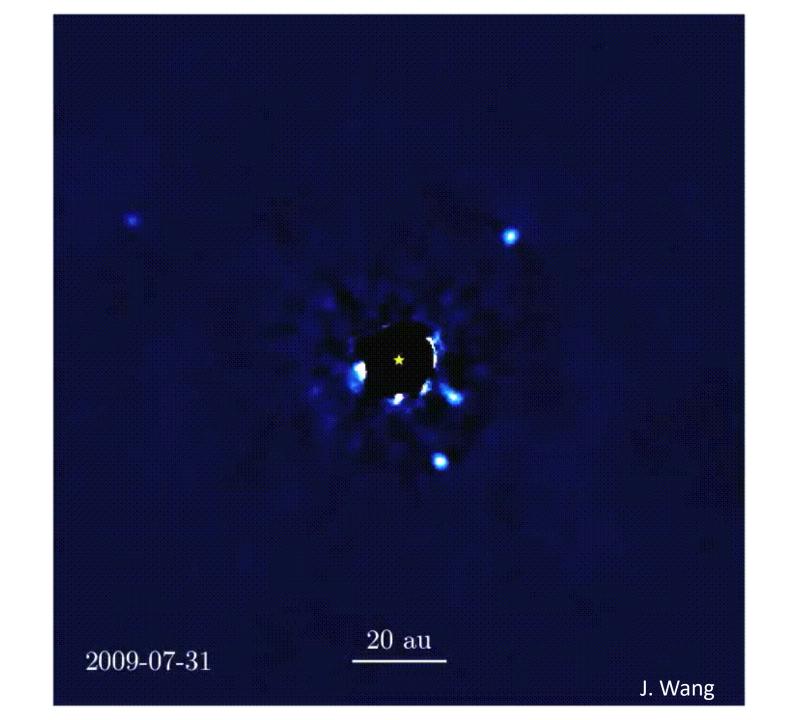
### Leiden EXoplanet Instrument

AstroTechTalk, Heidelberg, 22-02-2019

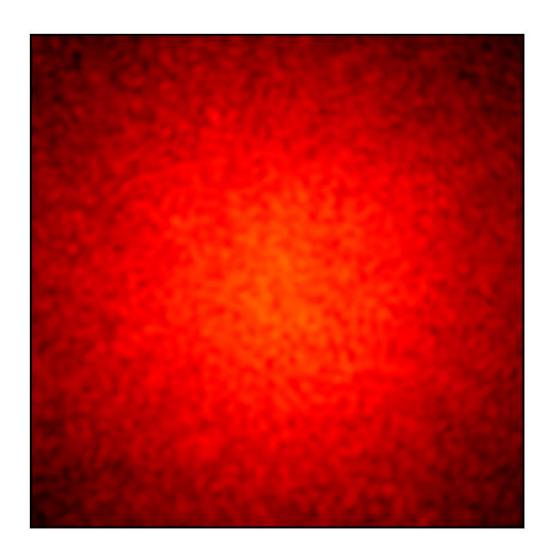
Sebastiaan Haffert, Christoph Keller, Ignas Snellen Leiden Observatory



