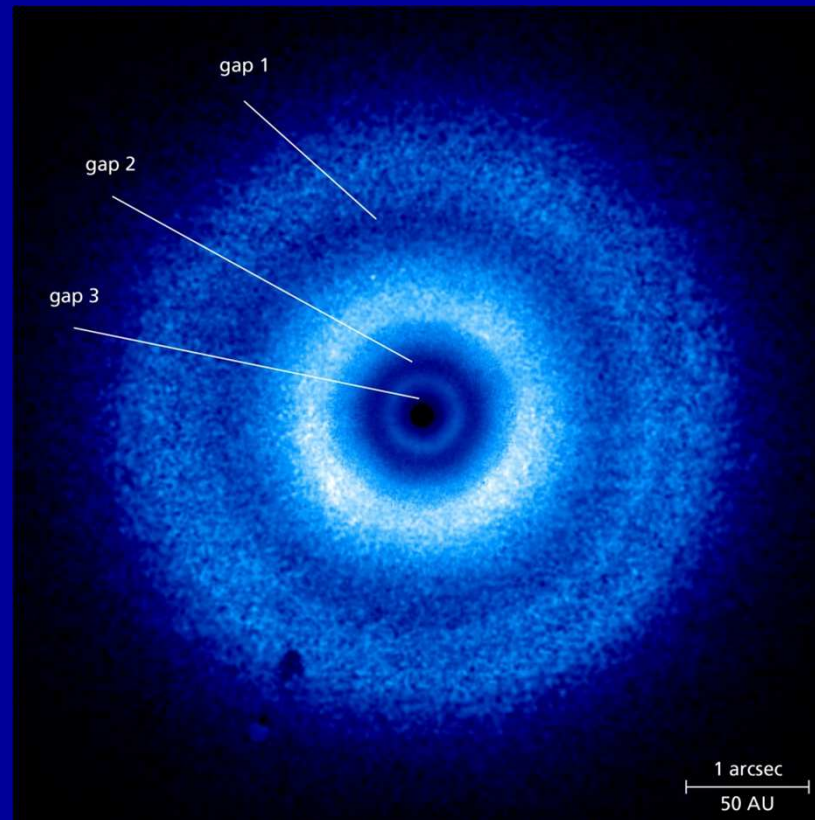


Perspektiven der Astronomie

Thomas Henning
Max-Planck-Institut für Astronomie



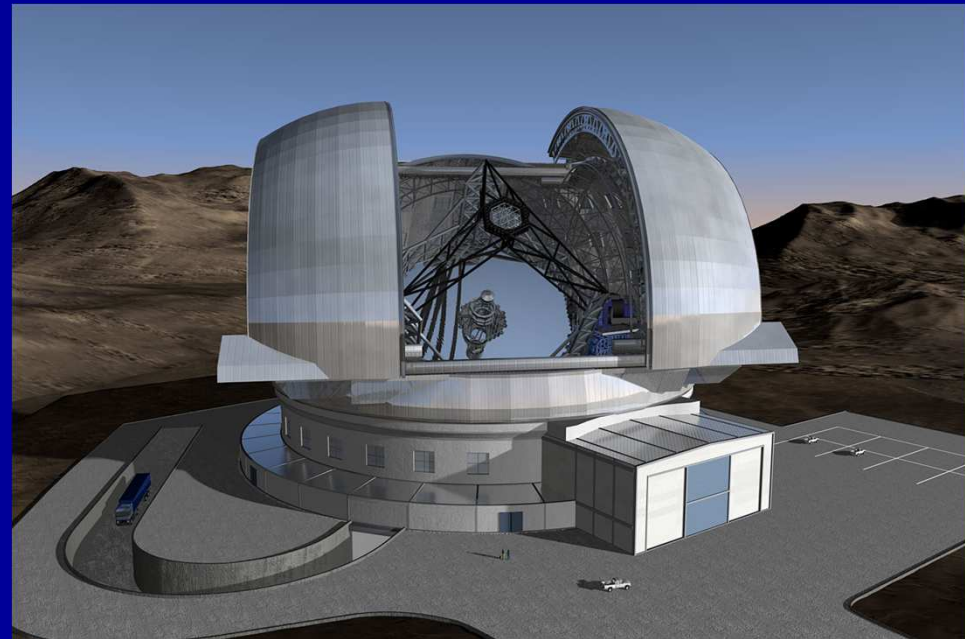
TW Hydrae Ringwelt

**SPHERE/IRDIS @ H band mit apodized Lyot Koronagraph (40 marcsec – 2.4 a.u.)
van Boekel, Henning, Menu et al. (2017)**

Landschaft in 2030

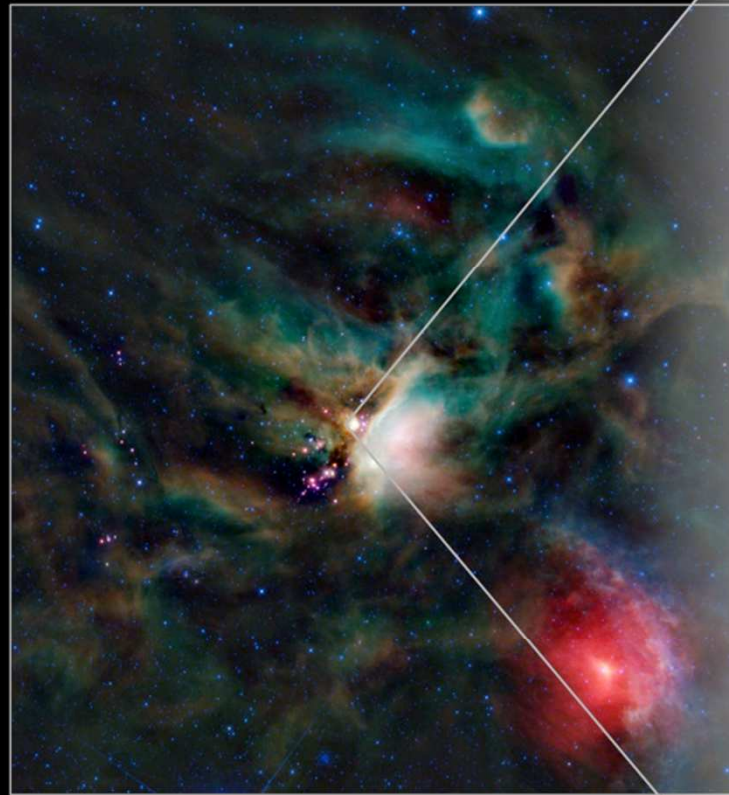


ALMA- Submm/mm
Imaging & Spektroskopie



ELTs – Vis-20 μm
AO-unterstütztes Imaging &
Spektroskopie

Von Scheibenstrukturen zur Diversität von Planeten



The Ophiuchus star-forming region
Image Credit: NASA/JPL-Caltech/WISE Team

Elias 2-27 as seen by ALMA

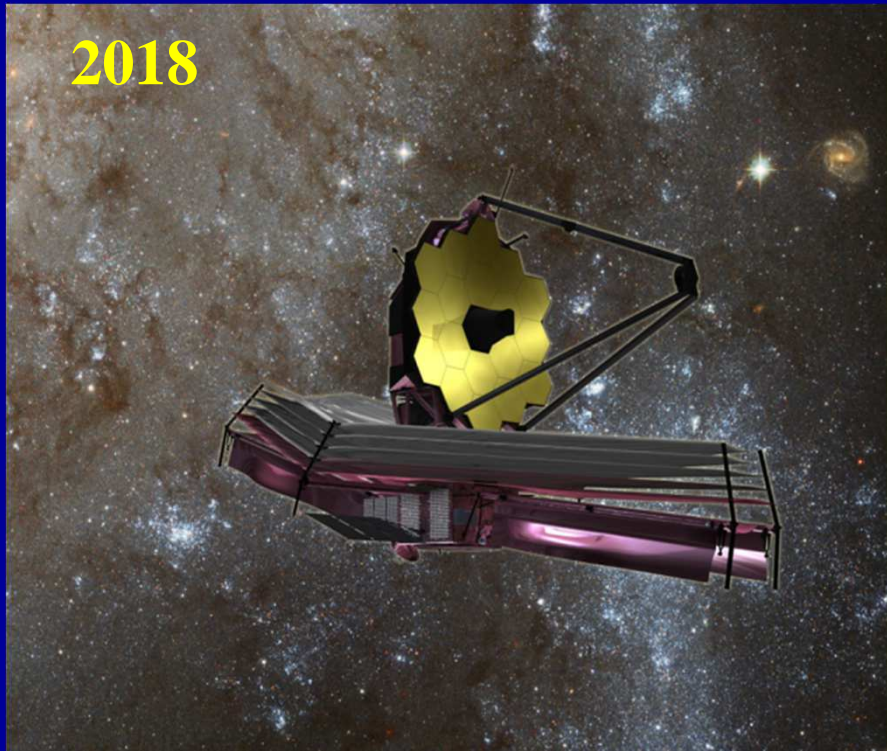
Kuiper Belt orbit



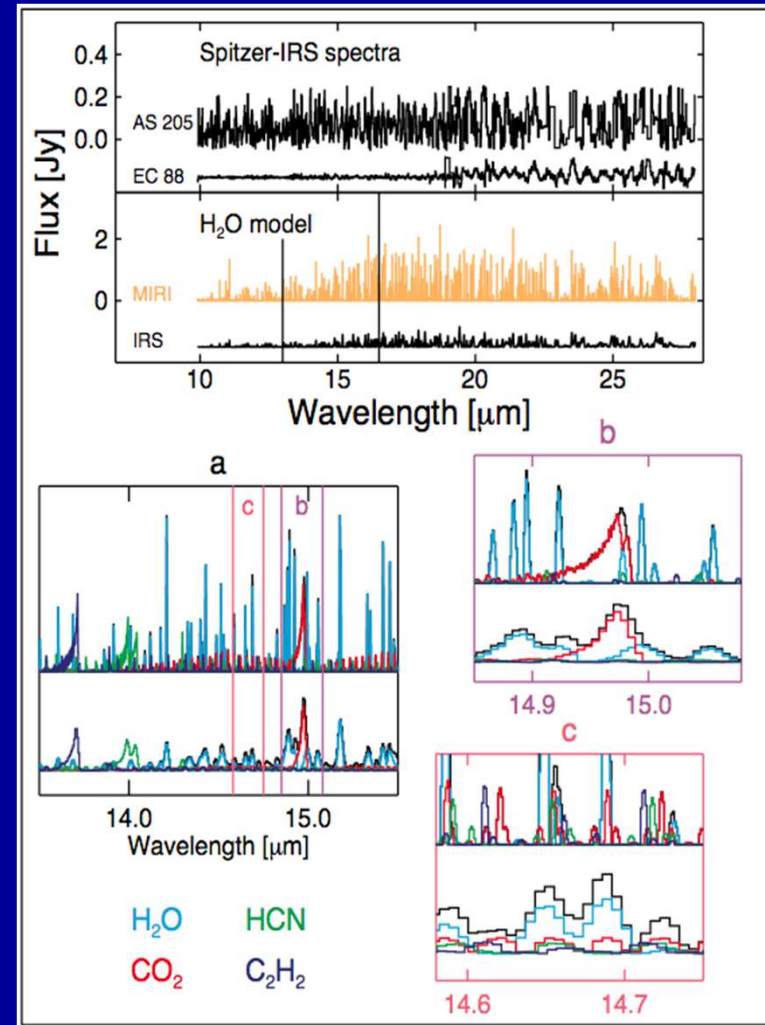
Credit: B. Saxton (NRAO/AUI/NSF);
ALMA (ESO/NAOJ/NRAO); L. Pérez (MPIfR)

Perez et al. (2016)

2030 – 10 Jahre JWST



James Webb Space Telescope
& WFIRST



Denkschrift Astronomie - 2017

- ELT
- Instrumente für La Silla/Paranal-Observatorium, ALMA
- SKA
- Weiterentwicklung von LBT, NOEMA, LOFAR

- Weltraum – Cosmic Vision Program
- Nationales Programm - SOFIA und eRosita



2020 Decadal Review in den USA ...

- LUVOIR (UV-optisches-IR-Teleskop)
- HAbEx (Habitable Exoplanet Mission)
(4-6,5 m-Teleskop – 250nm-1,8 μm für High Contrast Imaging, Starshade)
- Origins Space Telescope
(8-15 m-Teleskop, Aktiv gekühlt, 5 μm – 1mm)
- Lynx
(Röntgmission, 0.5“-Auflösung, Spektroskopie)

Was ist LUVOIR?

