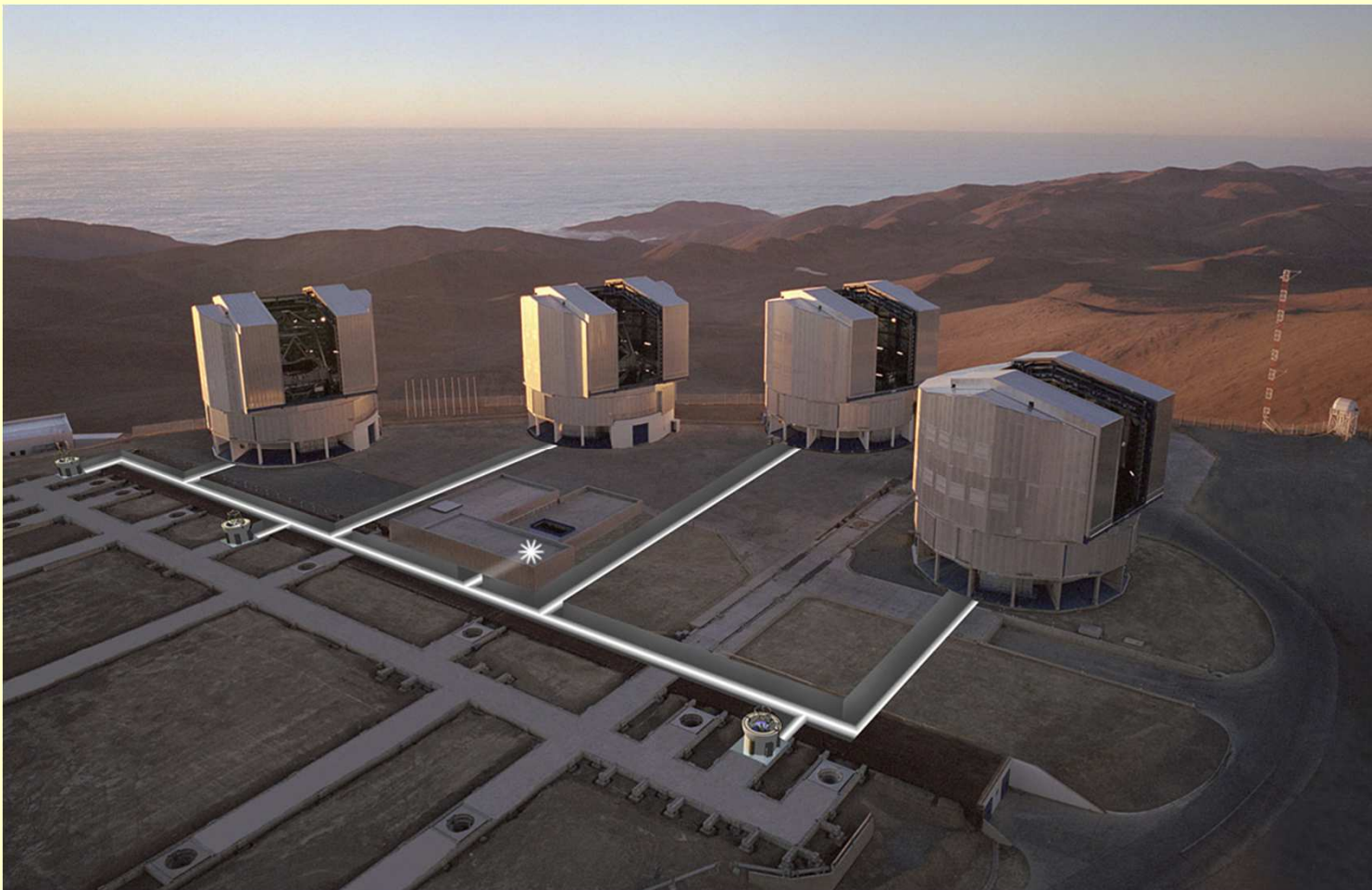
MATISSE?

- What does MATISSE do?
 - Interferometric observation with up to 4 telescopes

MATISSE is designed to be a mid-infrared interferometric instrument combining the beams of up to four telescopes (UTs or ATs) of the European Southern Observatory Very Large Telescope Interferometer (VLTI).

MATISSE?

- What does MATISSE do?
- Interferometric observation with up to 4 telescopes

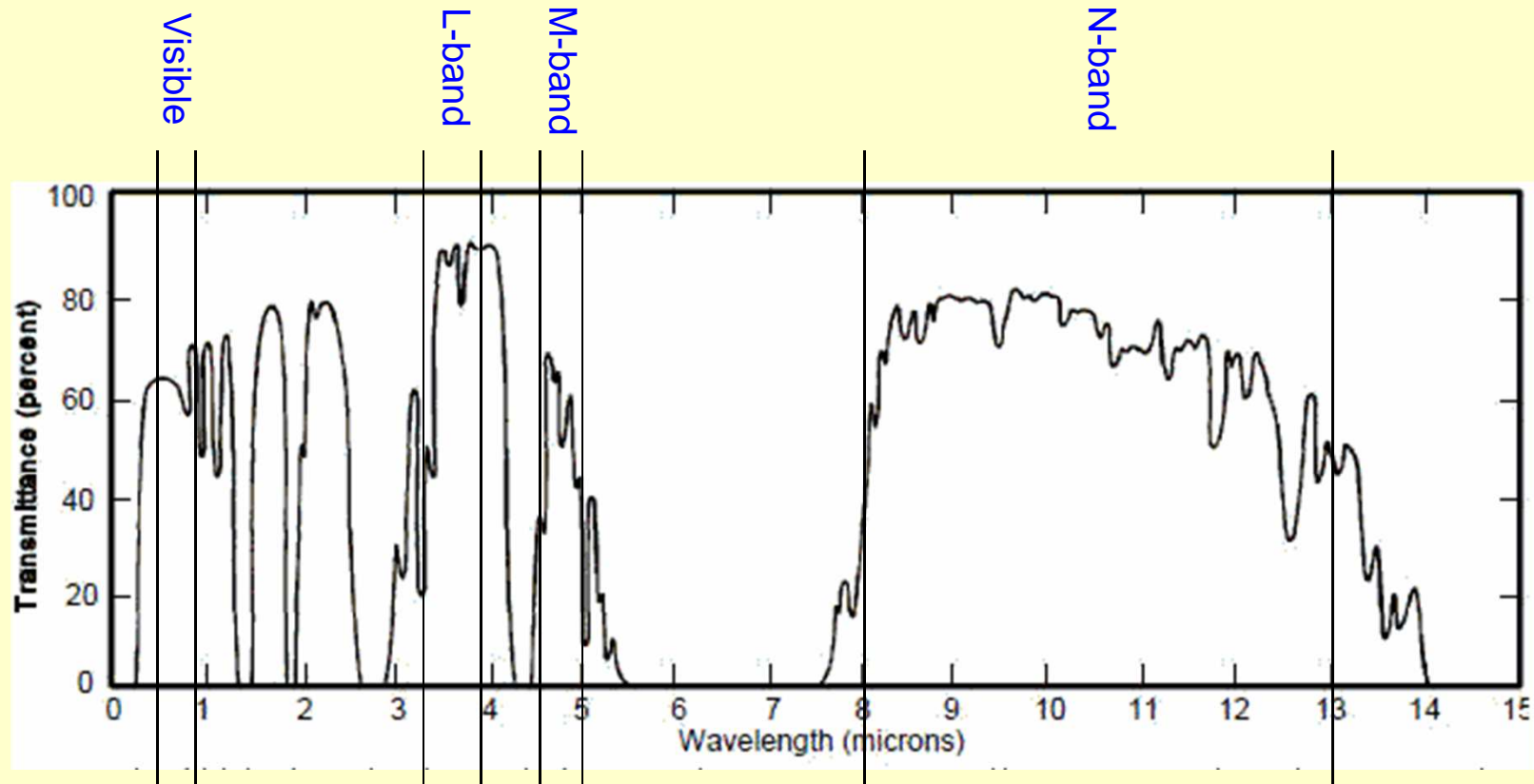


MATISSE?

- What does MATISSE do?
 - Interferometric observation with up to 4 telescopes
 - Wavelength range LM band
 - Wavelength range N band

MATISSE?

Atmospheric transmission



Human eye

Hawaii2RG

Aquarius

MATISSE?

- What does MATISSE do?
 - Interferometric observation with up to 4 telescopes
 - Wavelength range LM band with Hawaii 2 RG detector
 - Wavelength range N band with Aquarius detector

MATISSE team

- Team (hardware parts only)
 - OCA, Nice, France
 - PI institute
 - Warm optics
 - Astron, Dwingeloo, Netherlands
 - Cold optics
 - ESO, Garching (MPIfR, Bonn)
 - Detectors
 - MPIA
 - Instrument control electronics
 - **Cryostats**

MATISSE history

- Successor of MIDI
 - Installation 2002
 - Decommissioning 2015
- Phase A kick-off Nov 2006
- Phase A final review July 2007
- PDR Dec 2010
- FDR optics and cryogenics Sep 2011
- Transport to Nice
 - LM-band cryostat July 2014
 - N-band cryostat Nov 2014
- **Transport to Paranal June 2017**







MATISSE cryostats



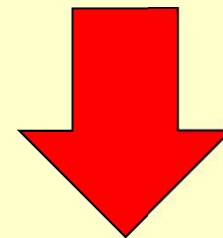
- Design driver
 - Temperature requirements
 - Space requirements
 - Vibration stability

MATISSE cryostats

- Temperature requirements

- LM-band optics 80 K 
- N-band optics 40 K 
- Hawaii2RG 40 K 
- Aquarius 8 K 

liquid nitrogen 77 K
 cooling machine
 cooling machine
 cooling machine



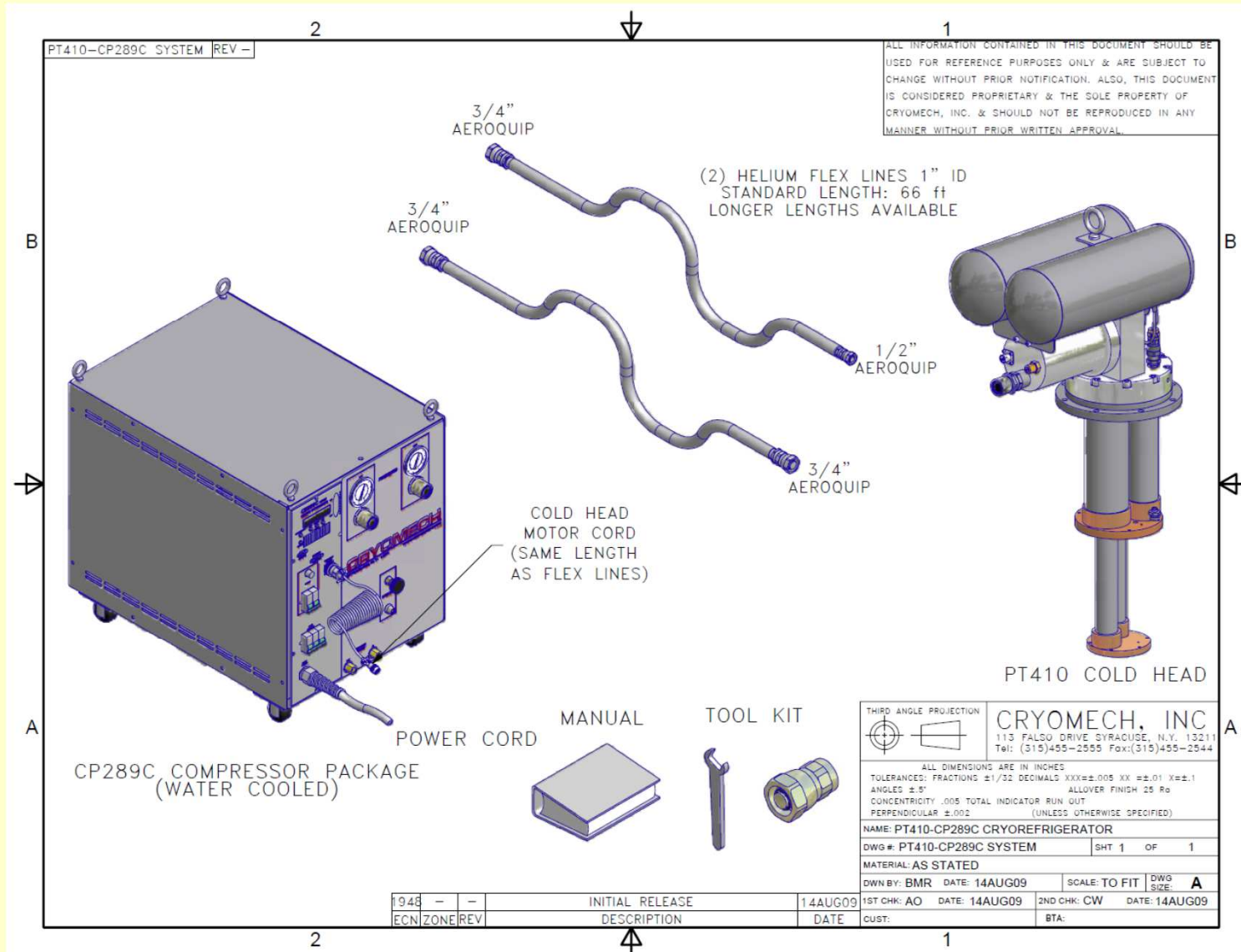
vibration



PT cooler

- Cooling with a pulse tube cooler
 - Low vibration
 - $T_{\min} < 40$ K (1st stage)
 - $T_{\min} < 6$ K (2nd stage)