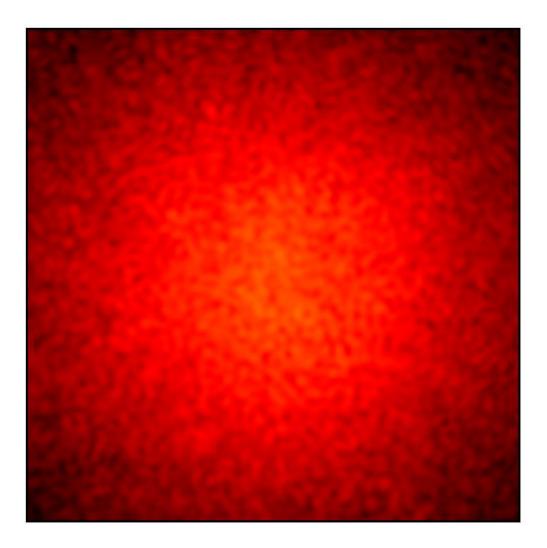
Credits: ESO/M. Kornmesser



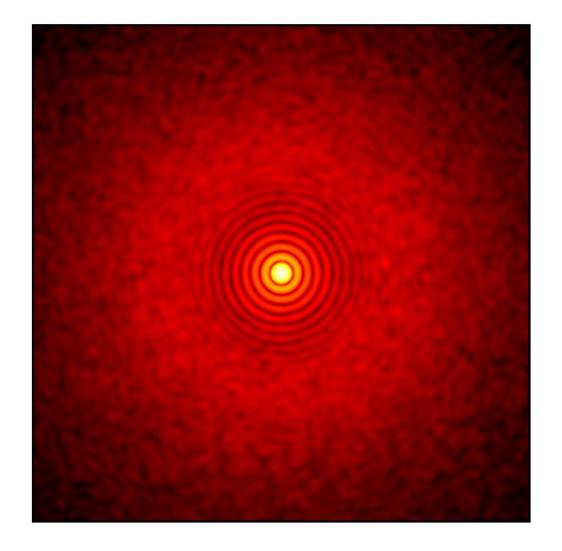
# Imaging through the atmosphere

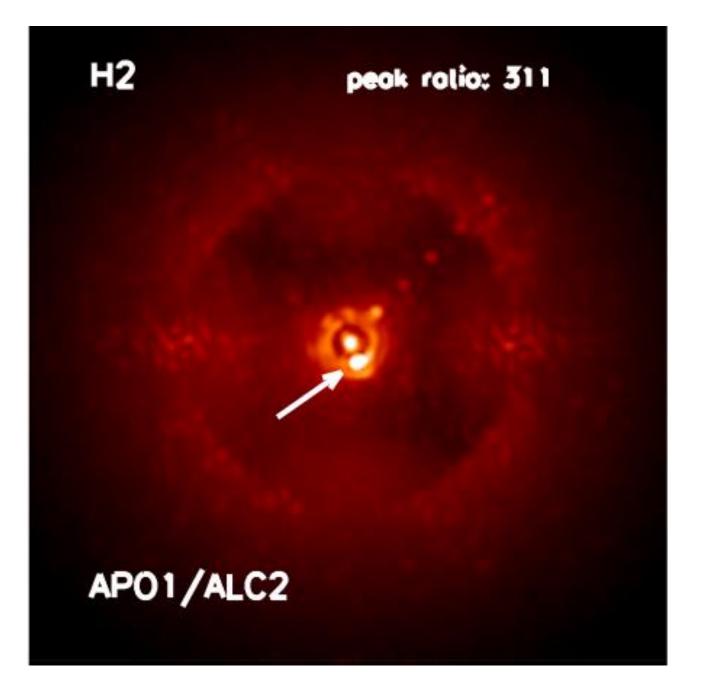


## Imaging through the atmosphere

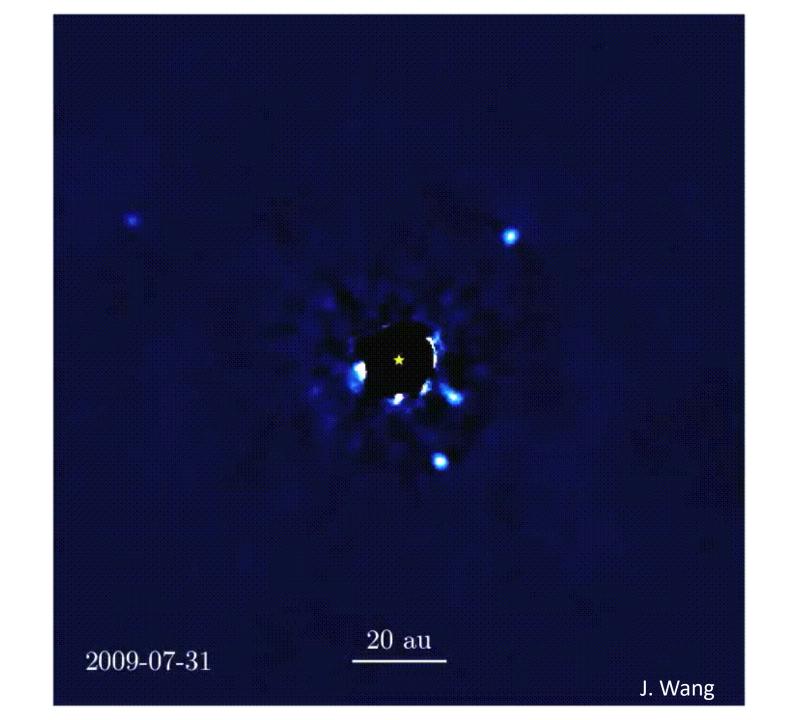