**Large UV / Optical / Infrared Surveyor
(LUVOIR) –
Teleskopkonzept in Tradition von Hubble**



Breite wissenschaftliche Möglichkeiten

Fernes UV bis Nahes IR

~ 8 – 16 m Spiegeldurchmesser (A – 15 m & B – 9m)

Imager & Spectrographen

Service Missions und Weiterentwicklung möglich

Gastbeobachterprogramm wie bei Hubble

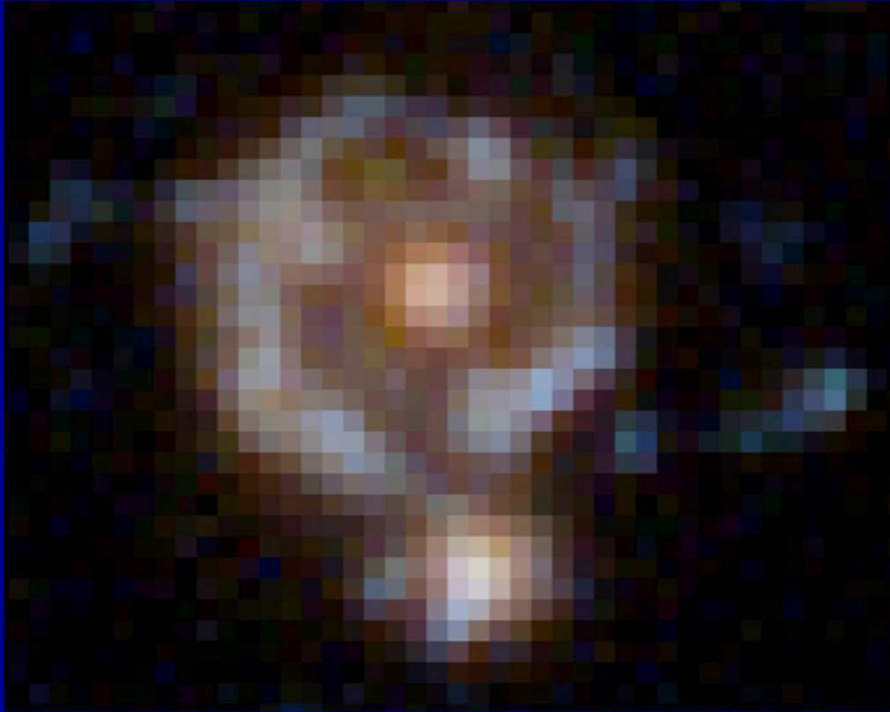
3 Hauptinstrumente

- Optischer/NIR Koronograph (bis zu Faktor 10^{-10} Unterdrückung)
- Multiobjekt UV-Spektrograph (100-400 nm)
($R=500 - 45.000$)
- Imager (2x3 arcmin) – 0,2 to 2,5 μm
- Optischer/NIR-Spektrograph
(verschiedene spektrale Auflösung bis 10^5)

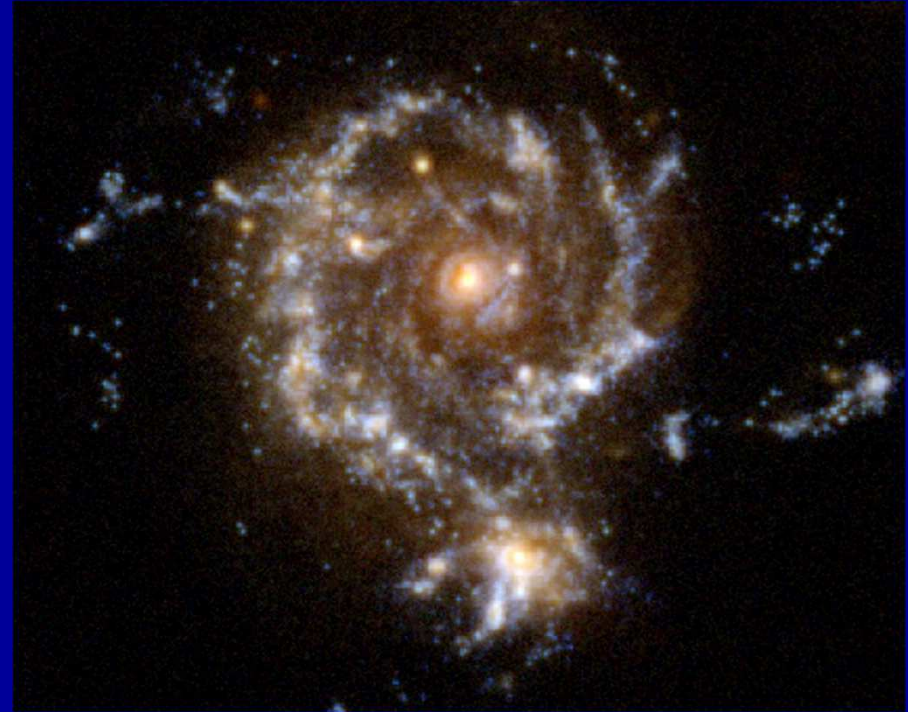


HST STIS UV instrument

Astronomie mit LUVOIR ...



Galaxy at $z = 2$
with HST



Galaxy at $z = 2$
with 12-m LUVOIR

Imagine Astronomy with LUVOIR ...