PT cooler

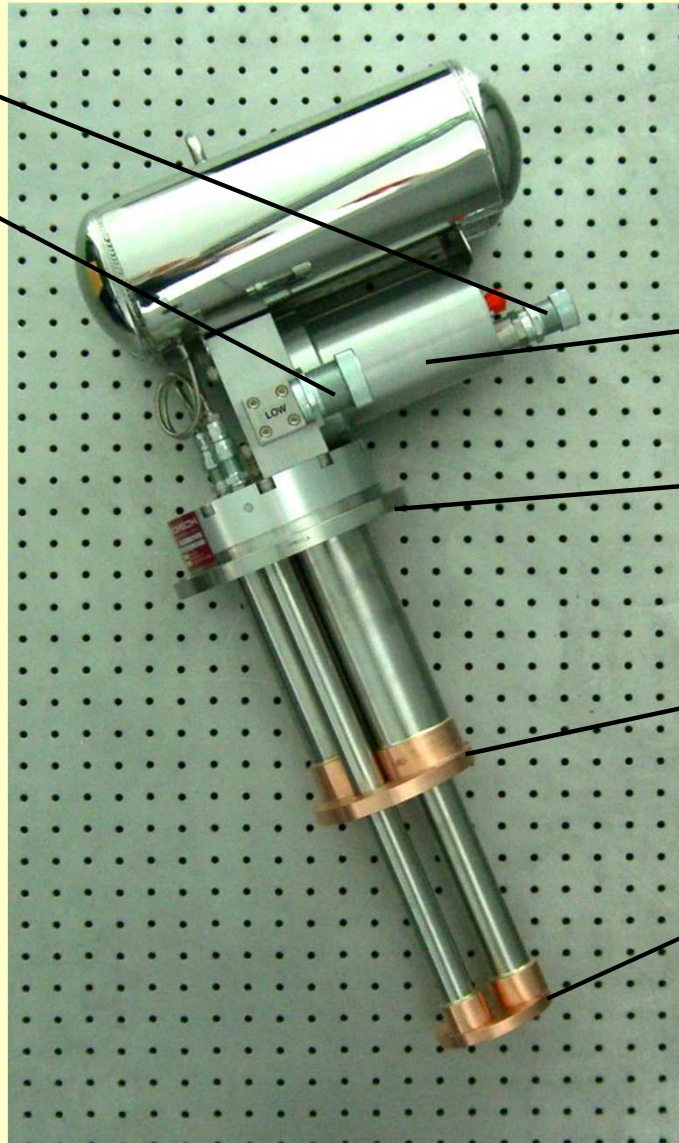


PT cooler

helium high pressure

helium low pressure

PT410
Cryomech



motor

vacuum flange
300 K

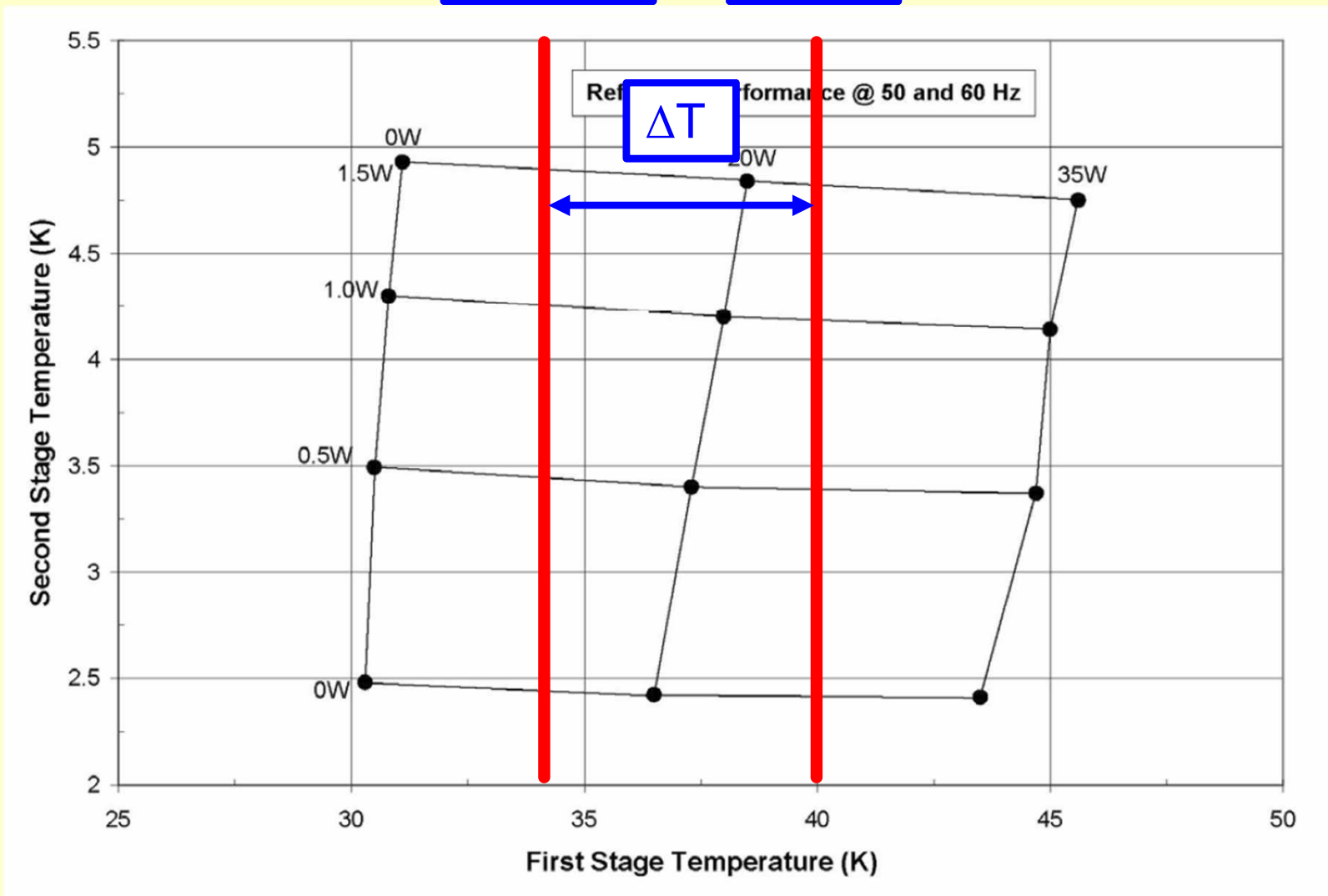
first stage
< 40 K

second stage
< 6 K

PT cooler

T cooler

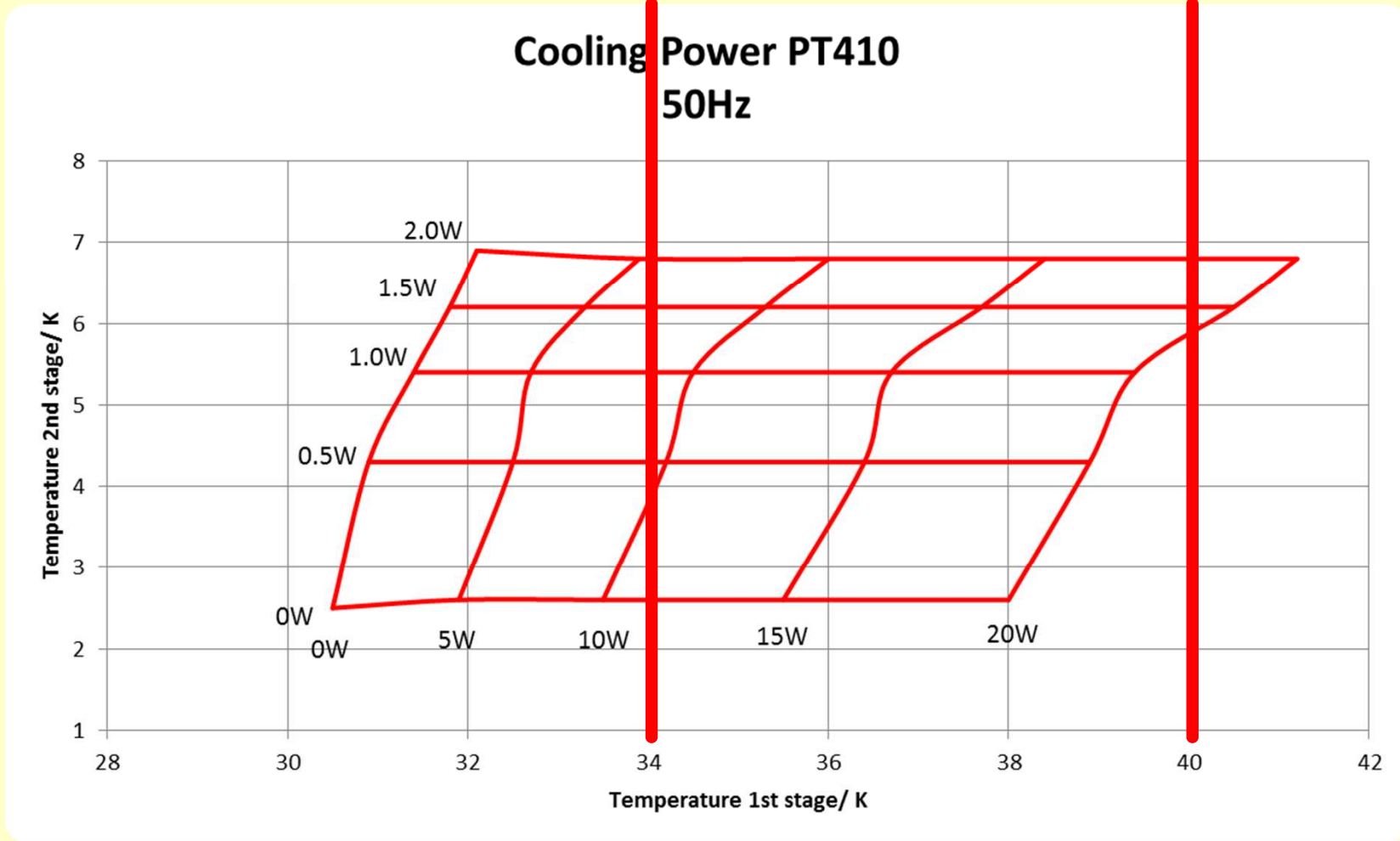
T max



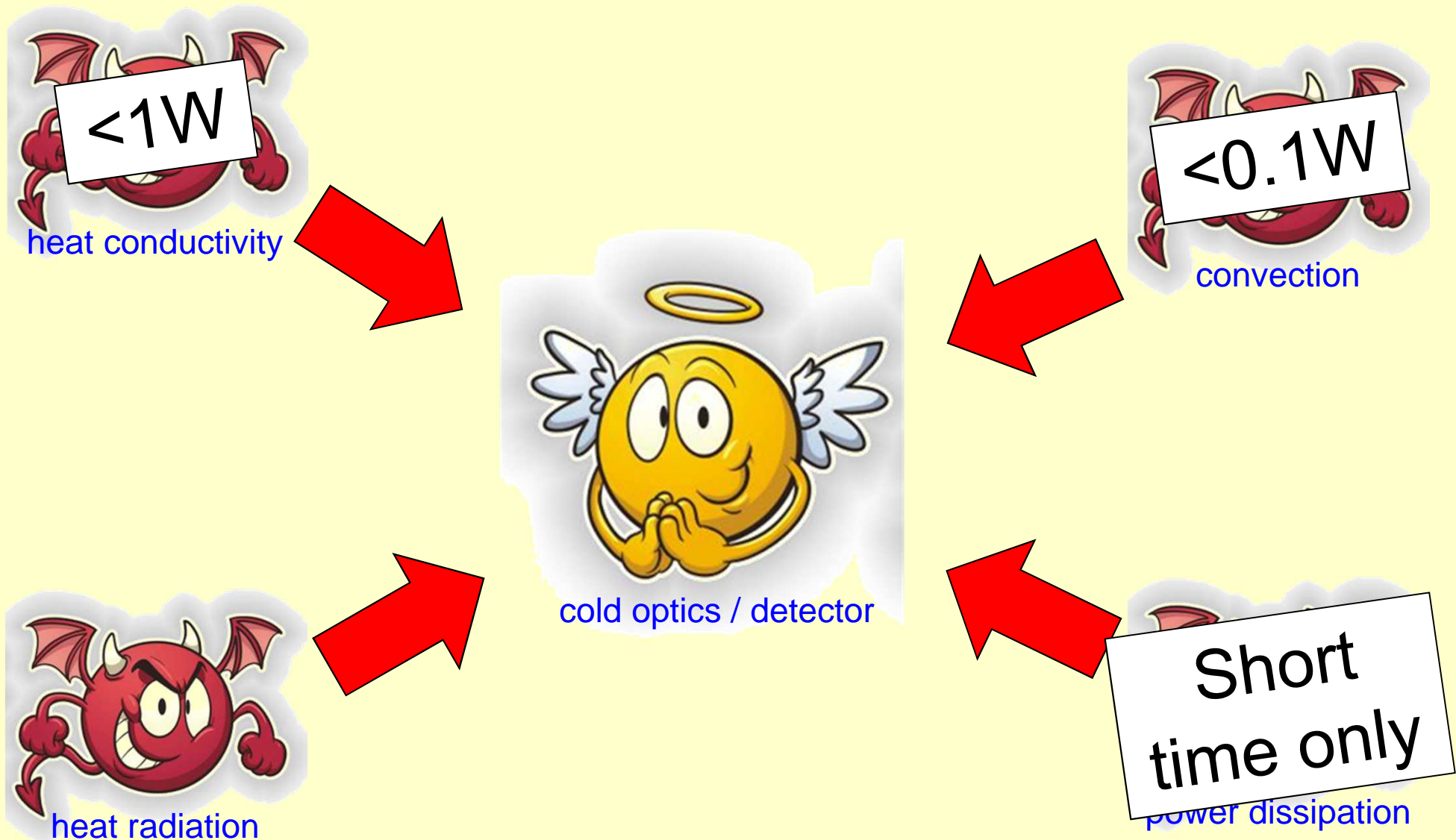
PT cooler

T cooler

T max

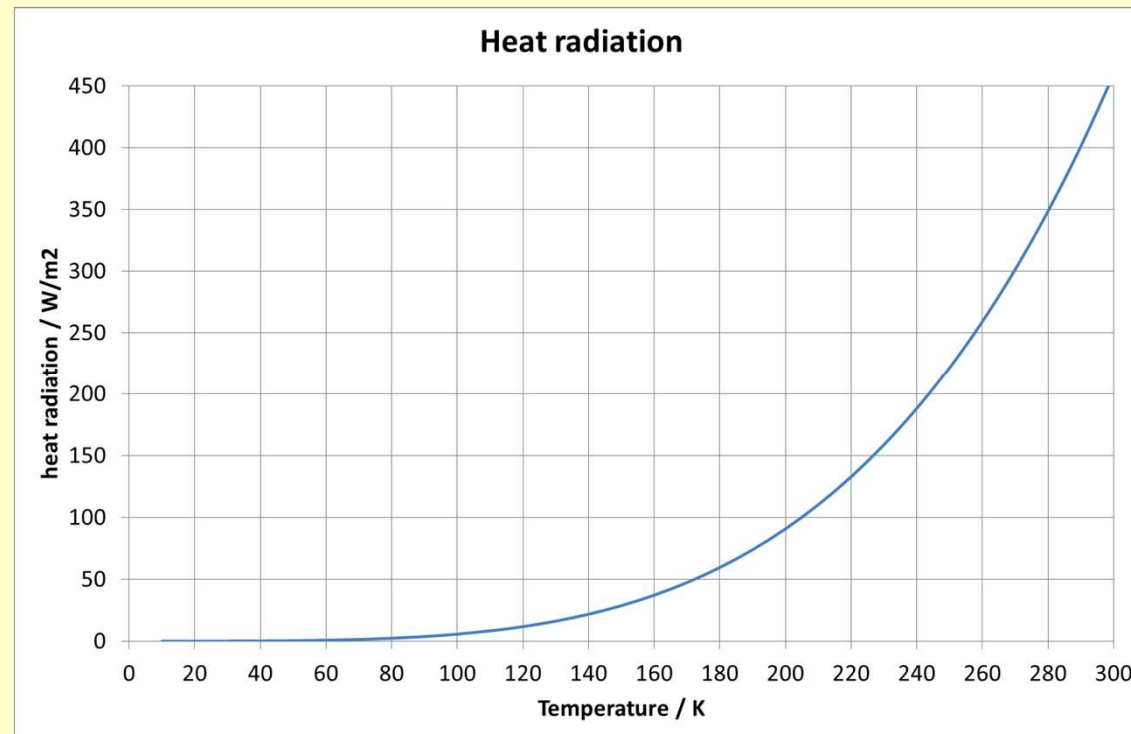


Cryostats heat input



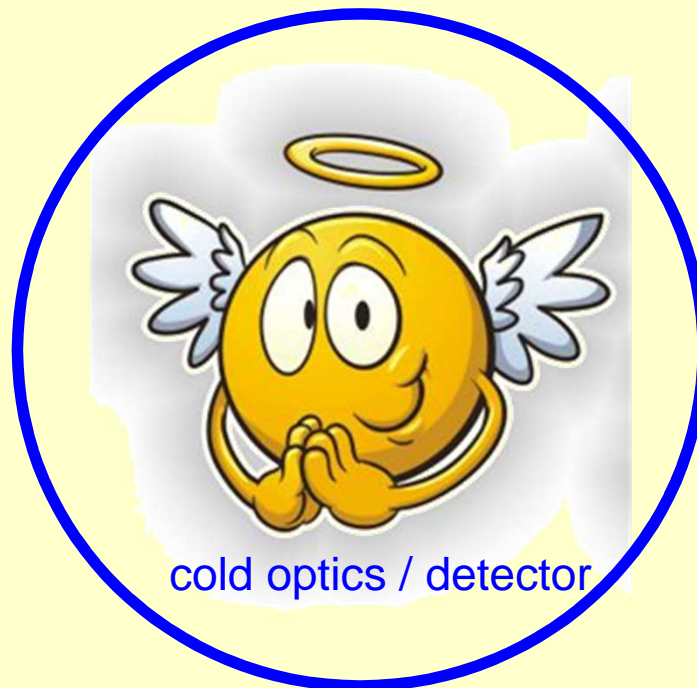
Cryostats heat input

- Heat radiation
 - Radiation from 290 K about 400 W/m²
 - With 90% reflection still 40 W/m²
 - Changing with T⁴