### Adaptive optics corrected

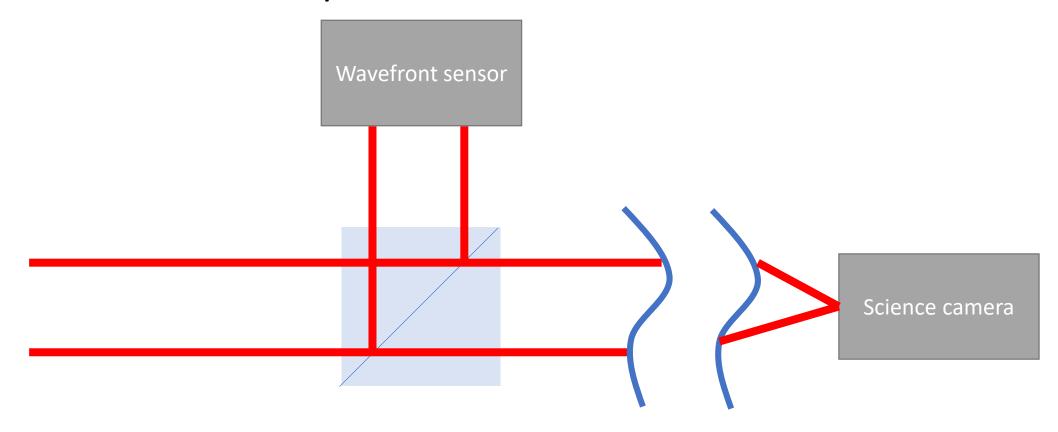




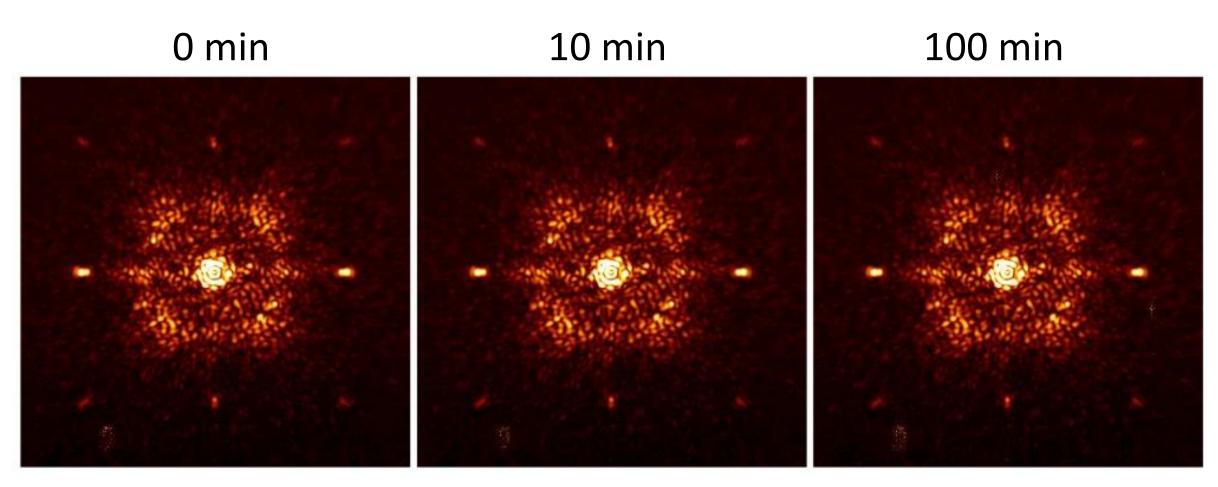
J-L. Beuzit et al 2019



### Non-common path abberations



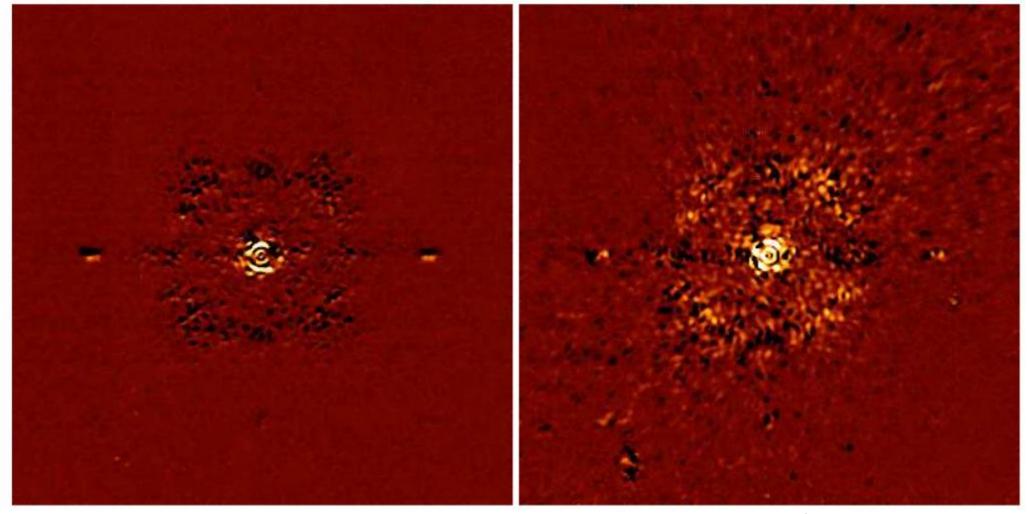
### Temporal stability



Martinez et al. 2013