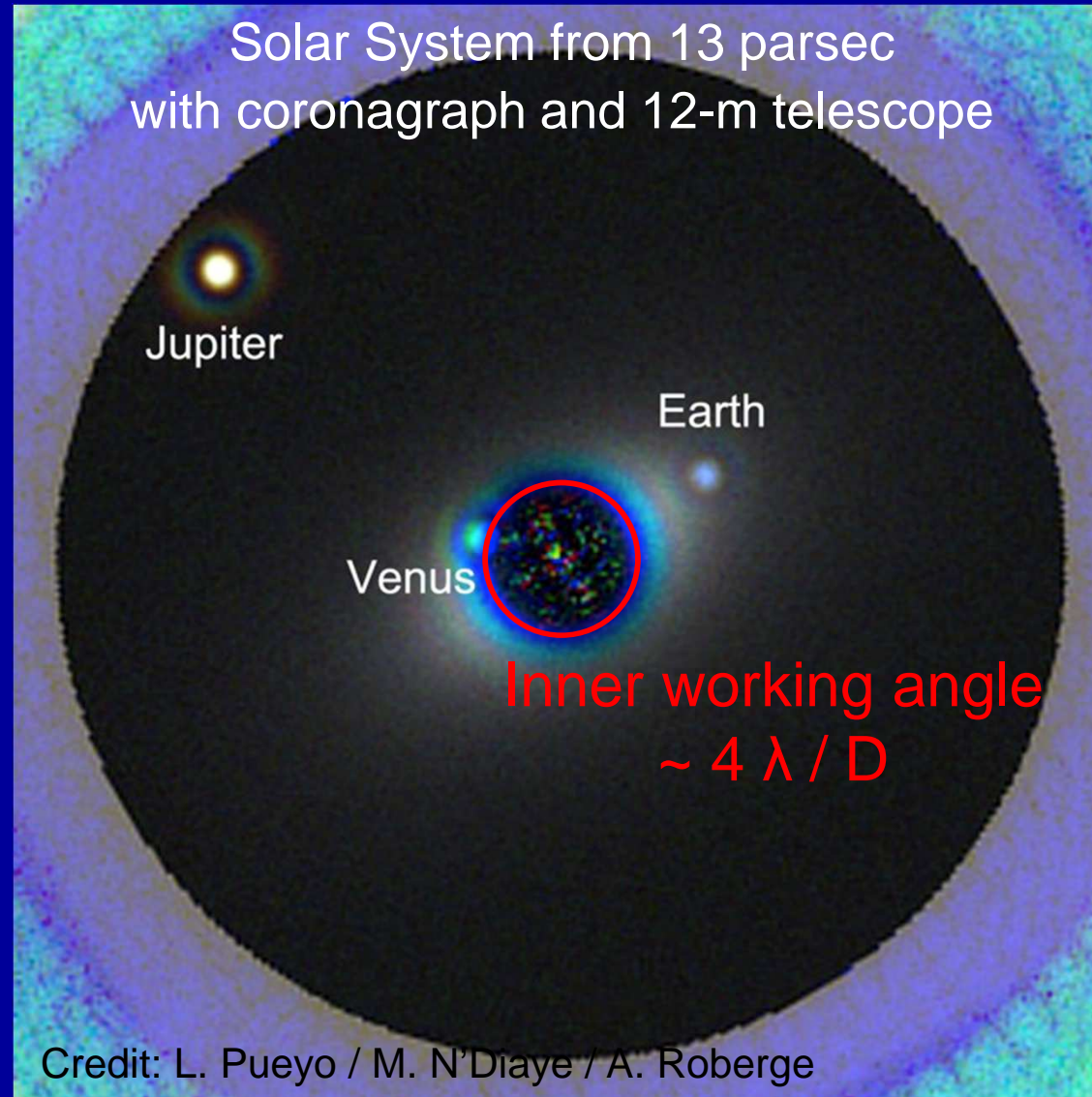


Pluto with HST



Pluto with 15-m LUVOIR

Imaging Earth 2.0



The search for life – biosignatures

Rayleigh scattering

Spectrum of Modern Earth

O₂



BLUE

RED

0.4 microns

2.4 microns

The search for life – biosignatures

Rayleigh scattering