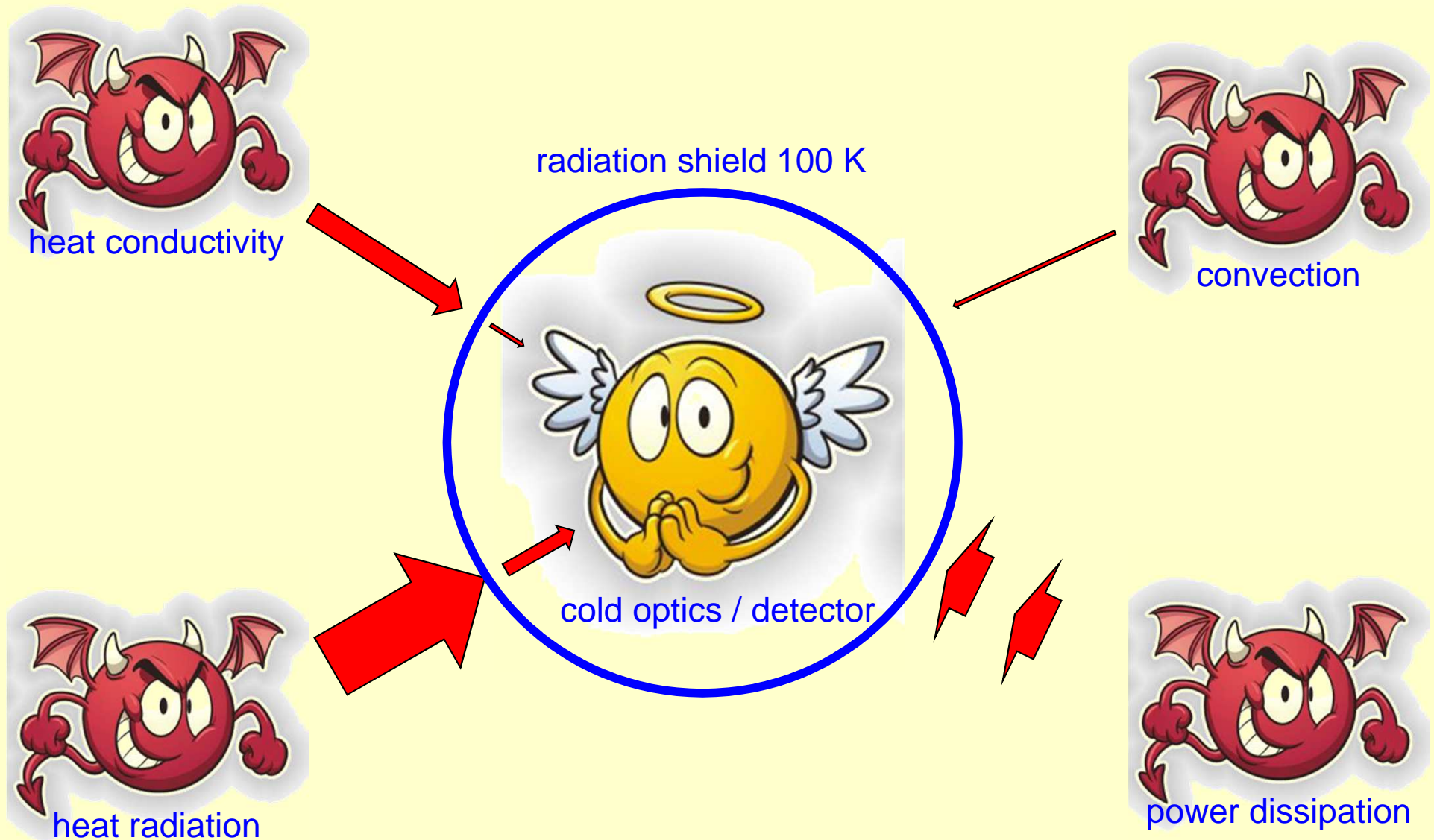


Cryostats heat input

Radiation shield cooled by LN₂



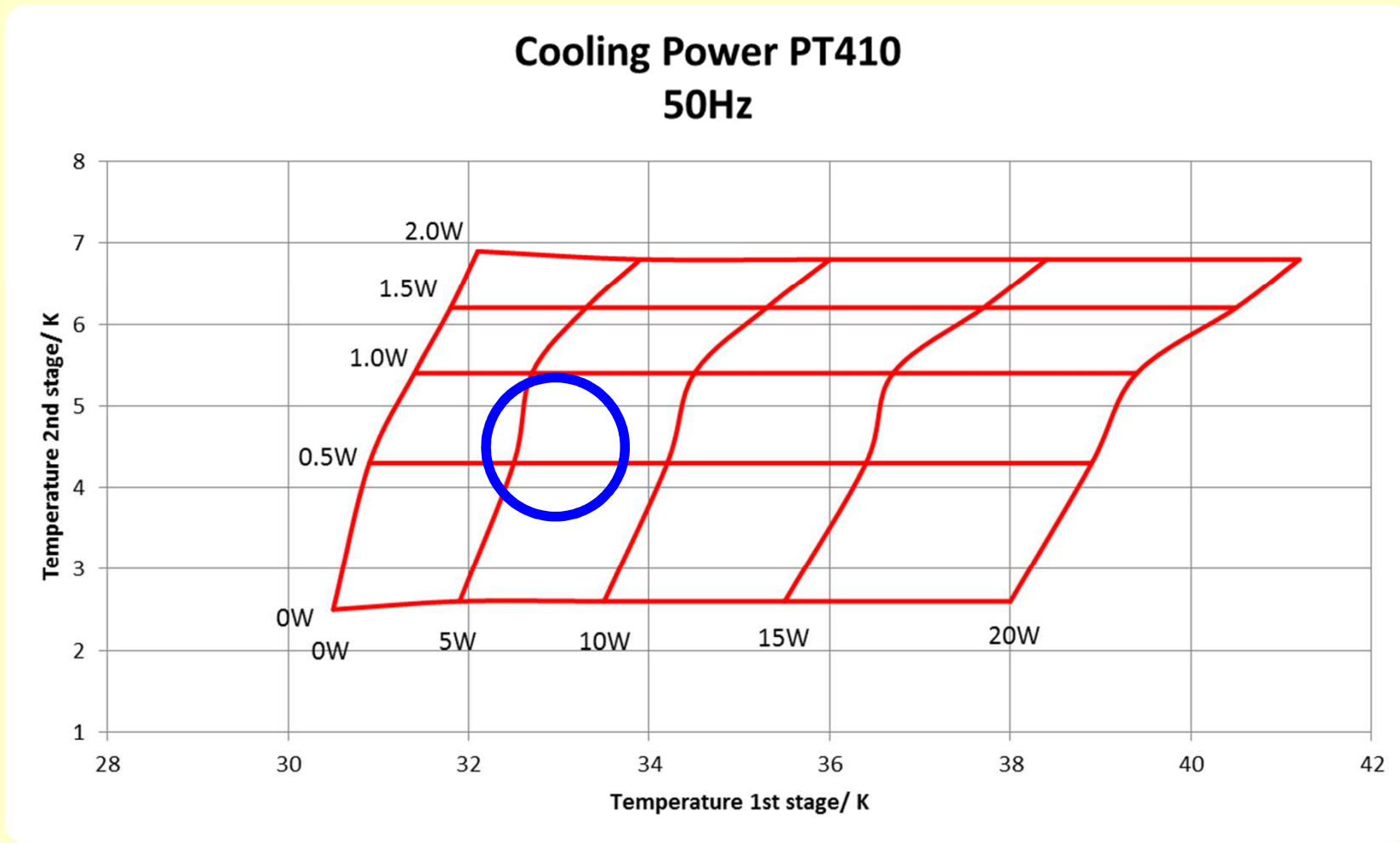
Cryostats heat input



Cryostats heat input

- Heat input
 - Radiation
 - On surface ○ 3 W
 - Through windows ○ 1.2 W
 - Conductivity
 - Spacer ○ <1 W
 - Cables ○ <1W
 - Dissipation
 - Convection
-
- 6.2 W

Cryostats heat input



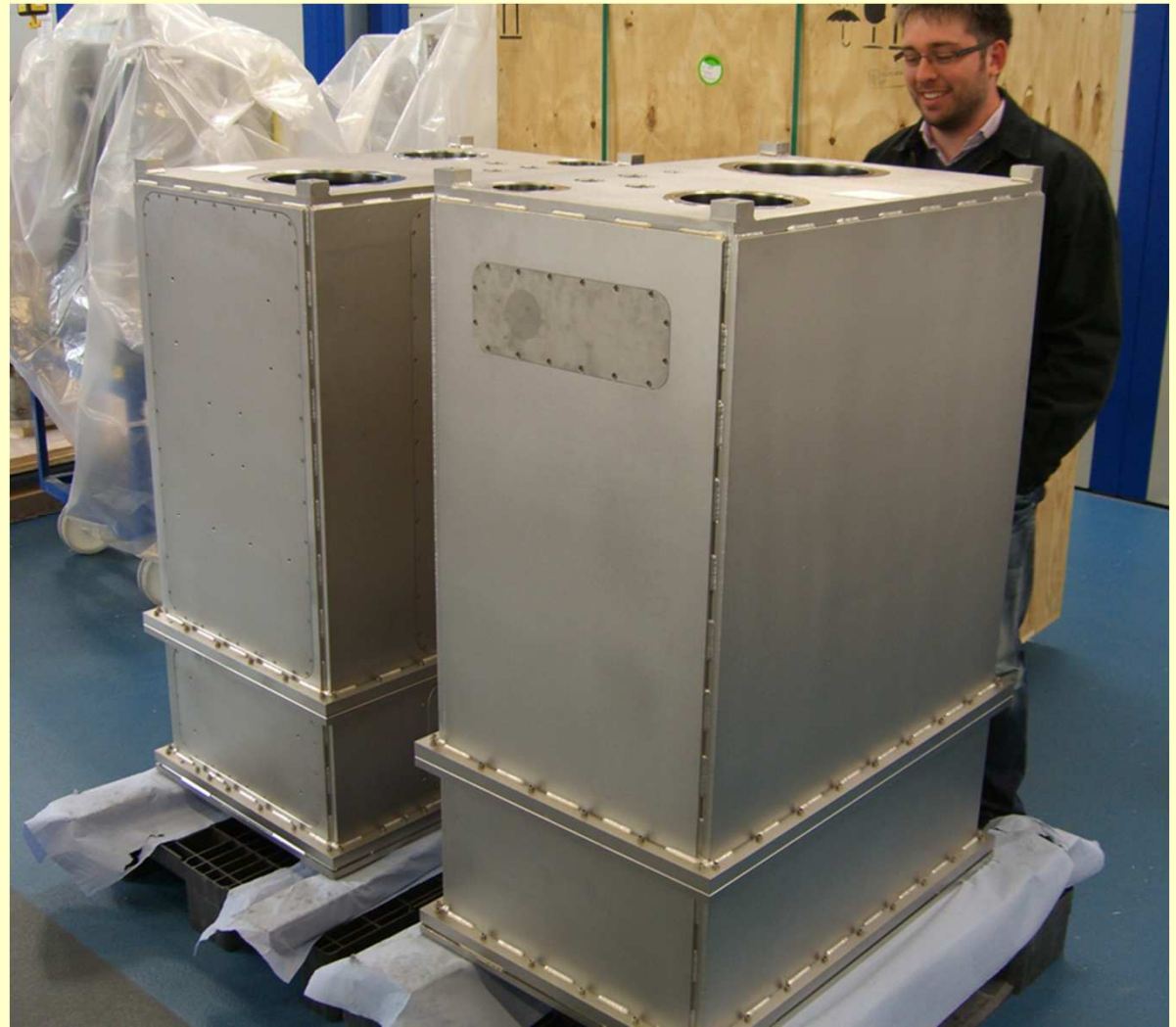
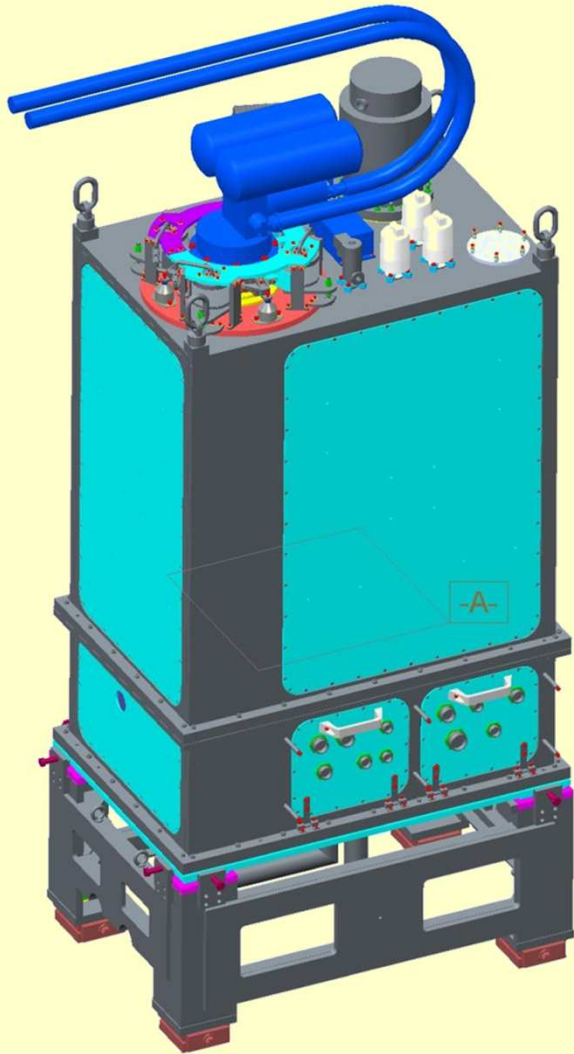
MATISSE cryostats