### Temporal stability 10 min – 0 min

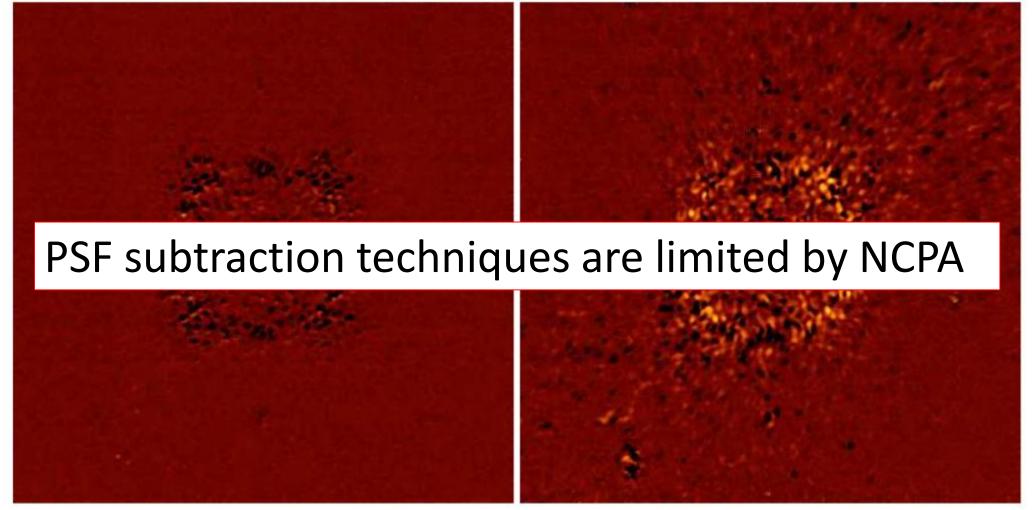
100 min – 0 min



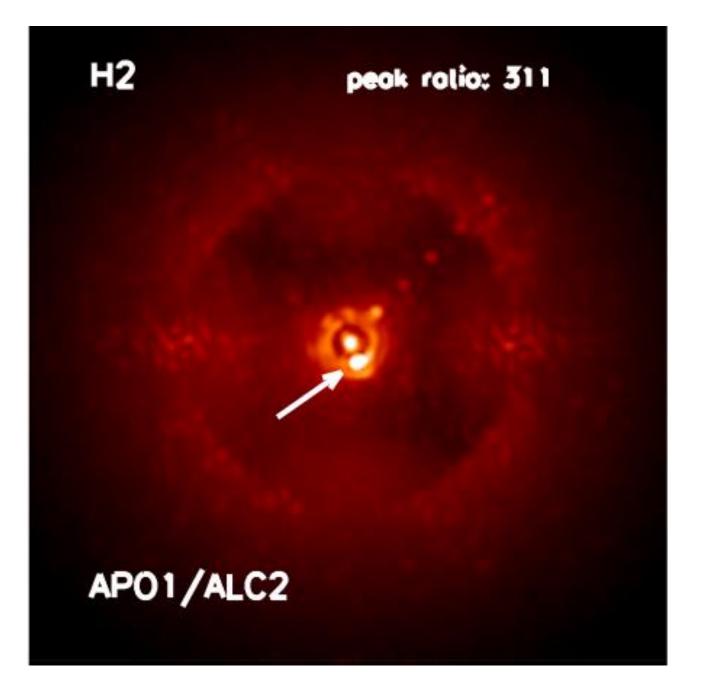
Martinez et al. 2013

### Temporal stability 10 min – 0 min

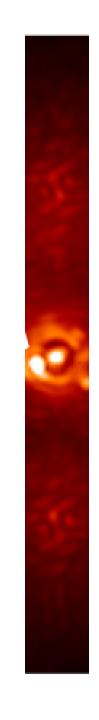