Spectrum of Modern Earth

O₂

Water vapor



BLUE

RED

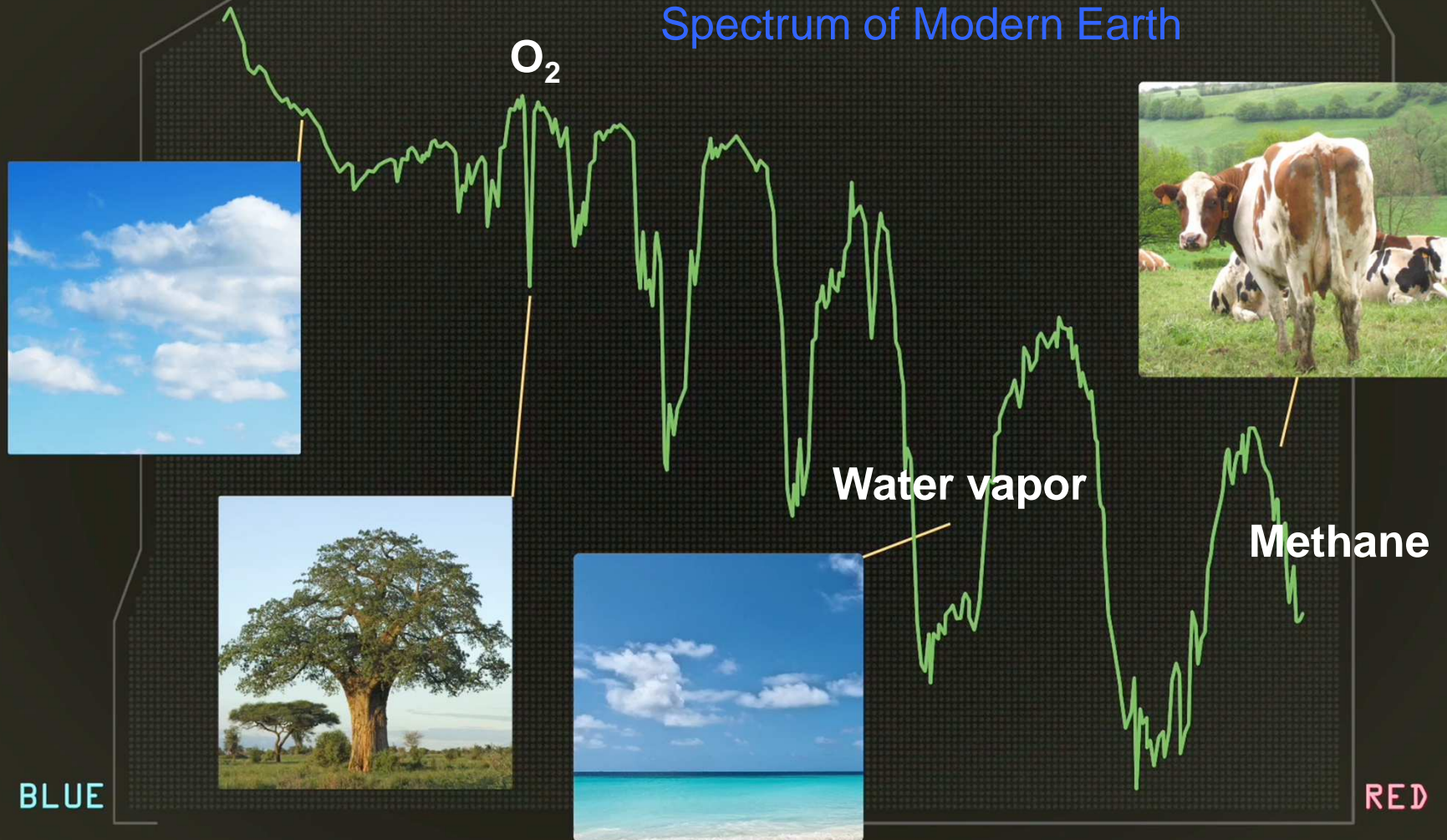
0.4 microns

2.4 microns

The search for life – biosignatures

Rayleigh scattering

Spectrum of Modern Earth



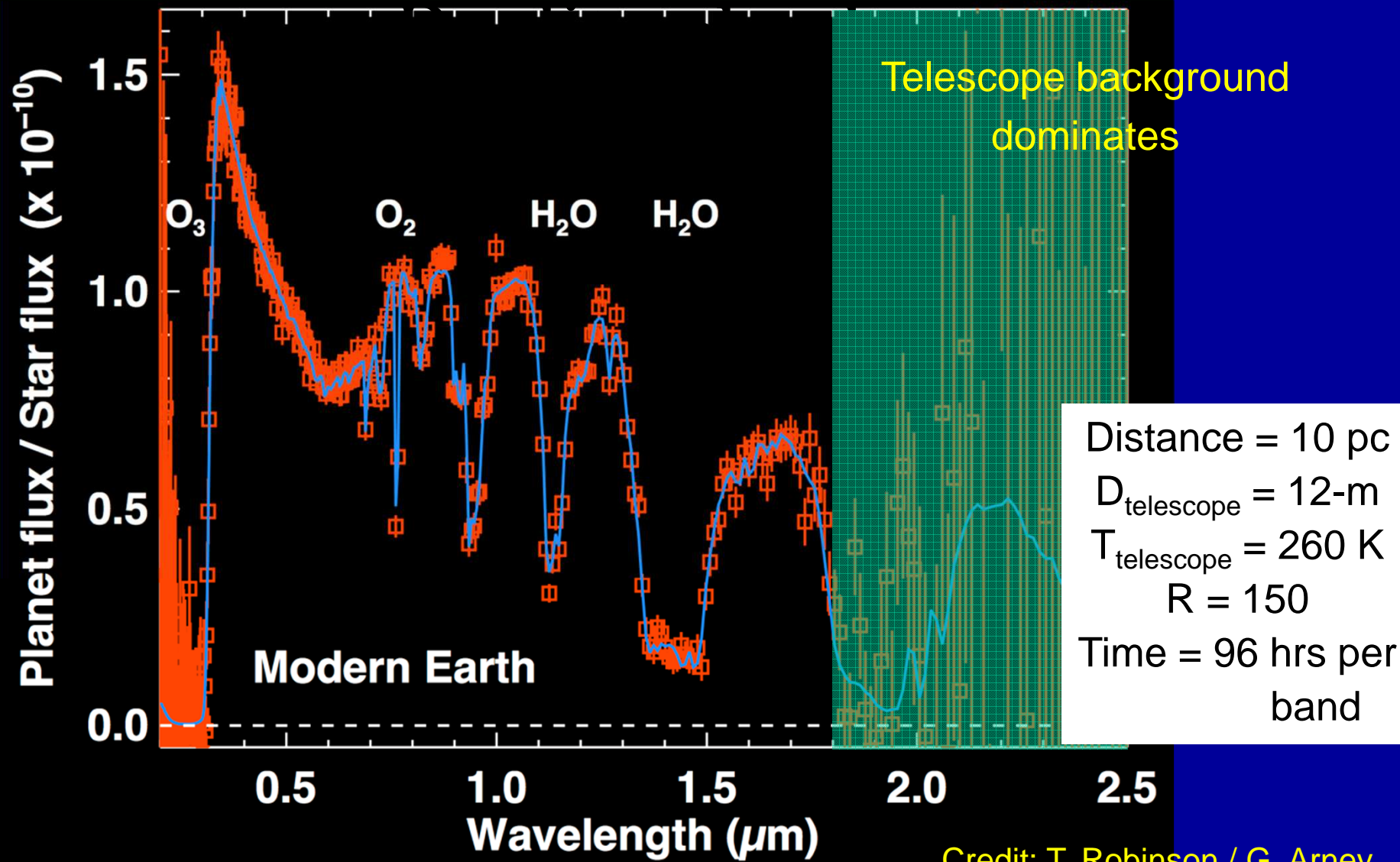
BLUE

RED

0.4 microns