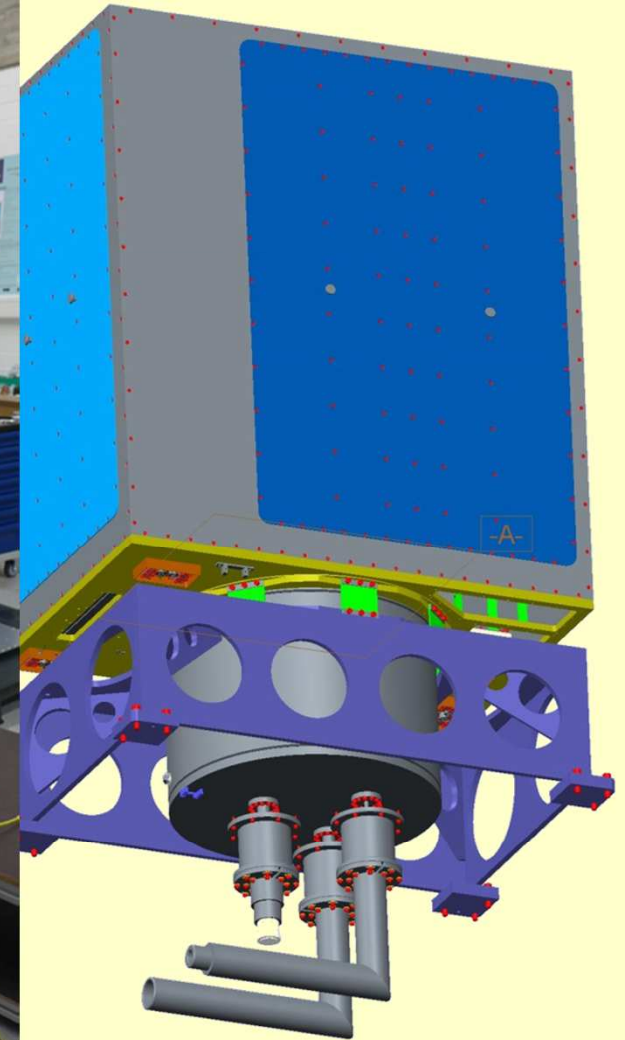
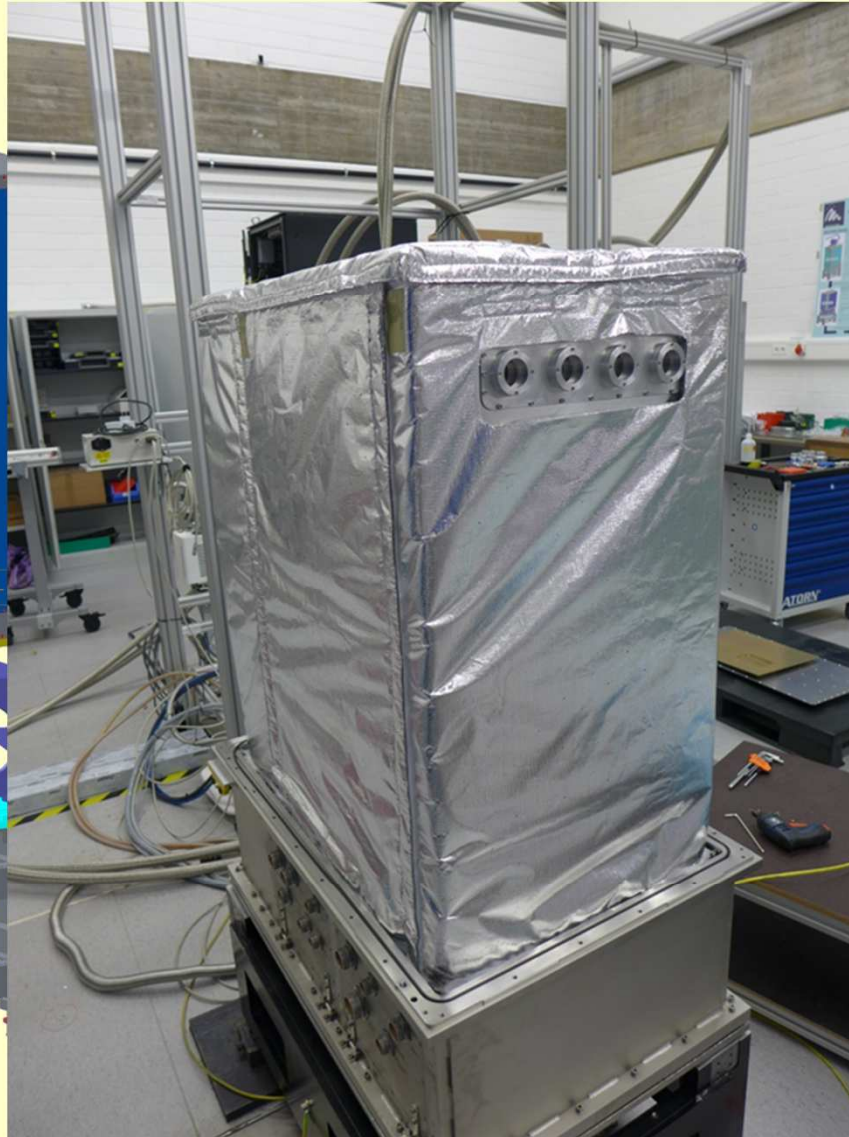
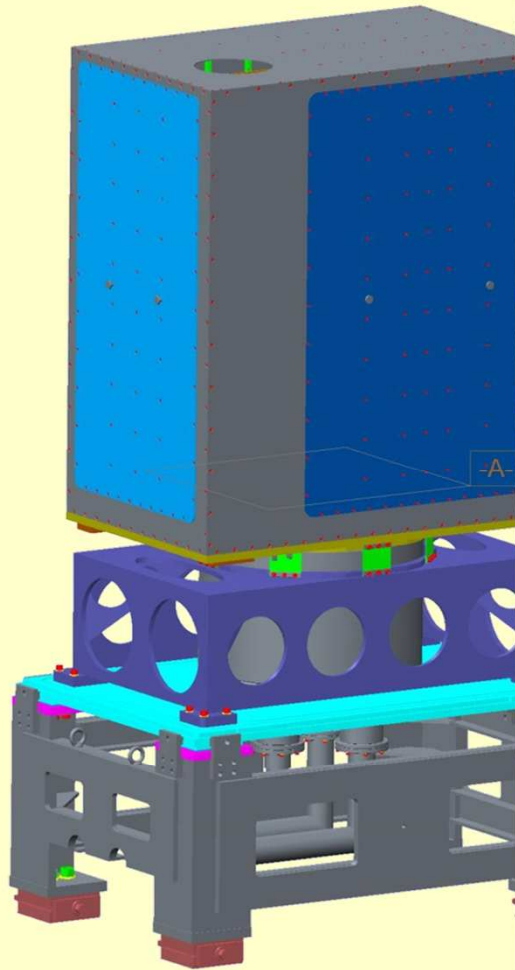
MATISSE cryostats



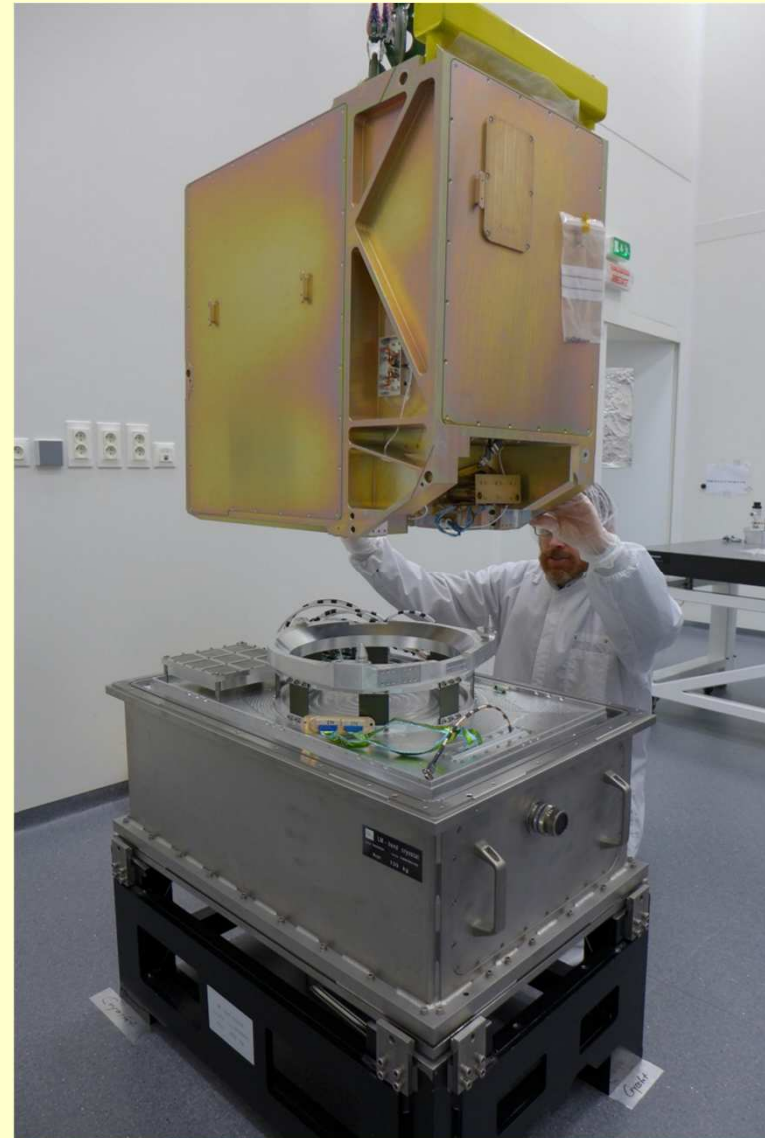
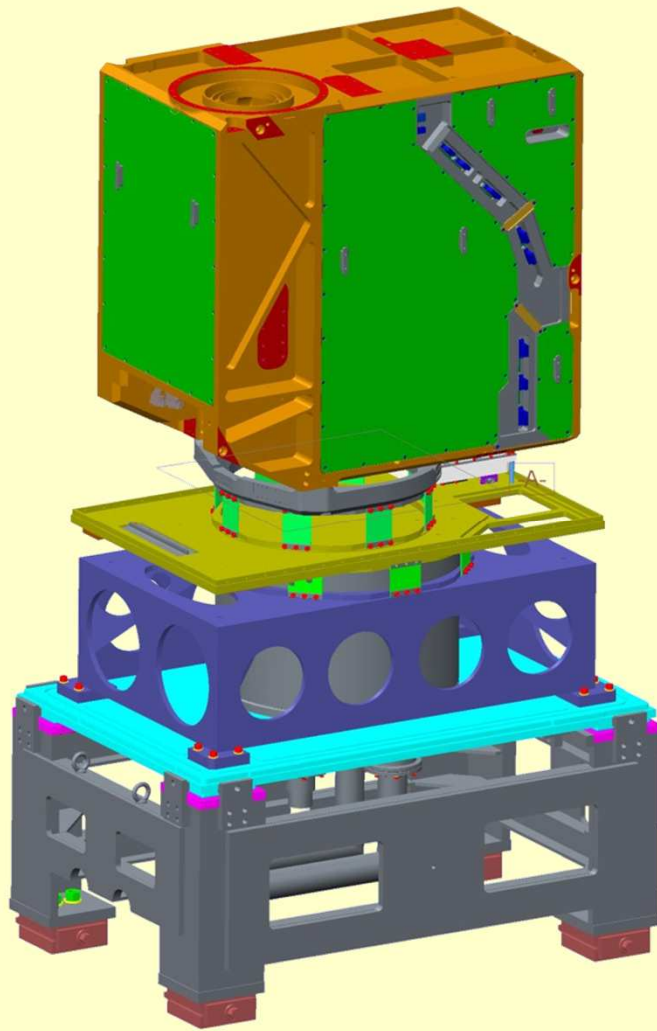
MATISSE vacuum vessel (290 K)



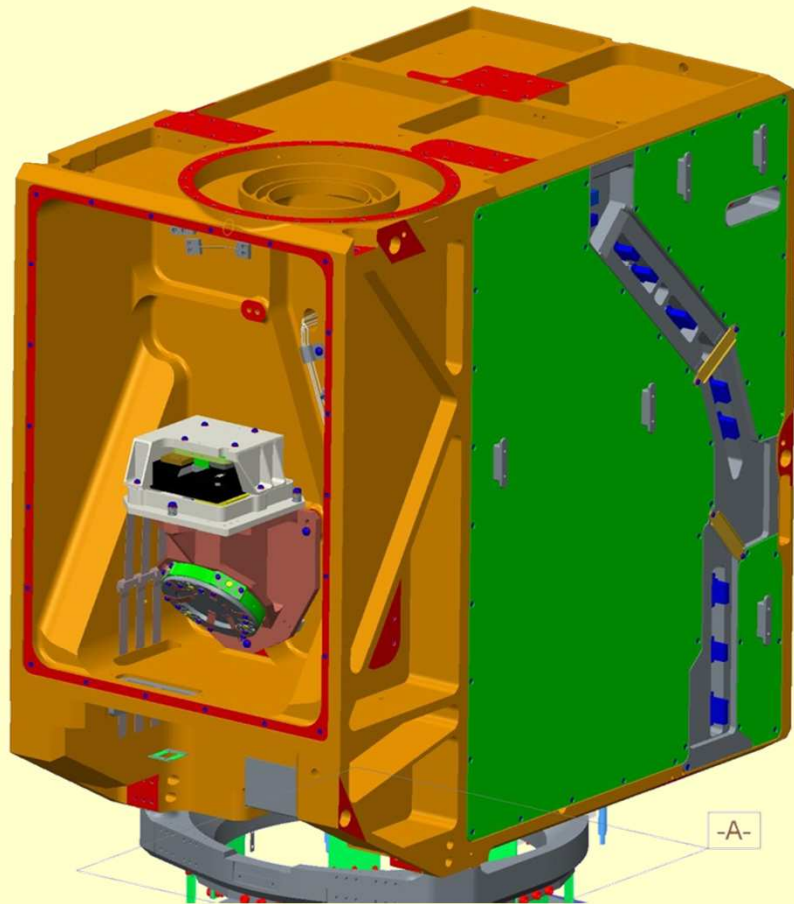
MATISSE radiation shield (100 K)



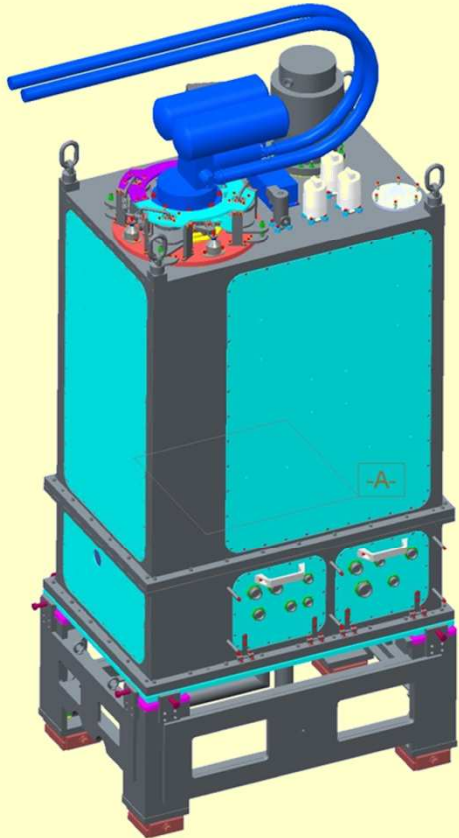
MATISSE cold optical bench (40 K)



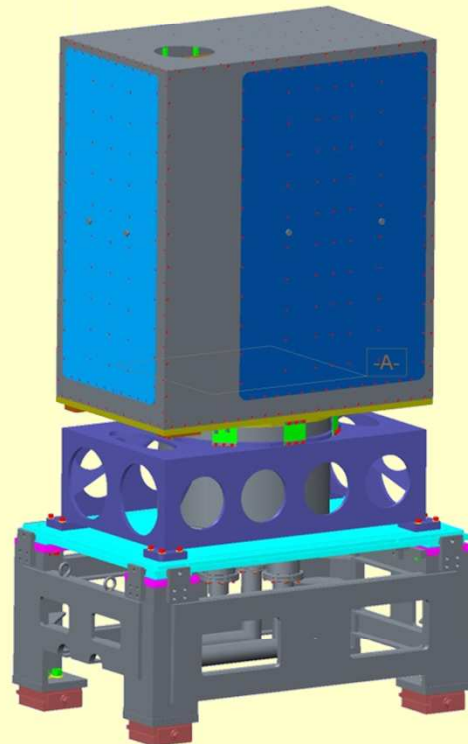
MATISSE detectors (Aquarius 8 K)