100 min – 0 min



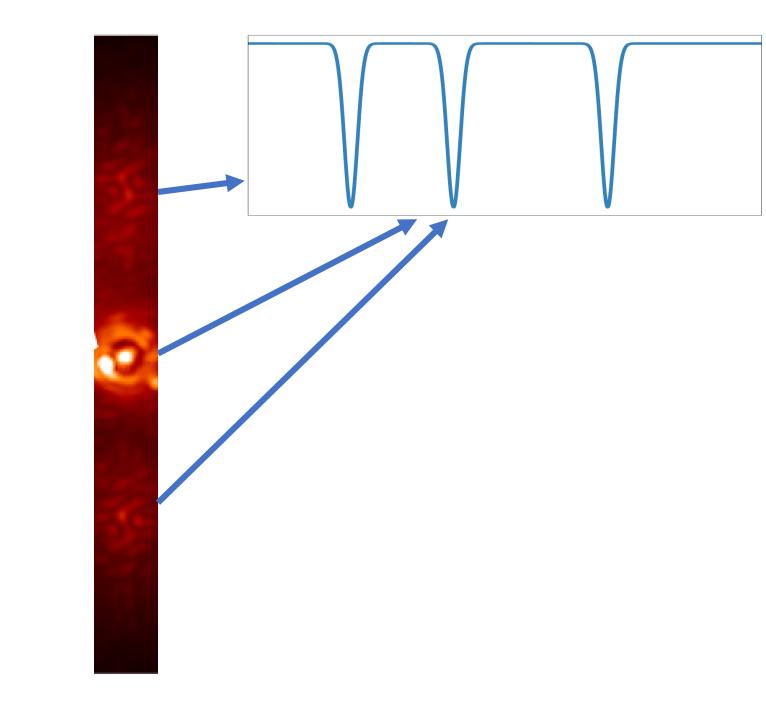
Martinez et al. 2013

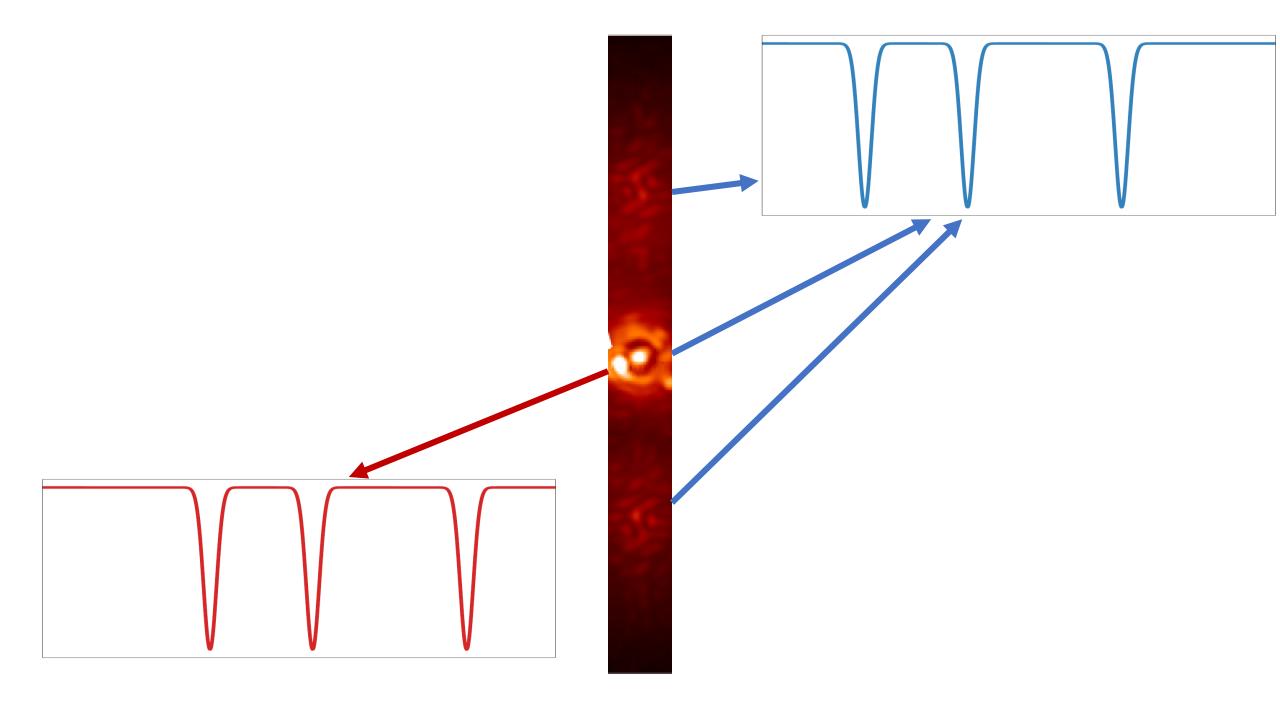


J-L. Beuzit et al 2019



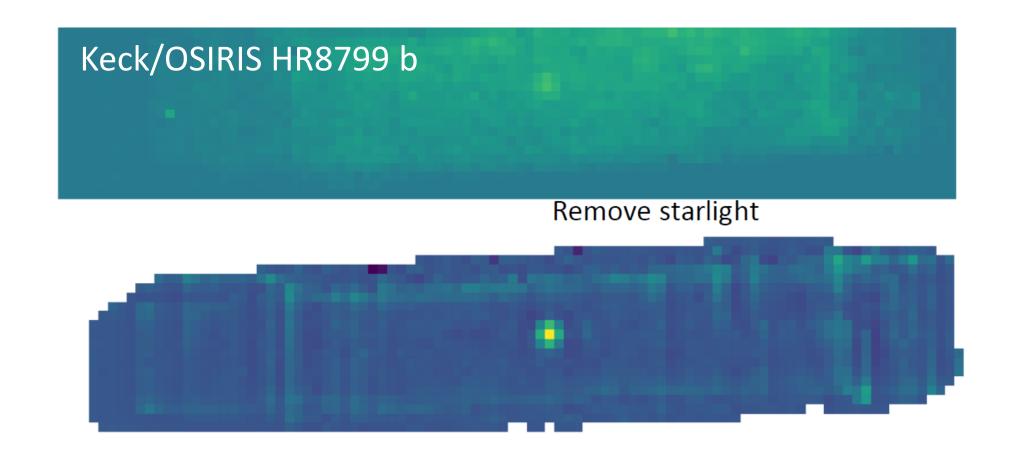
J-L. Beuzit et al 2019

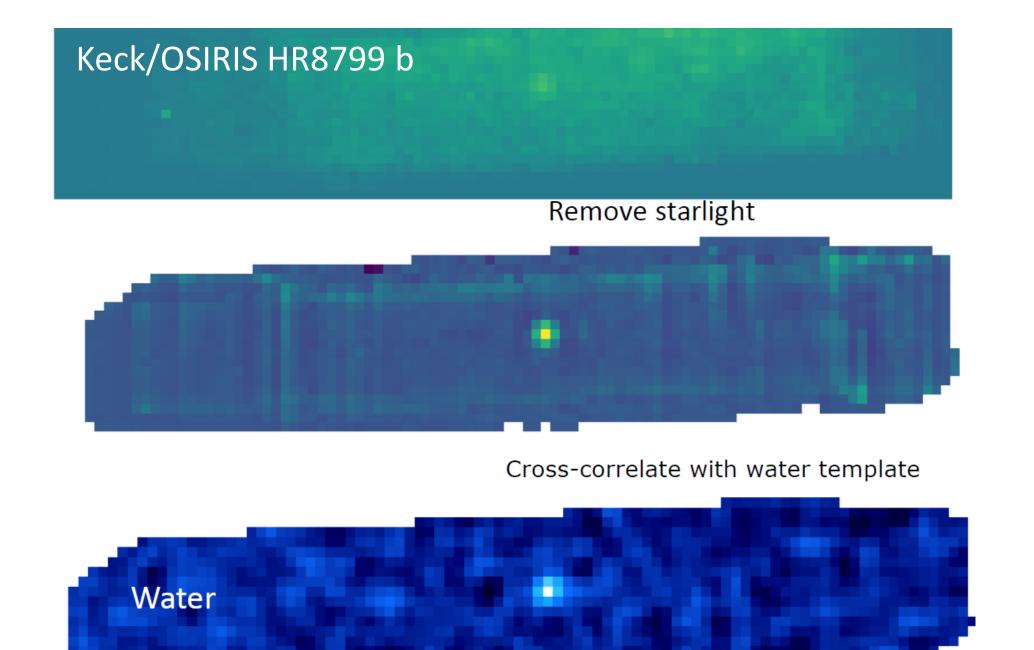