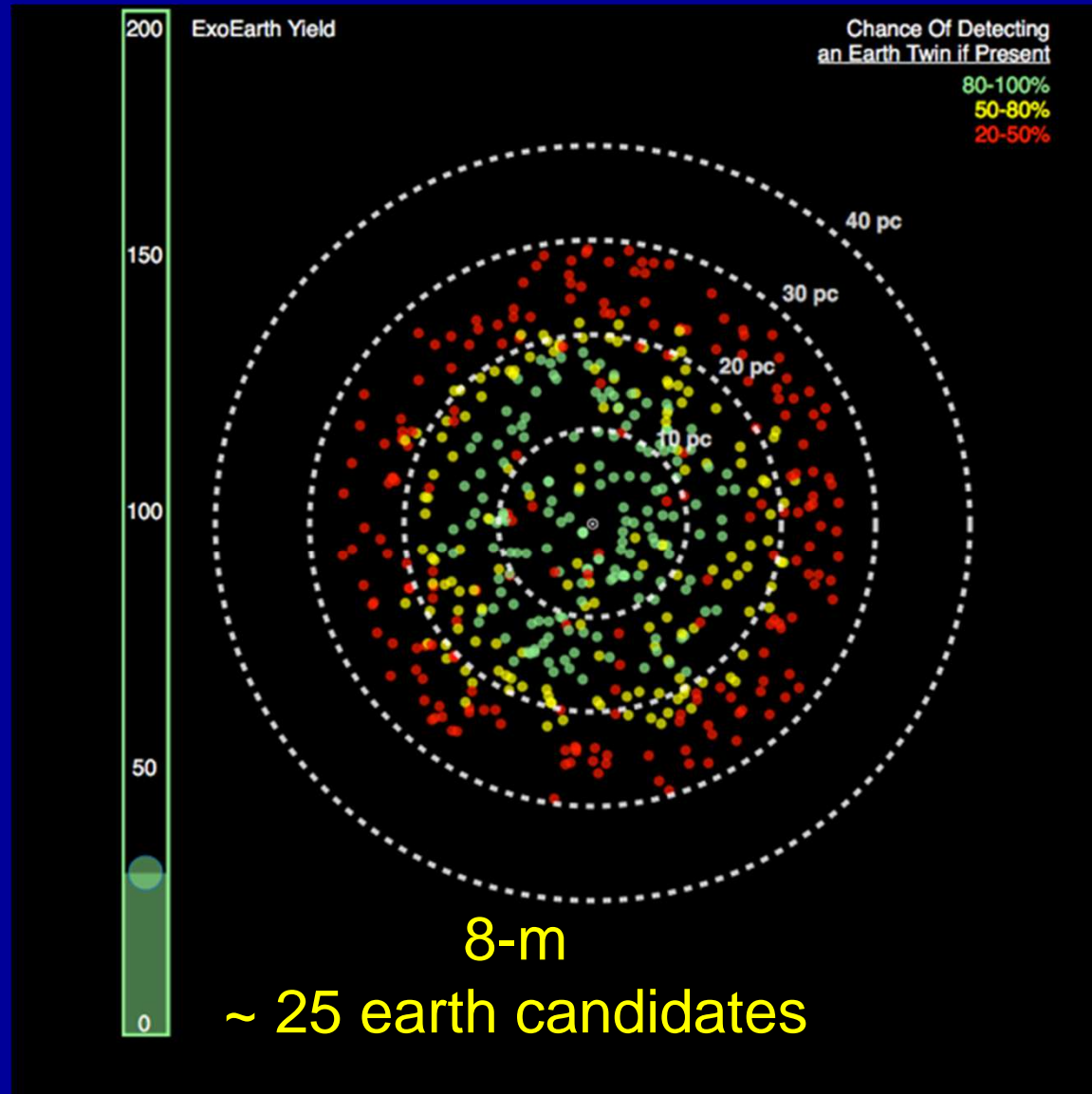
2.4 microns

Reality Check



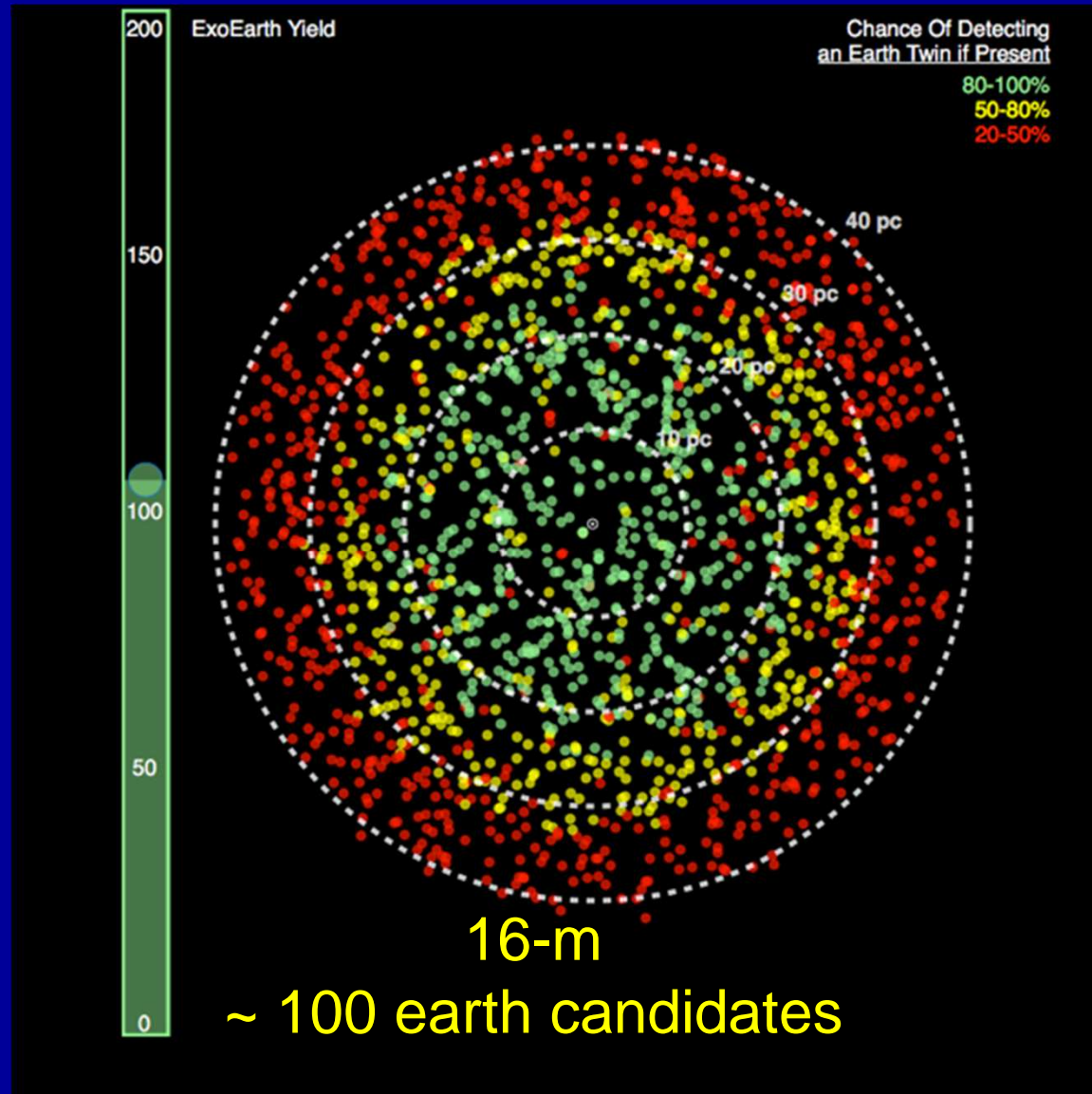
Credit: T. Robinson / G. Arney

ExoEarth candidates as function of aperture



Stark et al. (2014)

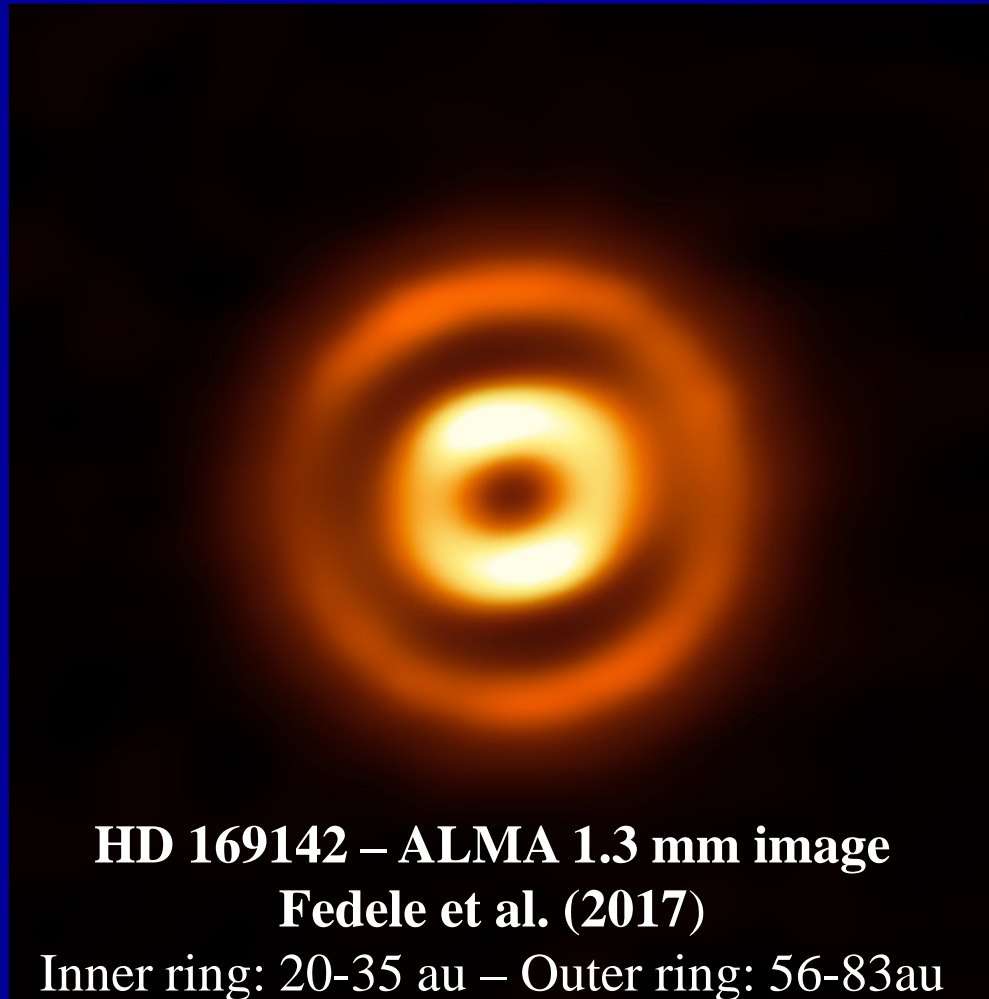
ExoEarth candidates as function of aperture