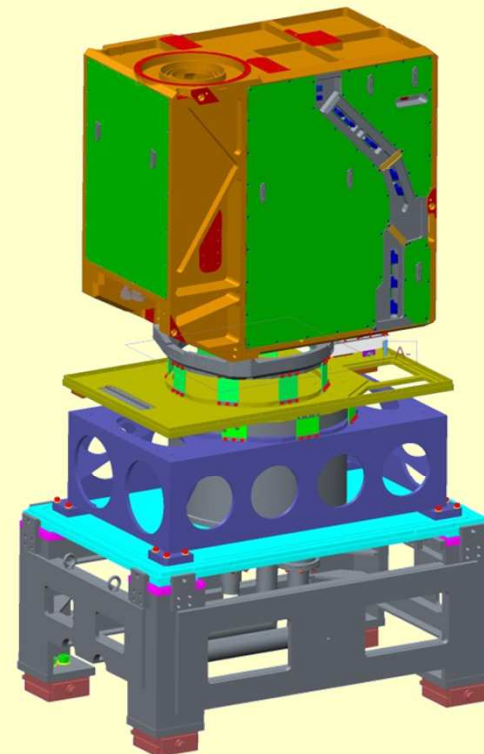
MATISSE cryostats



vacuum vessel
290 K

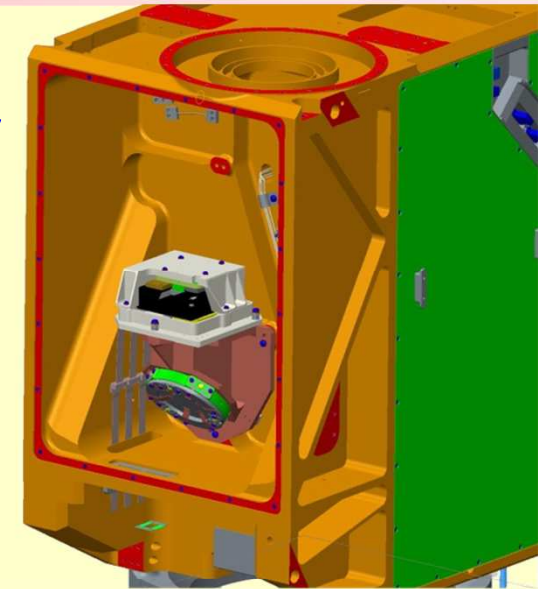


radiation shield
100 K



cold optics
40 K

detector
8 K



PT cooler

- Mounting and coupling of the PT cooler
 - Very soft mechanical connection due to vibrations
 - Damping system

Cryomech standard damping



Modified ESO damping system



Vibration damping system

- Damping system
 - Standard damping system from cryomech
 - Bellow with spring damping

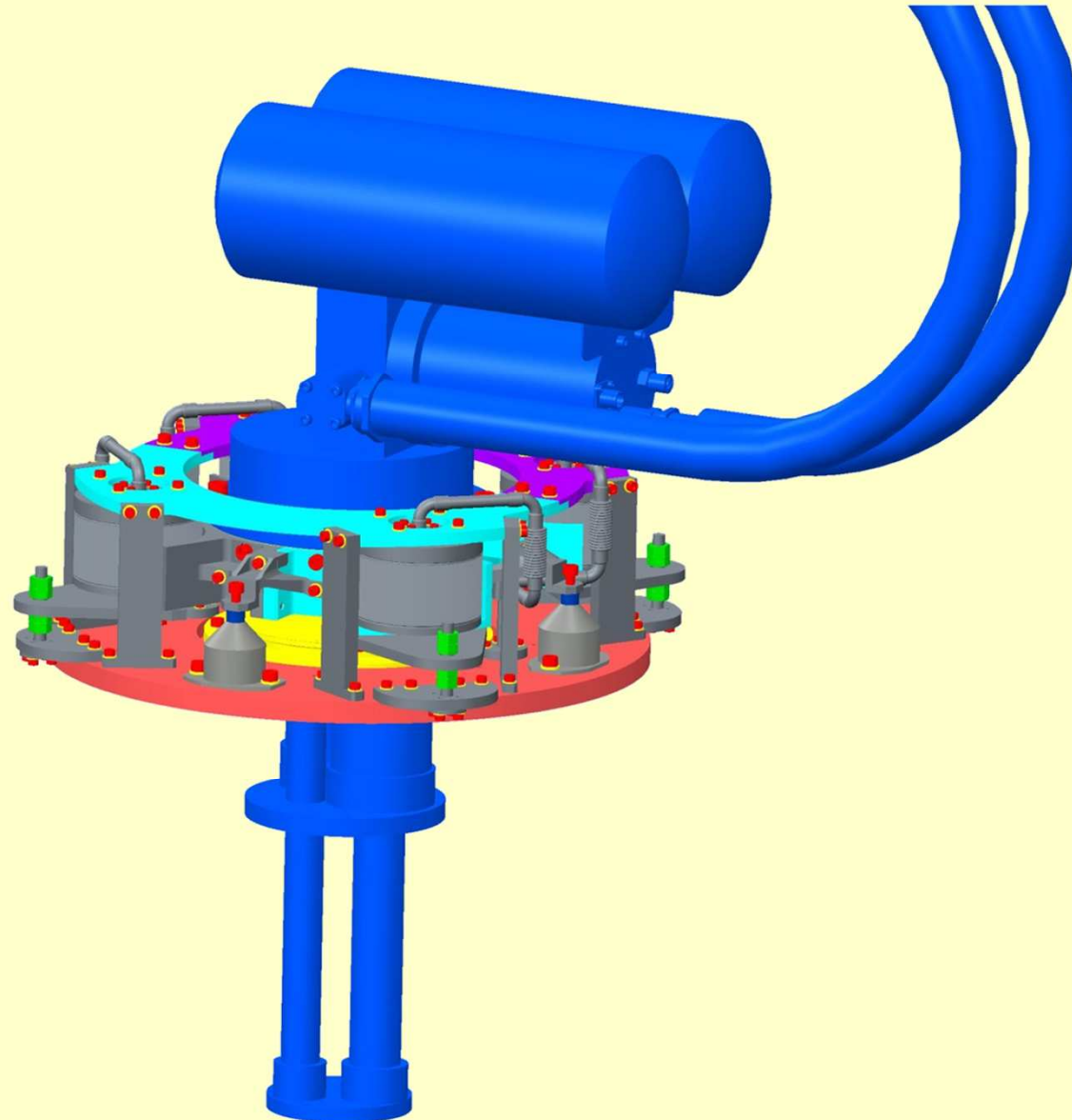


Vibration damping system

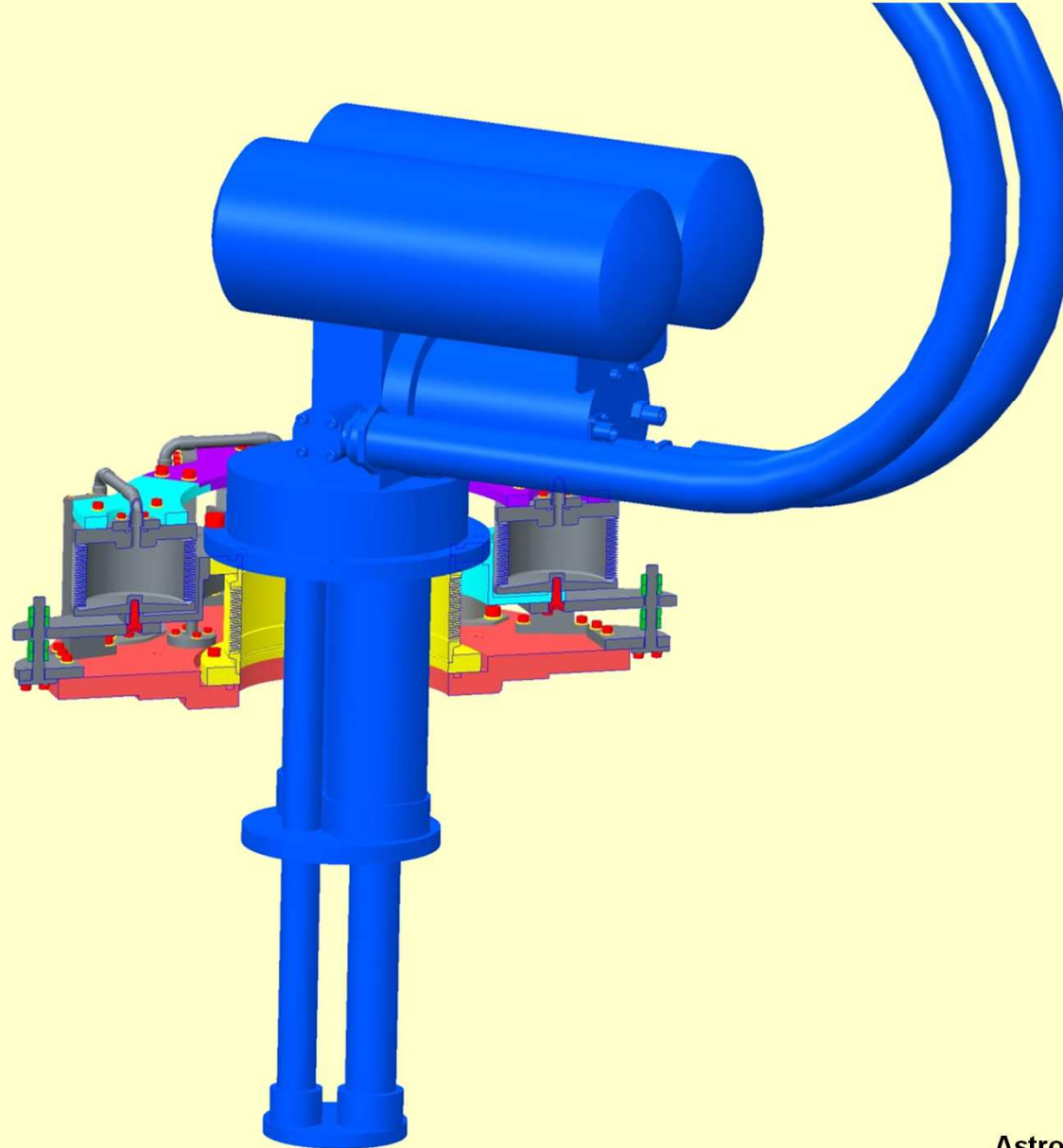
- Damping system
 - Modified ESO damping system
 - Balance between 1 central bellow pulling down and 4 small bellows pulling up



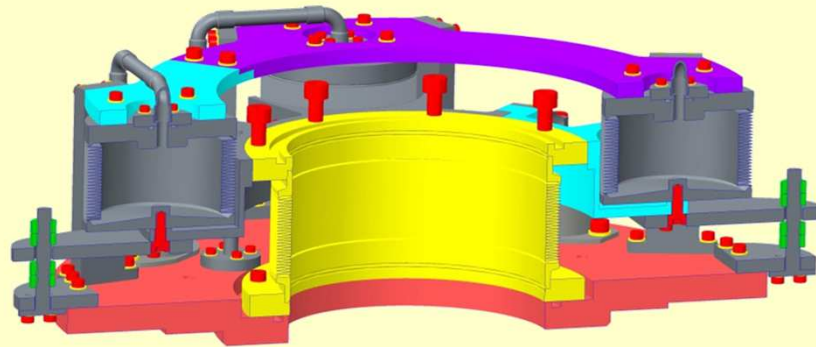
Vibration damping system



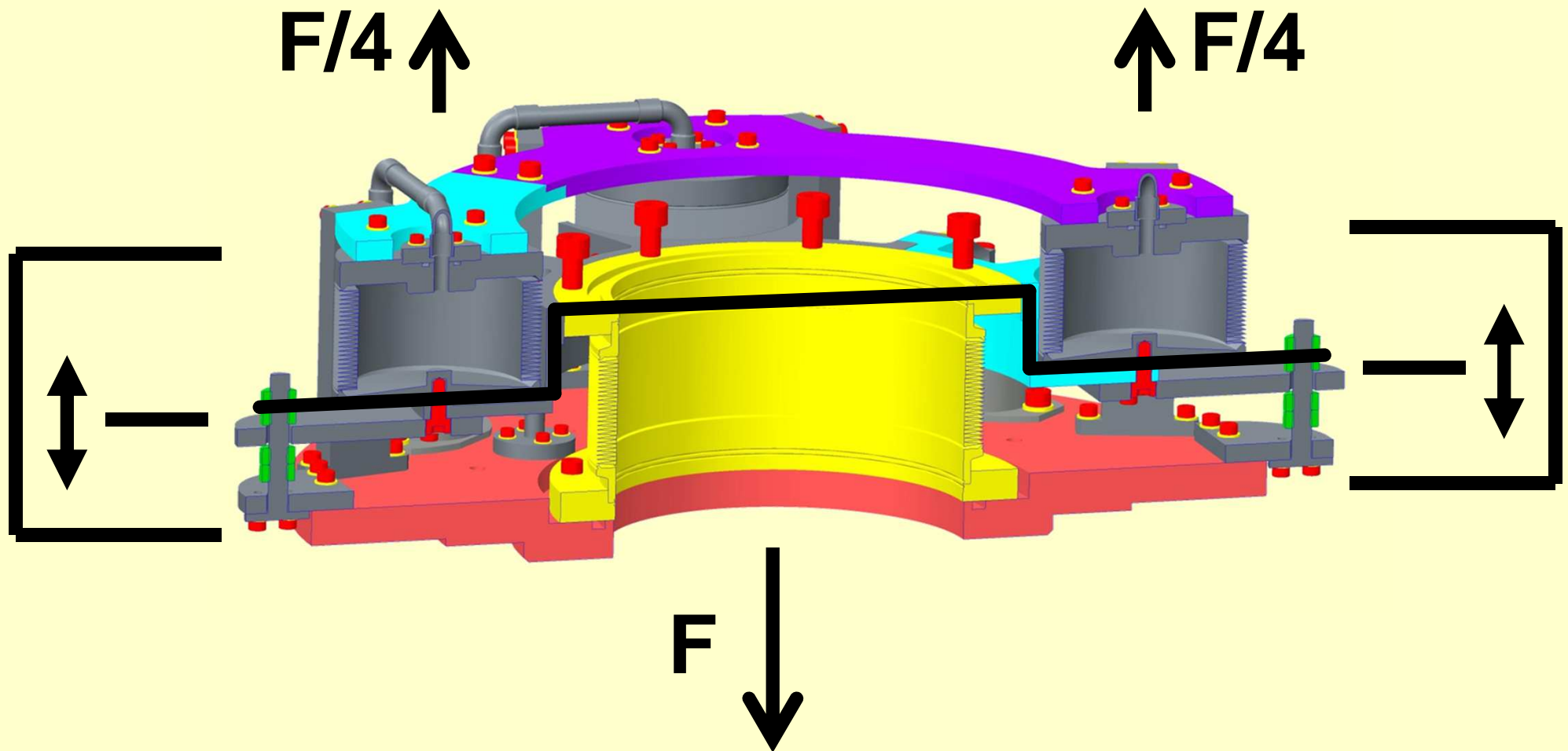
Vibration damping system



Vibration damping system



Vibration damping system

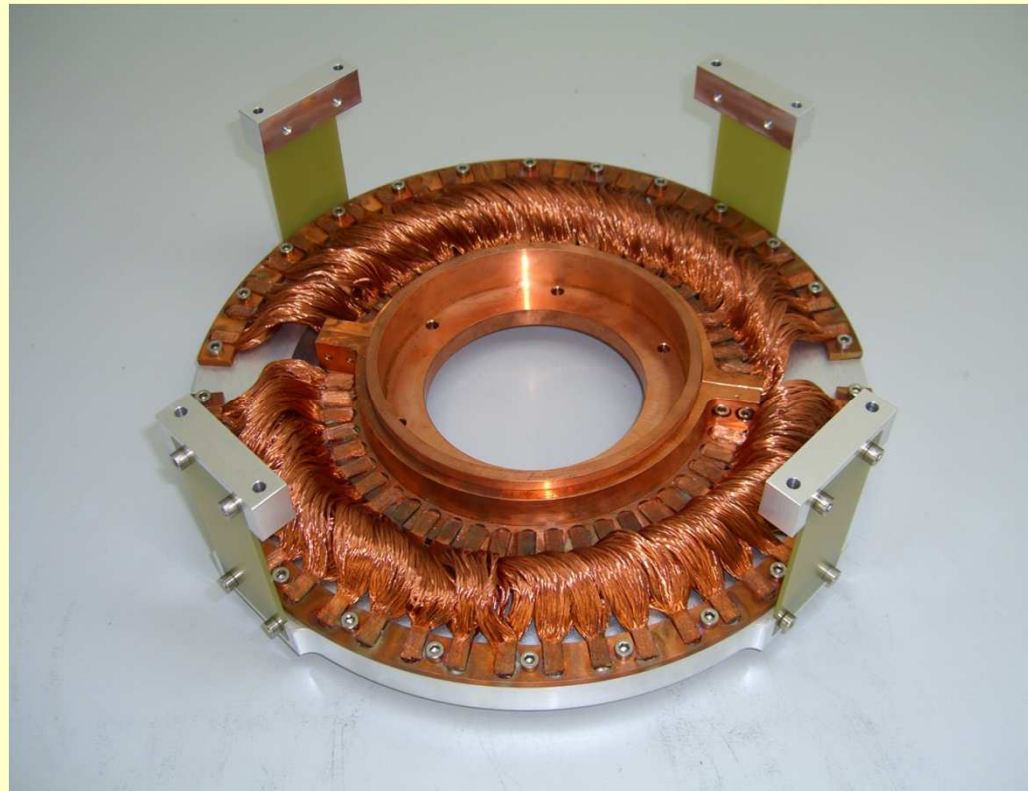


PT cooler

- Mounting and coupling of the PT cooler
 - Very soft mechanical connection
 - Damping system
 - Very soft but efficient thermal connection