Lets apply this technique

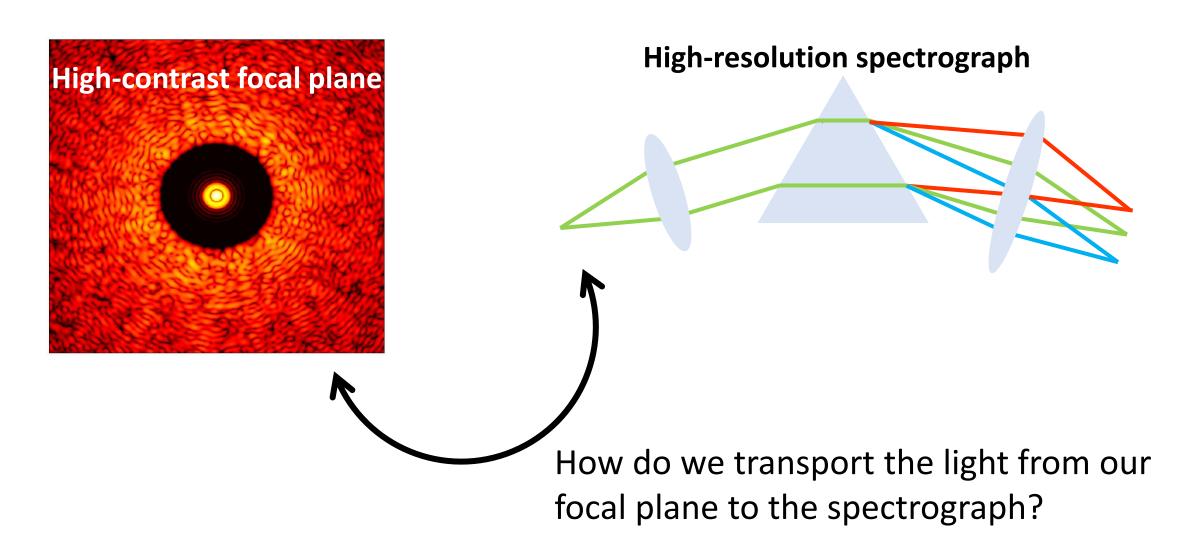
### Keck/OSIRIS HR8799 b



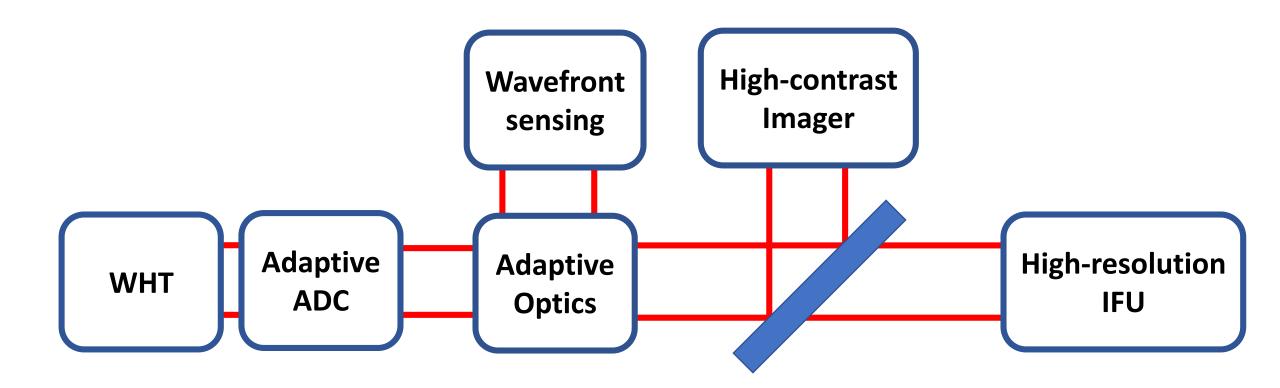


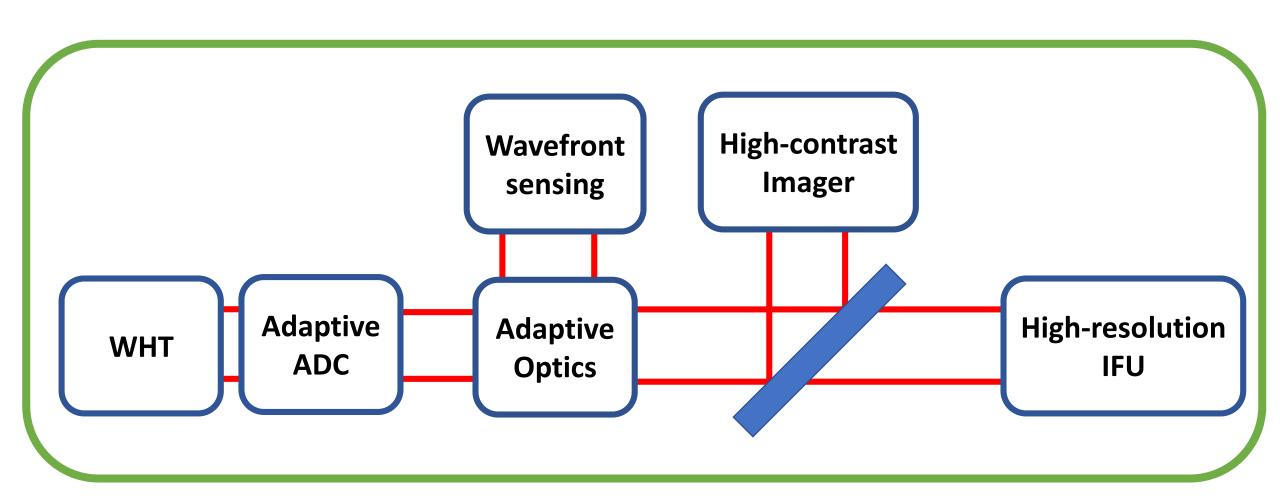
See Hoeijmakers et al 2018, Petit dit de la Roche et al. 2018