Stark et al. (2014)

Planeten in Scheiben

Niedrige Masse – Kleine Strukturen – Line/Kontinuum klein
Empfindlichkeit – Winkelauflösung – Hohe spektrale Auflösung



HD 169142 – ALMA 1.3 mm image

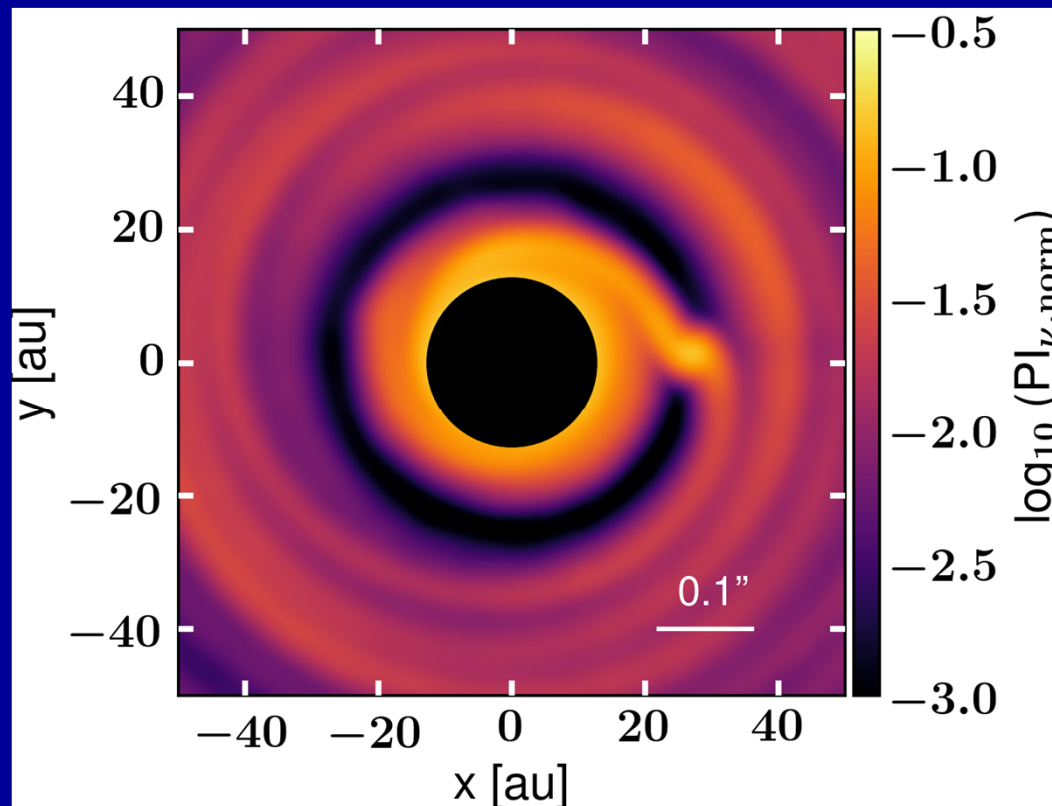
Fedele et al. (2017)

Inner ring: 20-35 au – Outer ring: 56-83au

LUVOIR Imaging

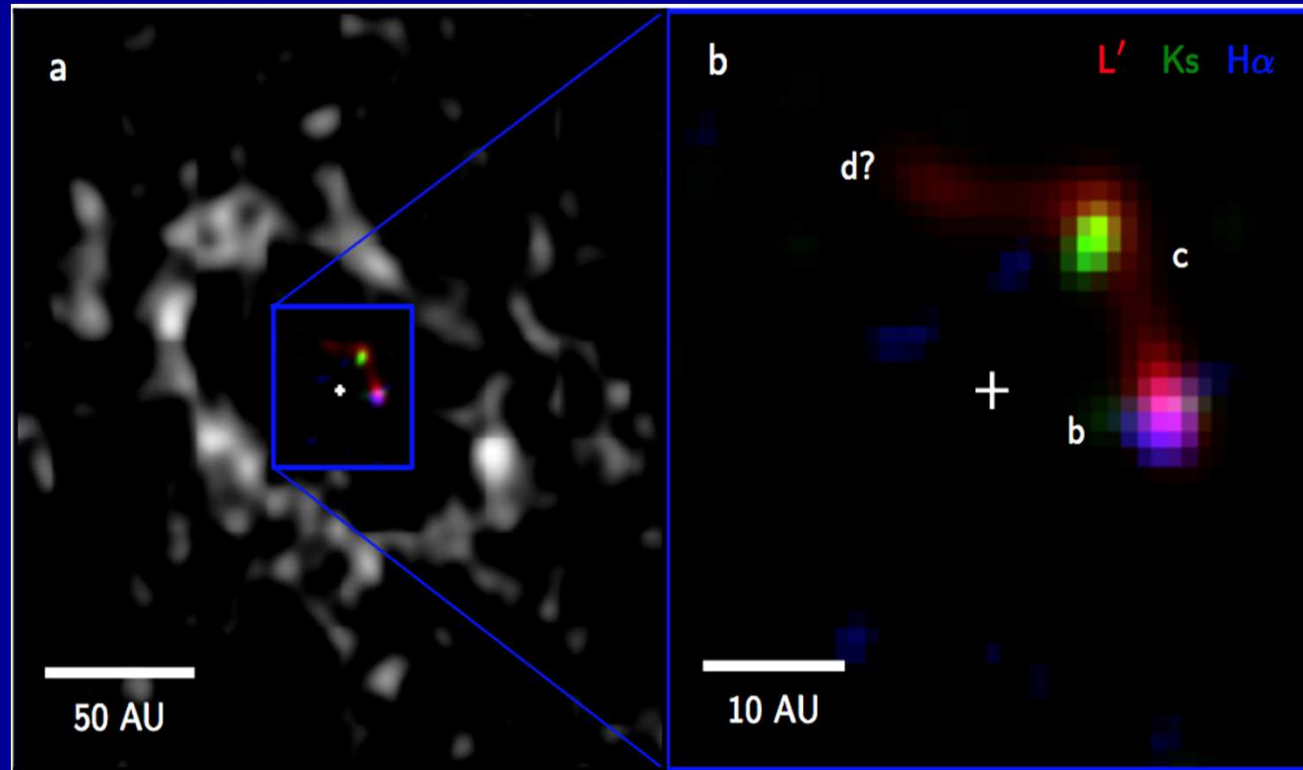
High-contrast **optical imaging** - Inner disk structures