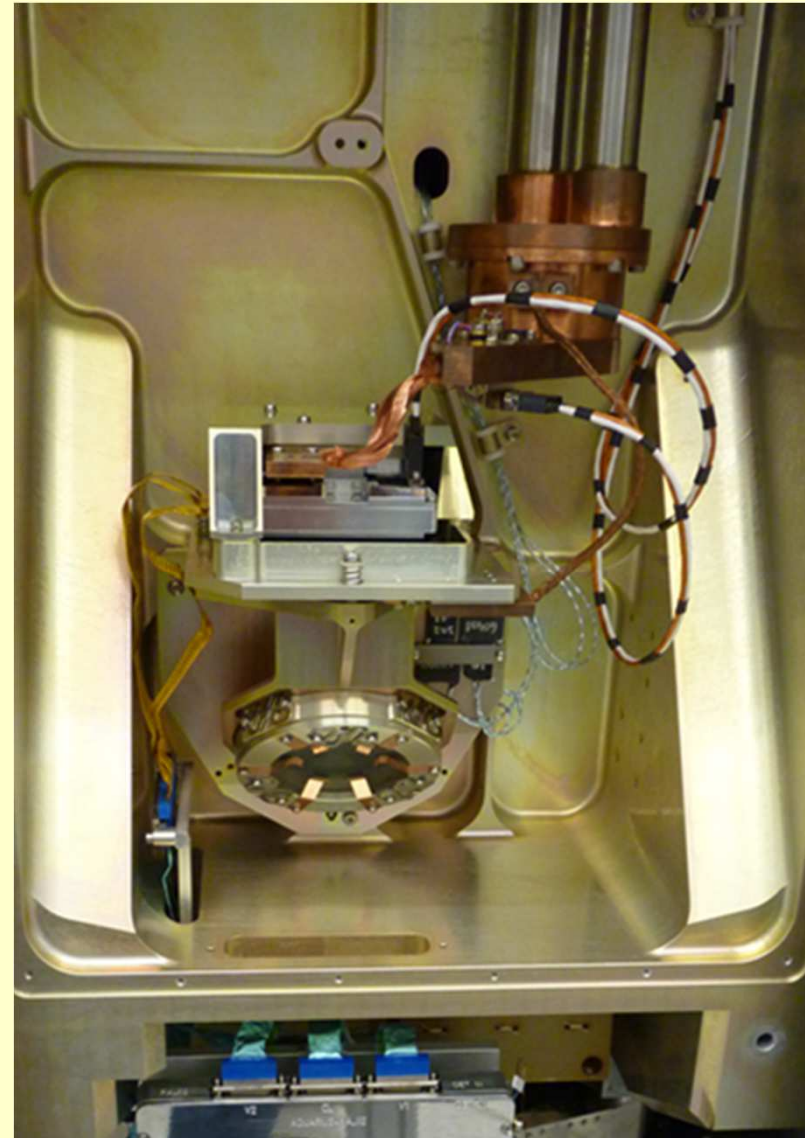
Thermal connection COB

- Very soft thermal connection
 - With copper braids $A/l \approx 20 \text{ mm}^2/\text{mm}$ for small ΔT
 - Solved by using 48 braids each $A= 20 \text{ mm}^2$, $l= 50 \text{ mm}$
 - 48 braids placed in a circle around the 1st stage of the cooler



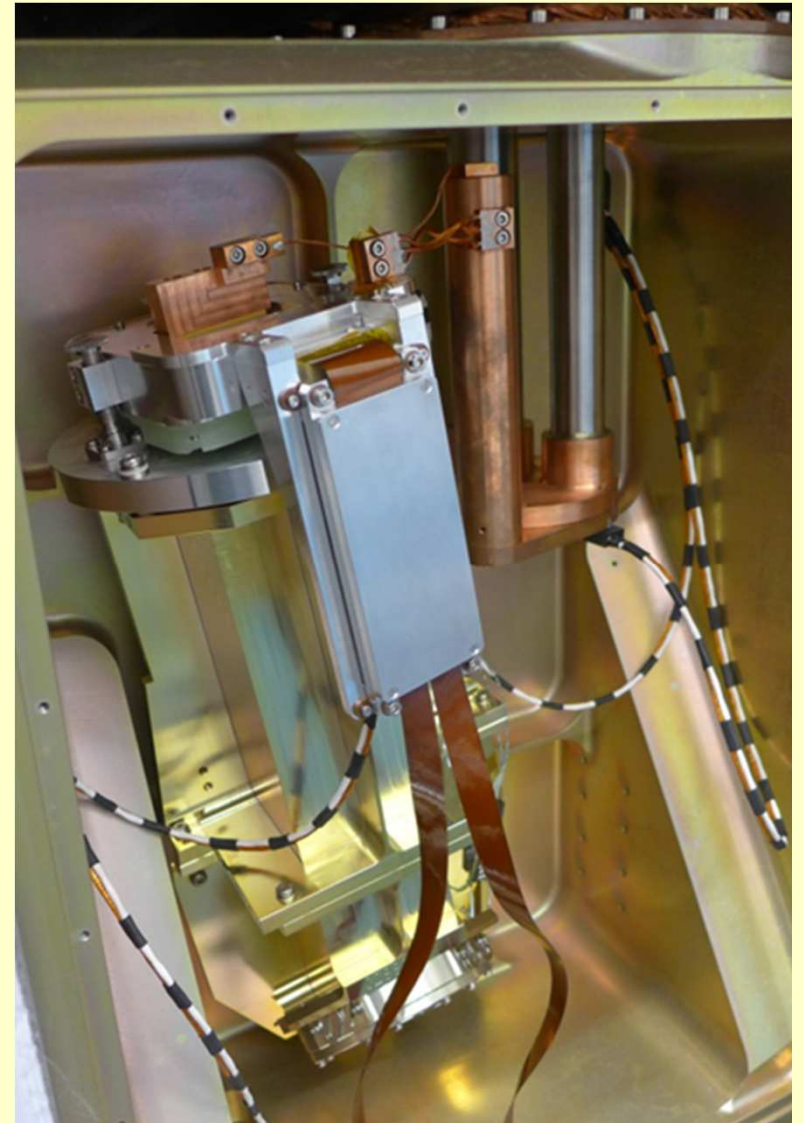
Thermal connection Aquarius

- 2nd stage to 8 K
- $\Delta T \approx 4$ K
- Copper braid
 $A = 16 \text{ mm}^2$, $l = 65 \text{ mm}$



Thermal connection Hawaii

- 2nd stage to 40 K
- $\Delta T \approx 36$ K
- Copper braid (detector)
 $A = 0.5 \text{ mm}^2$, $l = 85 \text{ mm}$
- Copper braid (preamp)
 $A = 2.5 \text{ mm}^2$, $l = 50 \text{ mm}$



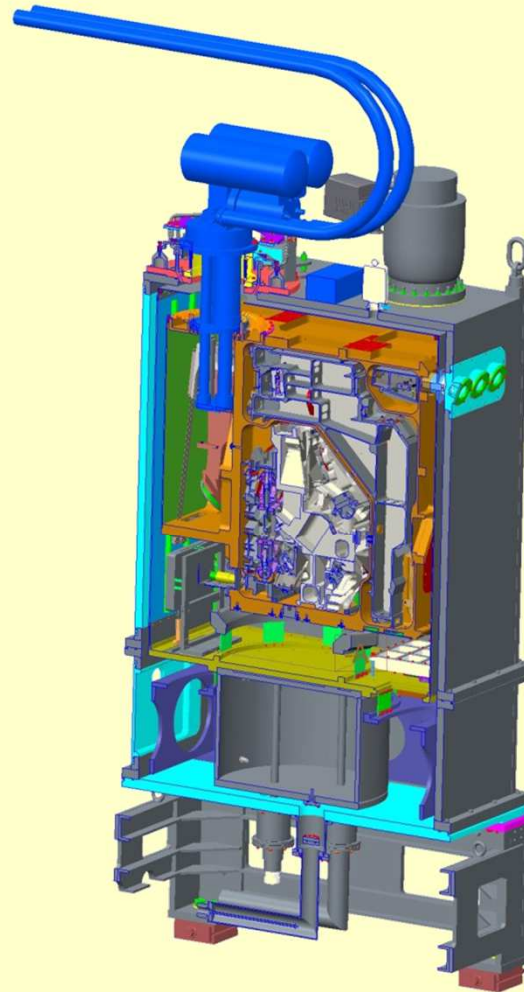
MATISSE cryostats

- Mounting and coupling of the PT cooler



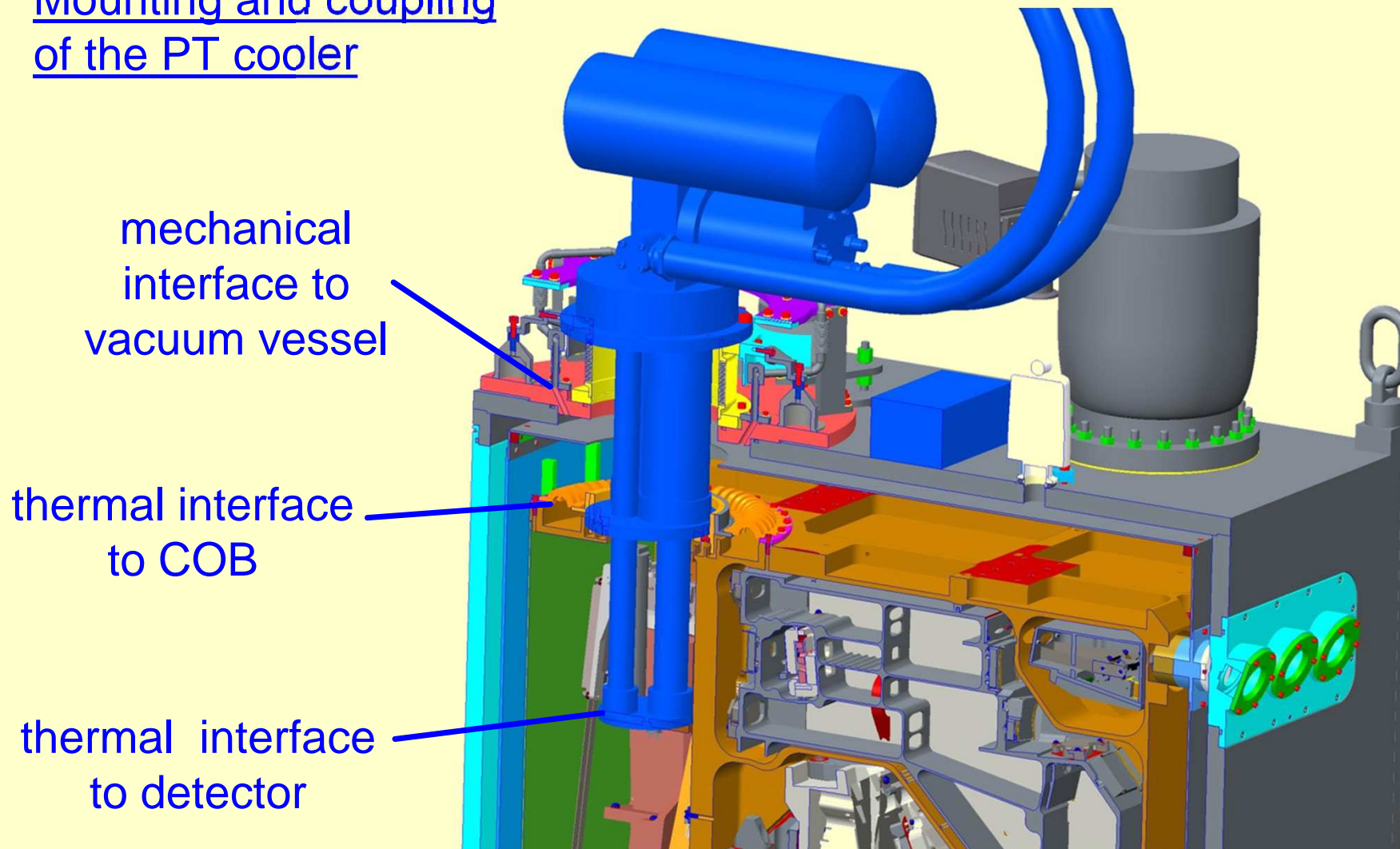
MATISSE cryostats

- Mounting and coupling of the PT cooler



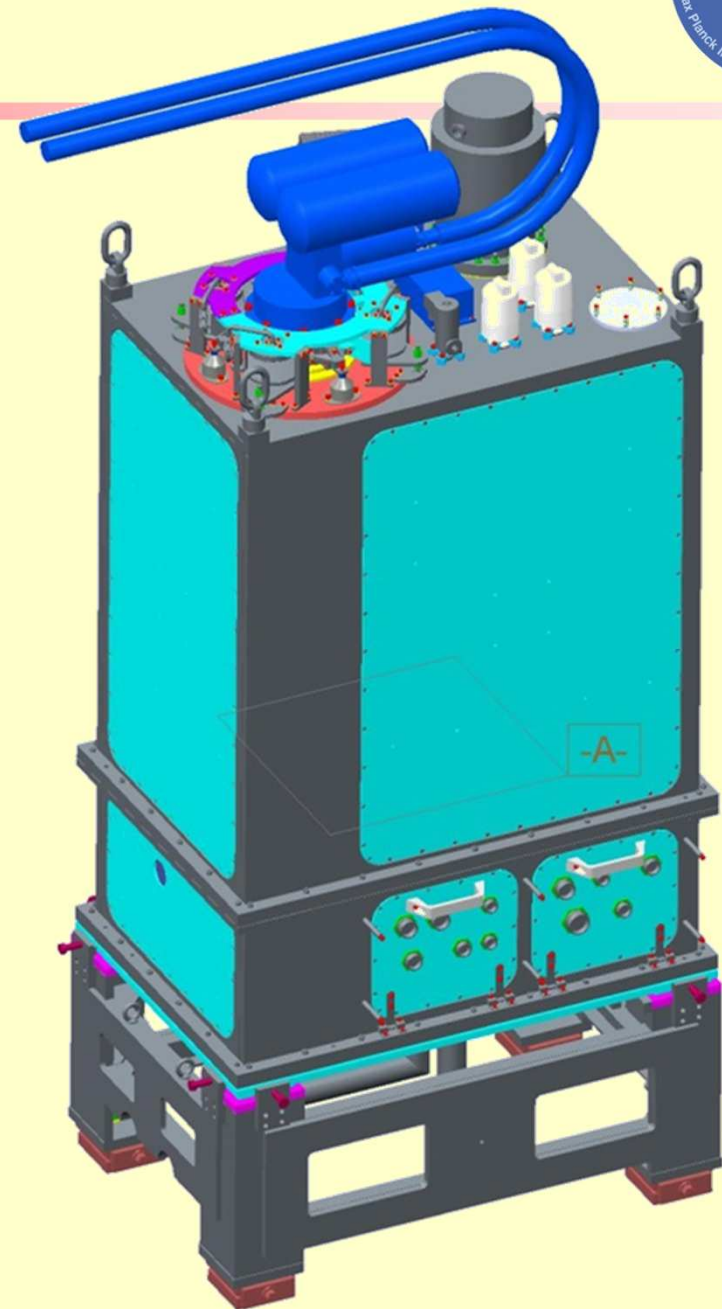
MATISSE cryostats

- Mounting and coupling of the PT cooler



MATISSE cryostats

- Accessibility to optics and detector
 - Side window for access to cold optics
 - Backside window for access to detector