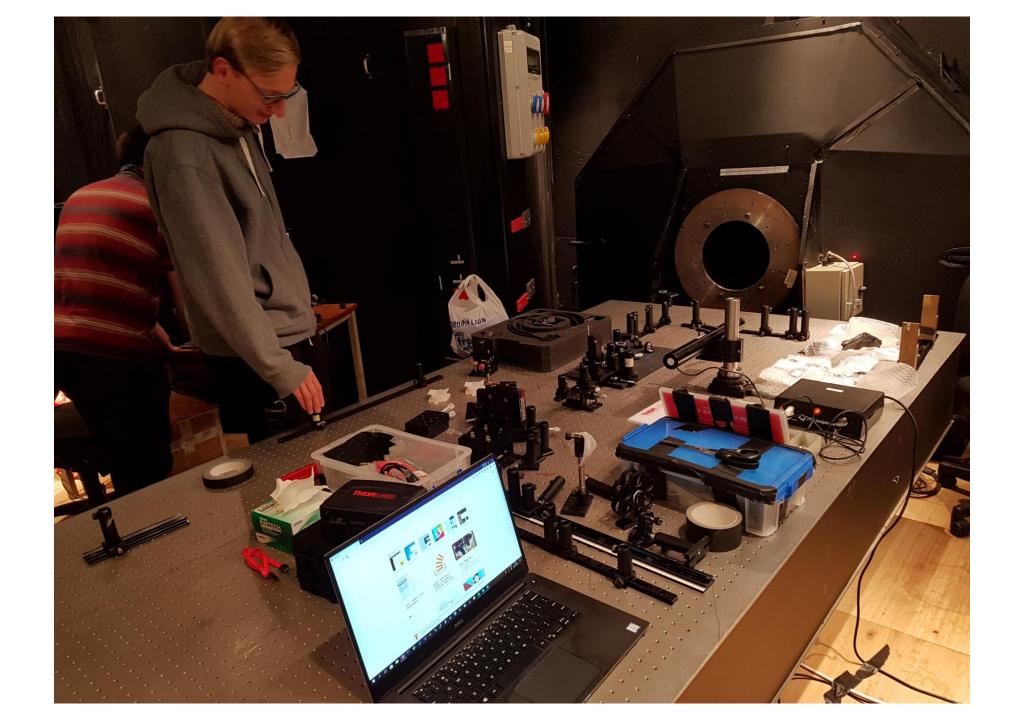
### How to couple HCI with HRS

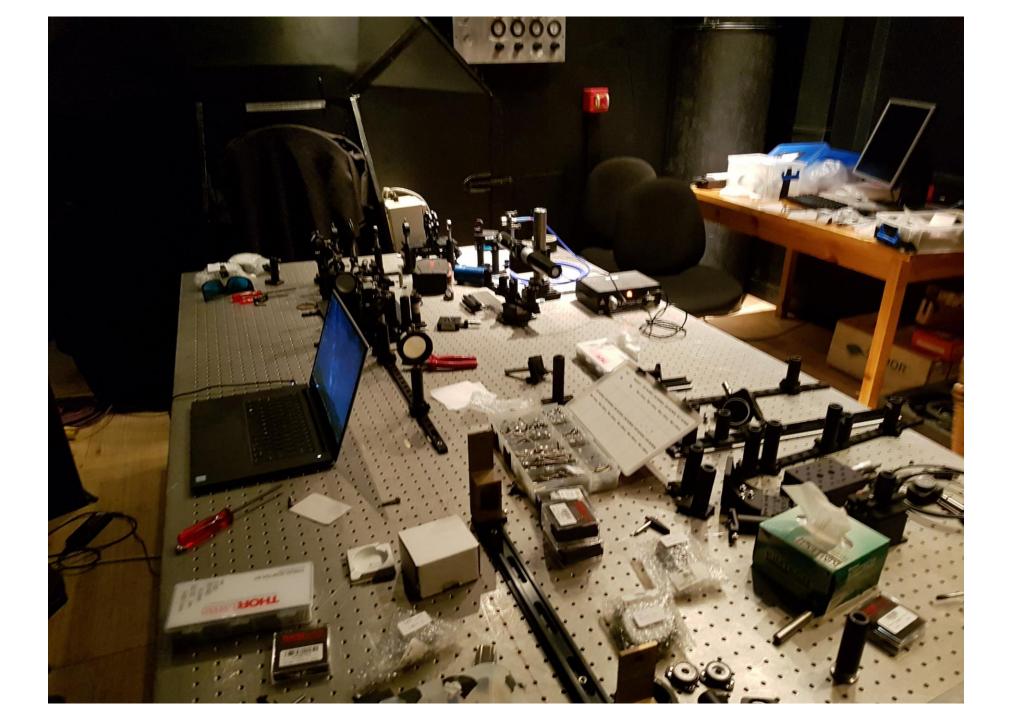