IWA @ 600 nm = $2 \lambda / D \approx 15$ to 30 mas = 2 to 4 AU @ $d = 140$ pc



PI Image (inner $0.2''$ has been masked) –
Convolved with Gaussian PSF with FWHM = 20 mas
(MPIA Simulation after Pohl et al. 2015)

Suche nach eingebetteten Planeten - LkCa 15

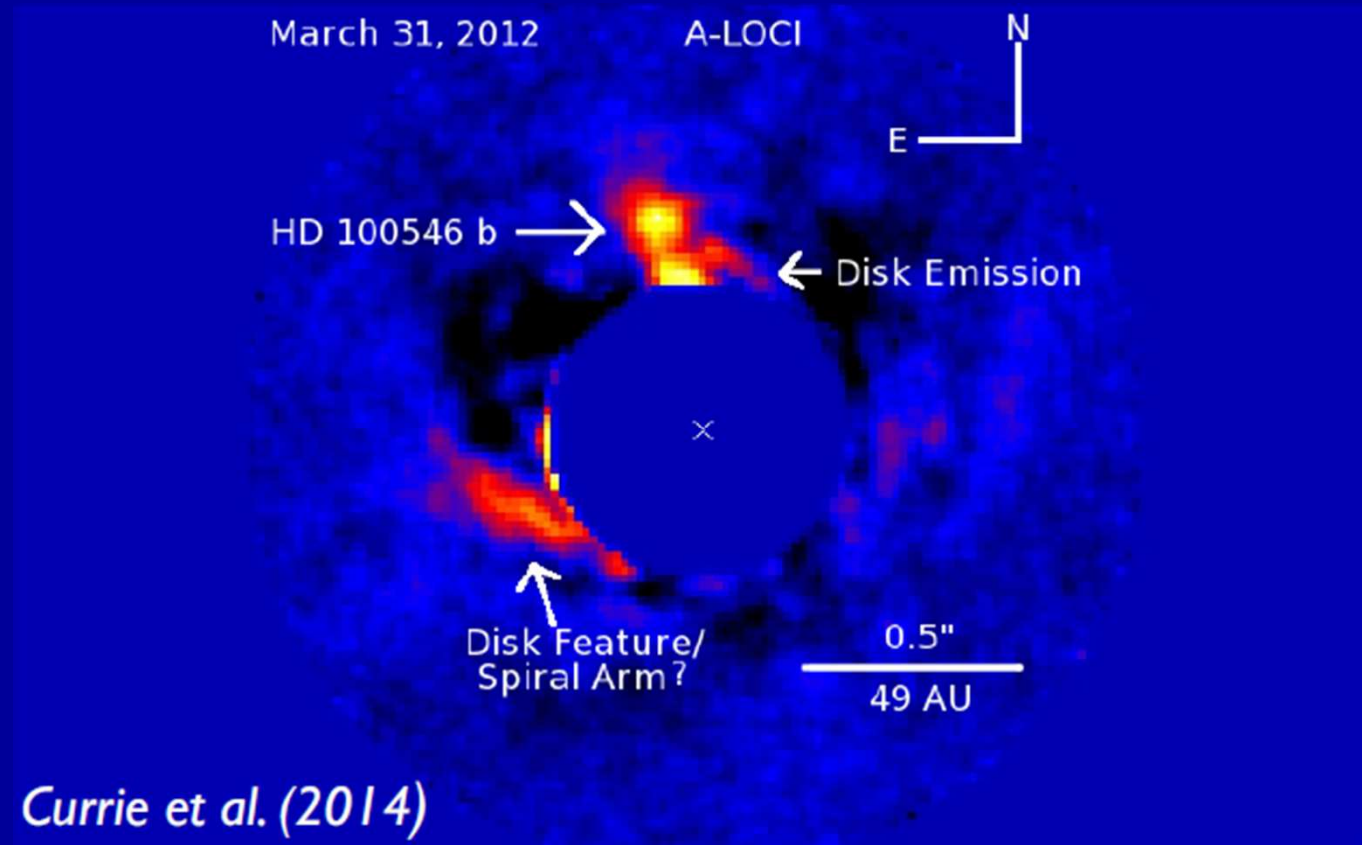


VLA Mm Observation of Disk H α (blue), Ks (green), L' (red)
Sallum et al. (2015)

Accretion rate = 5×10^{-10} Msun/yr : Planet/Star contrast for Jovian-mass planet is 10^{-4}
(with extinction up to a factor 100 smaller)

MagAO: 10^{-4} at 200 mas; Needed: 10^{-6} at 10 marcsec (20 marcsec)

Planeten in Protoplanetaren Scheiben



Inner Planet Candidate: 20-30 M_J @ 10 AU (Mulders et al. 2013, MIDI/VLTI,
See also Acke & van den Ancker 2006, Tatulli et al. 2011, Brittain et al. 2014)

Outer Planet Candidate: 15 M_J @ 70 AU (Quanz et al. 2013, 2015,
Currie et al. 2014, but Rameau et al. 2017)

Technologische Herausforderungen

- Kompatibilität zwischen UV & Koronagraphie
- Extrem hoher Kontrast mit segmentierten Teleskop
- Lichtwegkontrolle auf einem neuen Niveau
- Optische und Infrarotdetektoren mit sehr geringen Rauschen
- Effiziente Teleskopbetrieung für vernünftiges Sample



Technologische Entwicklungen

Entfaltung eines segmentierten Teleskops



Credit: A. Jones (GSFC)

LUVOIR