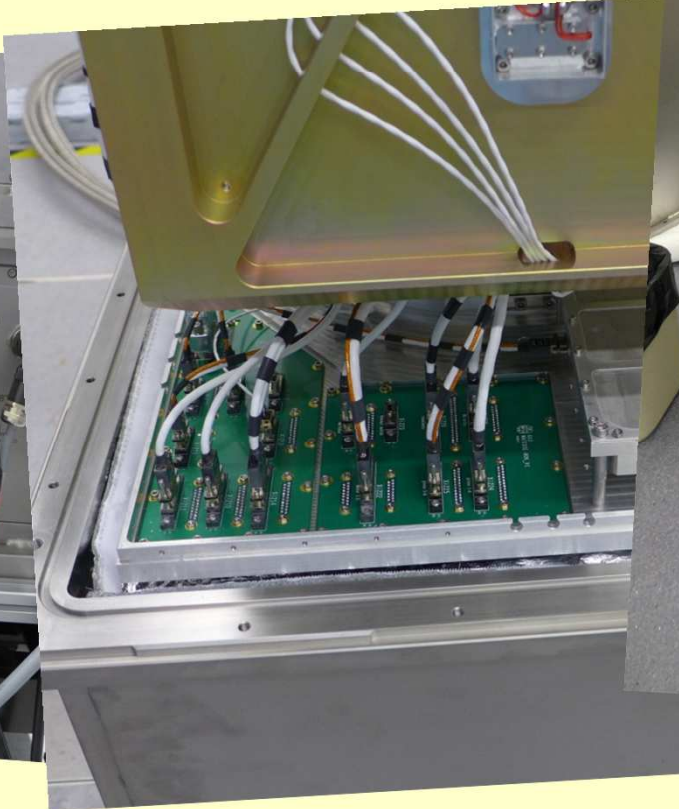
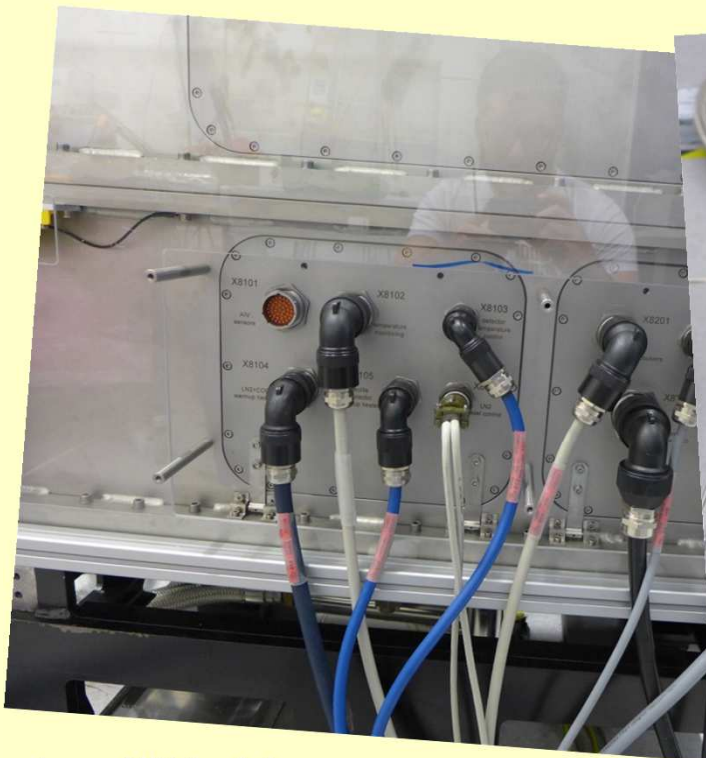
MATISSE cryostats

- Access
- Side
- Back



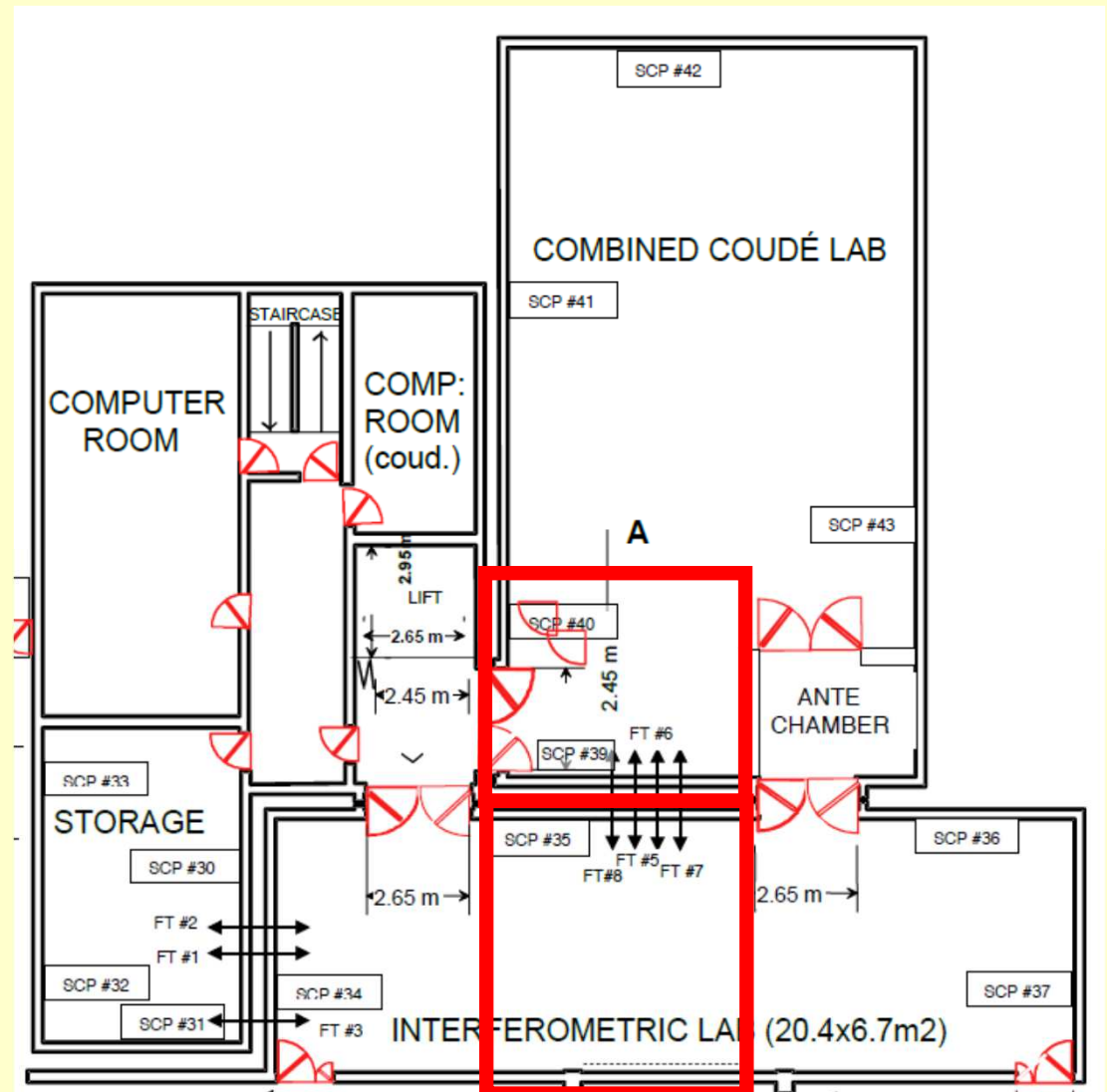
Cabling

- Cabling for the ICE
 - All feedthroughs on the lower part of the vacuum
 - Thermal coupling on the radiation shield
 - Distribution panel mounted to the base plate of
 - Connectors on the side wall of the COB

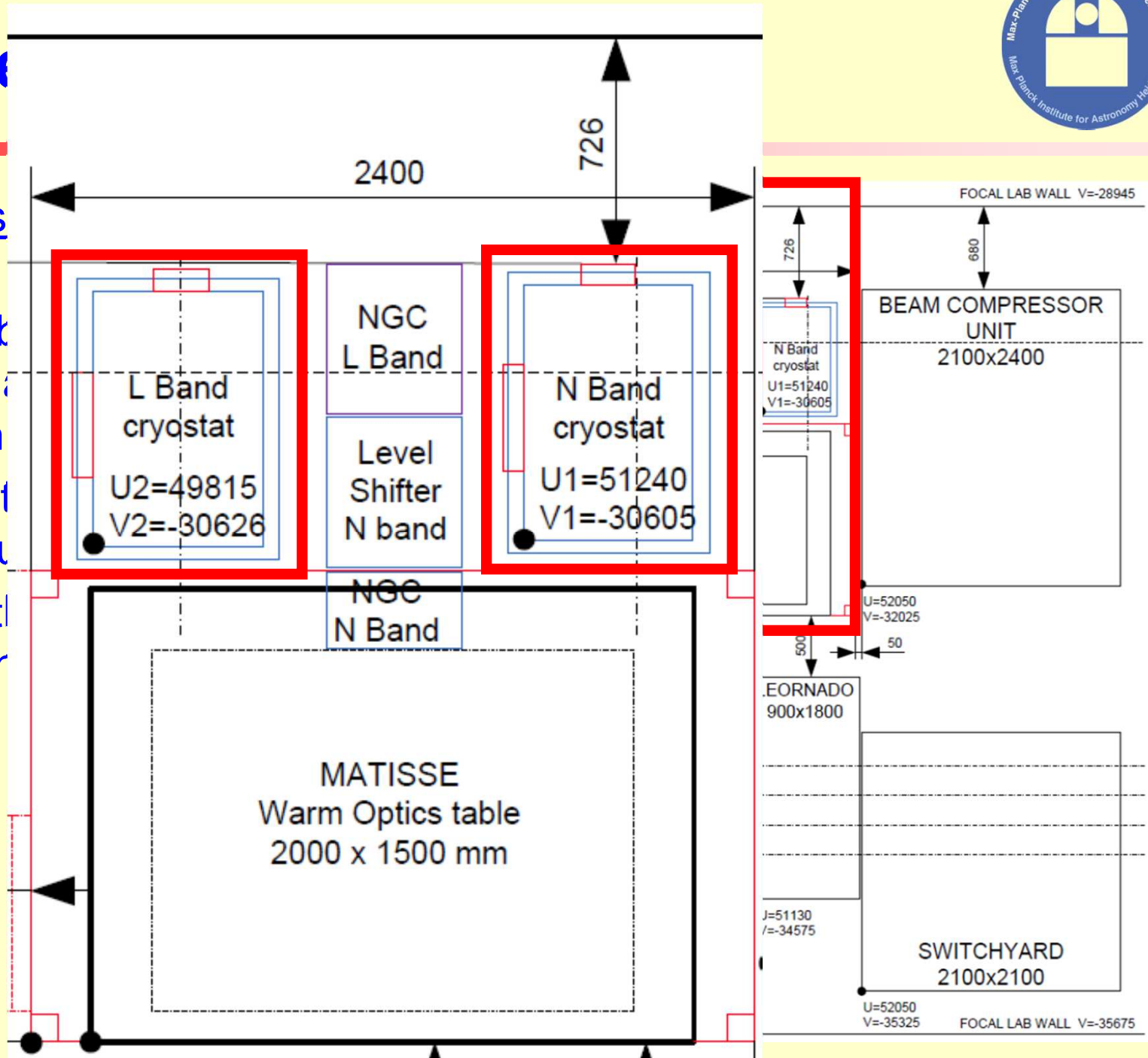


Space requirements

- Space requirements
 - Warm optical bench and cryostats inside interferometric lab
 - Electronic cabinets and helium-compressors inside combined coudé lab



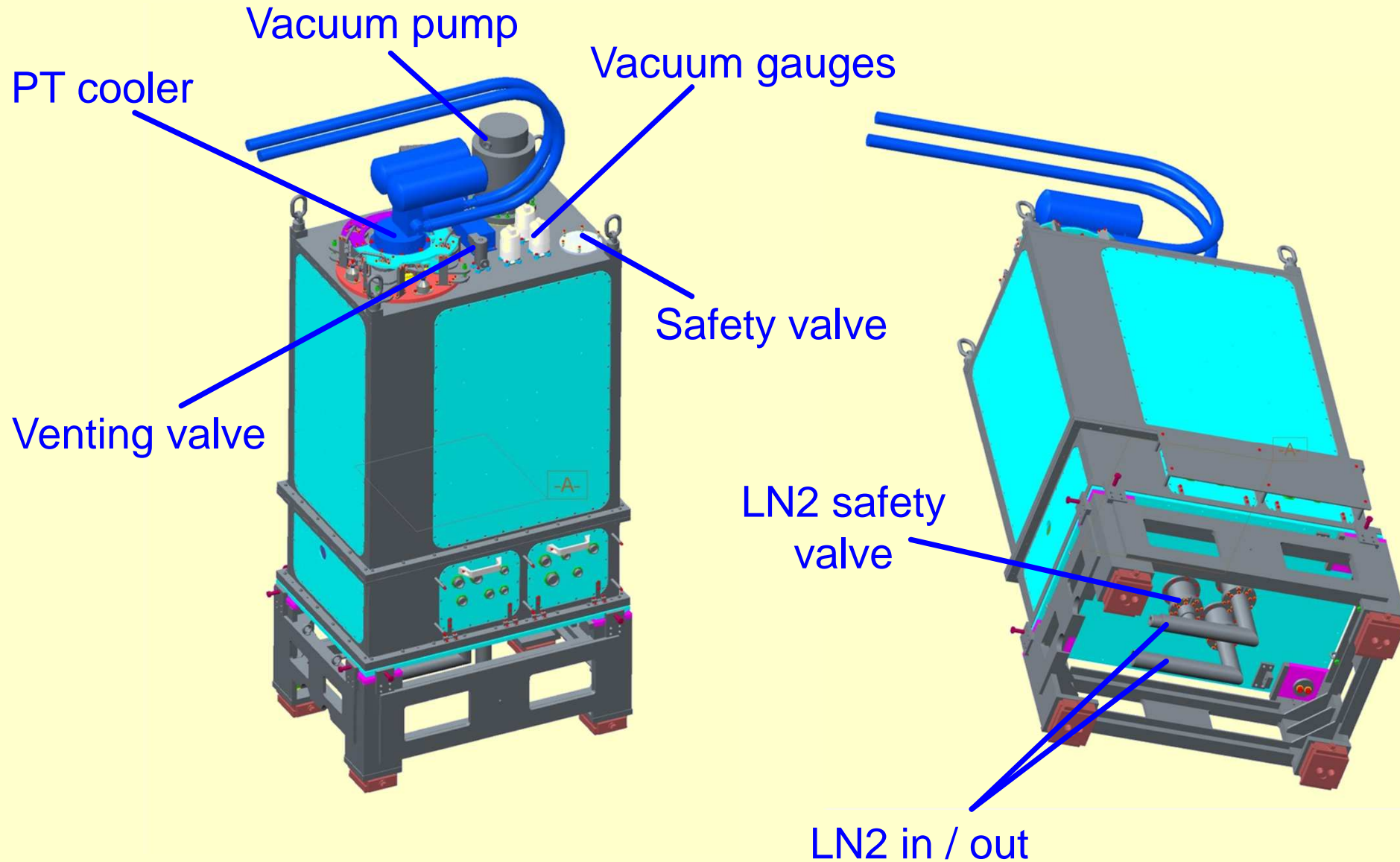
- Space requirements interferometric lab
- Warm optical table cryostats within a 2.65 m x 2.40 m
- Space required for access around the instrument
- Accessibility of the instrument from one side and back necessary



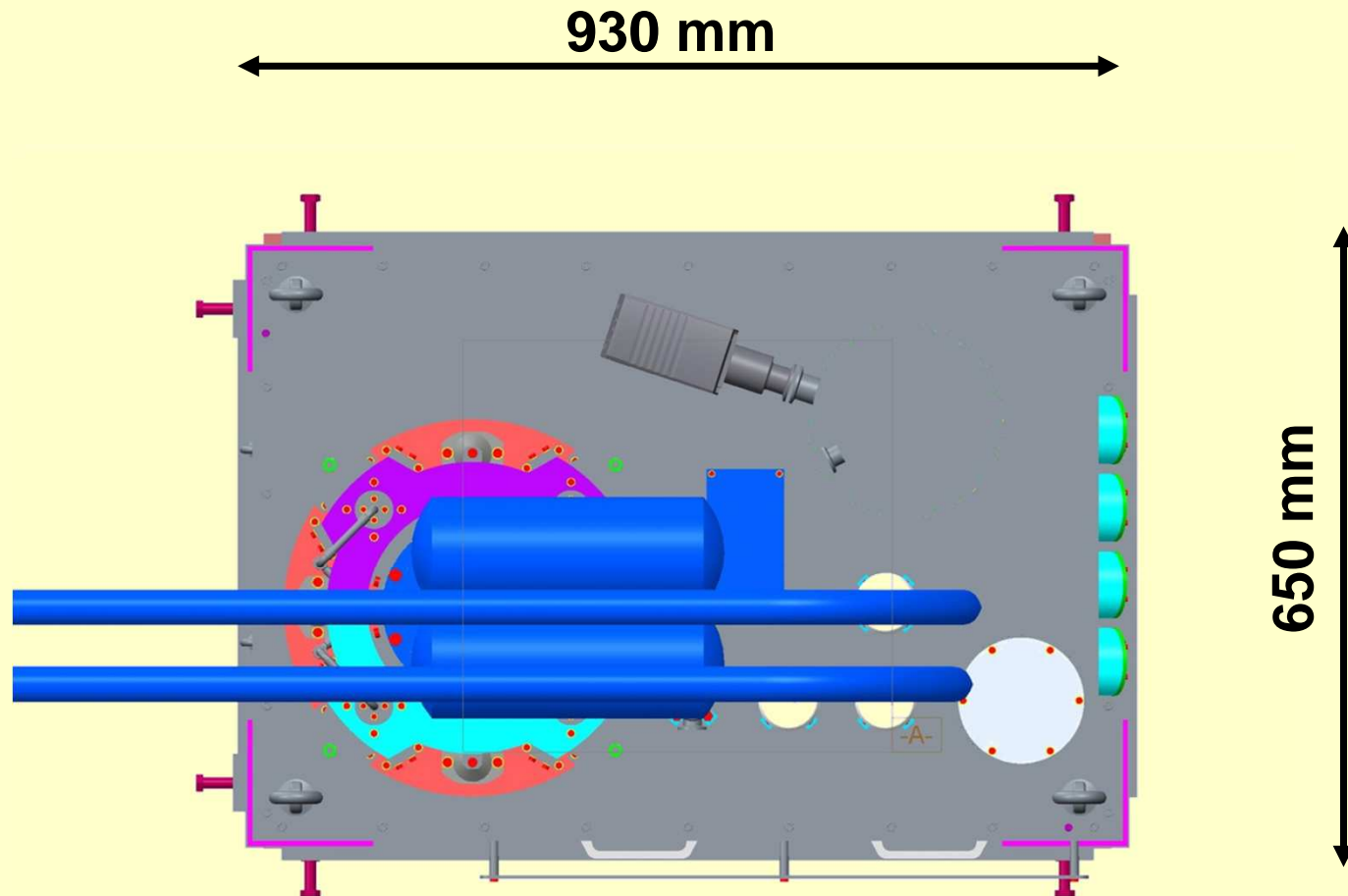
Space requirements

- Methods to reduce size
 - Extremely compact optics with multiple folding in 3 dimensions
 - Vacuum vessel with flat side walls
 - All protruding installation on top or below the cryostat
 - Helium lines bend in a high bow above the pass way on the back

Space requirements

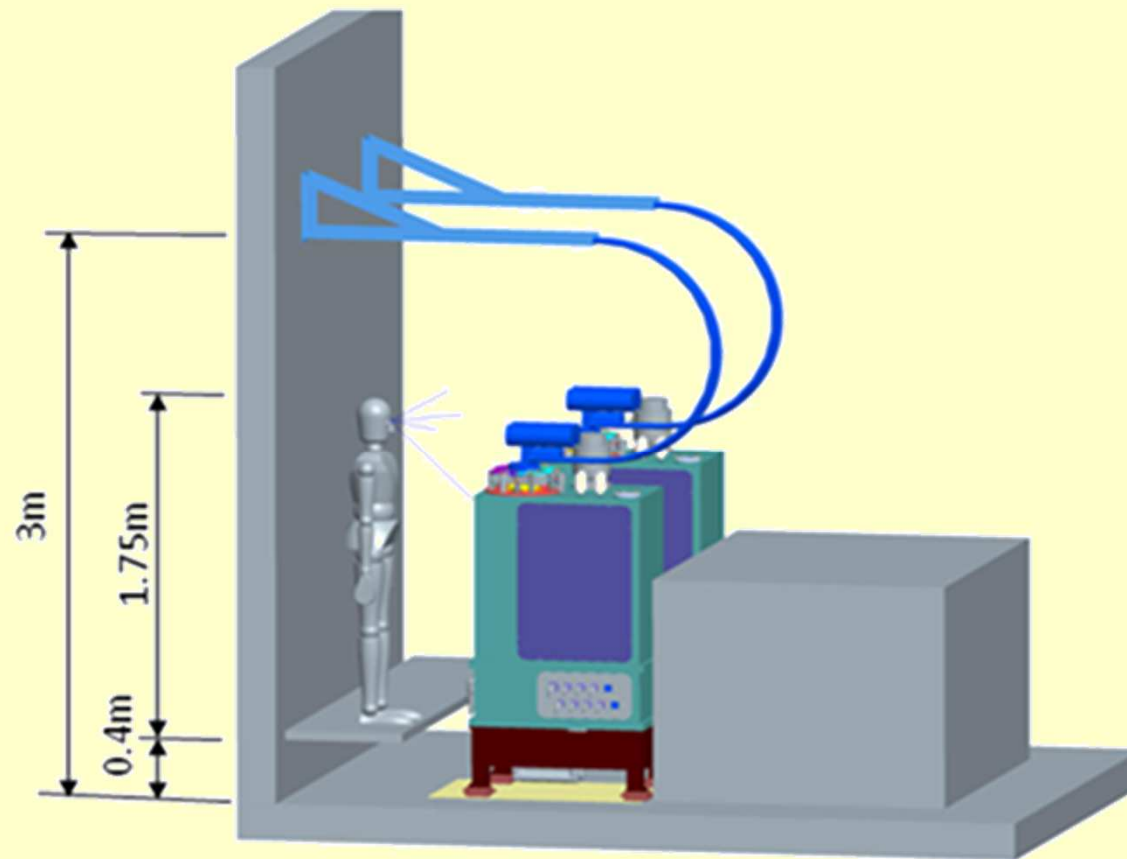


Space requirements



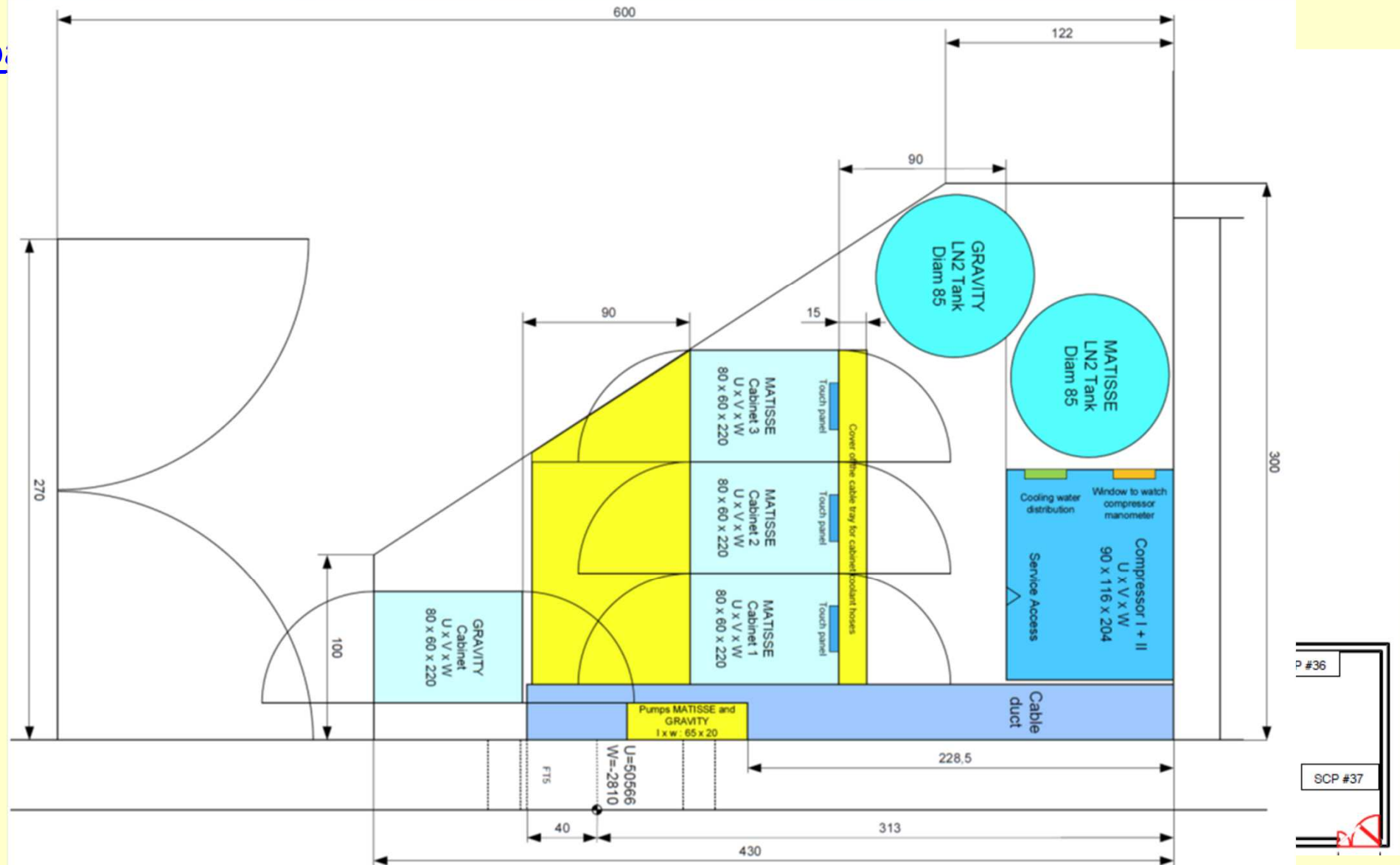
Space requirements

- Pass way behind the cryostat
- Helium lines bend up



Space requirements

- Sp
-
-



Space requirements

- Space inside combined coudé lab
 - Two compressors mounted in vibration and noise damping housing
 - Mounted on top of each other

Compressor housing



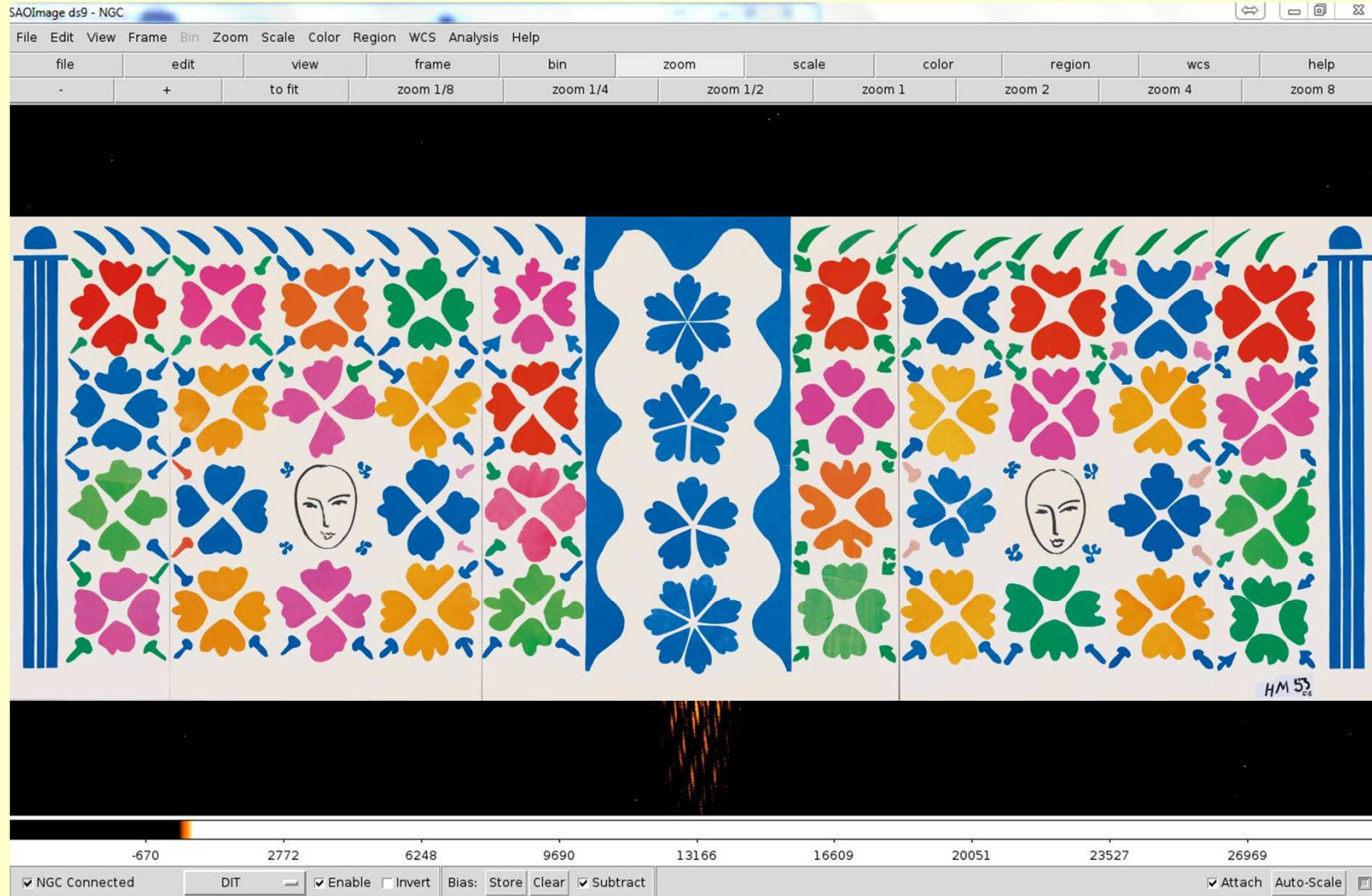
MATISSE in Nice

- Current status and set up
 - Both cryostats in Nice
 - Integration and alignment of warm optics almost finished
 - Soon start of tests with the complete setup



MATISSE in Nice

- First fringes in July 2015





MATISSE



Lessons learned

- Good planning is important
- Don't start manufacturing before the design is finished or it will result in technical compromise or remanufacturing of parts
- Don't work only on the solution of your own problems, think also about what your partners need
- "higher authority" is needed to keep discipline in
 - planning, testing, documentation
- Make realistic planning for resources like personnel, lab space, money...
 - Fight for getting those resources

Future

- Future planning
 - PAE, May 11, 2017
 - Start packing: May 12, 2017
 - Start transport: June 16, 2017
 - Start integration, June 30, 2017

MATISSE on Paranal 2017





Questions?



Good bye Uwe