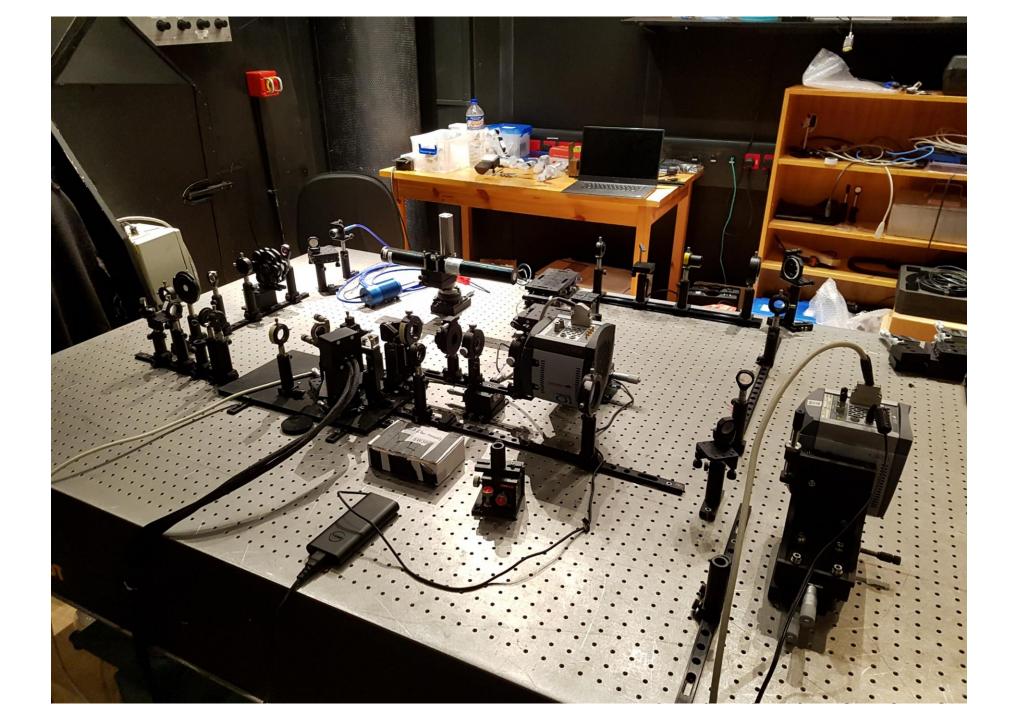
### The Leiden EXoplanet Instrument

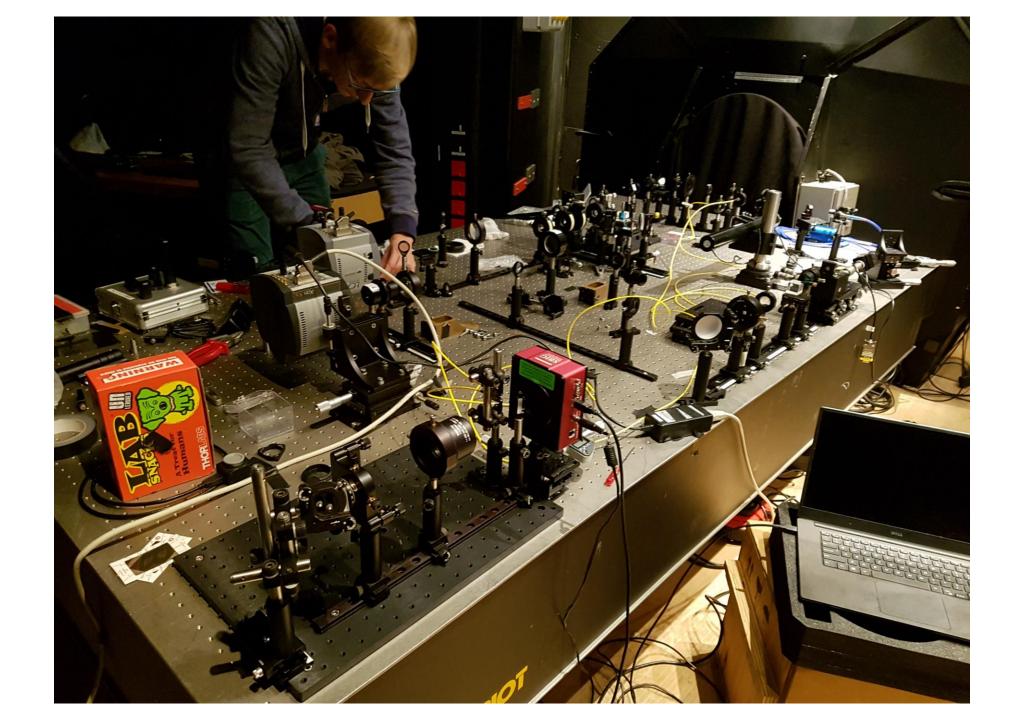


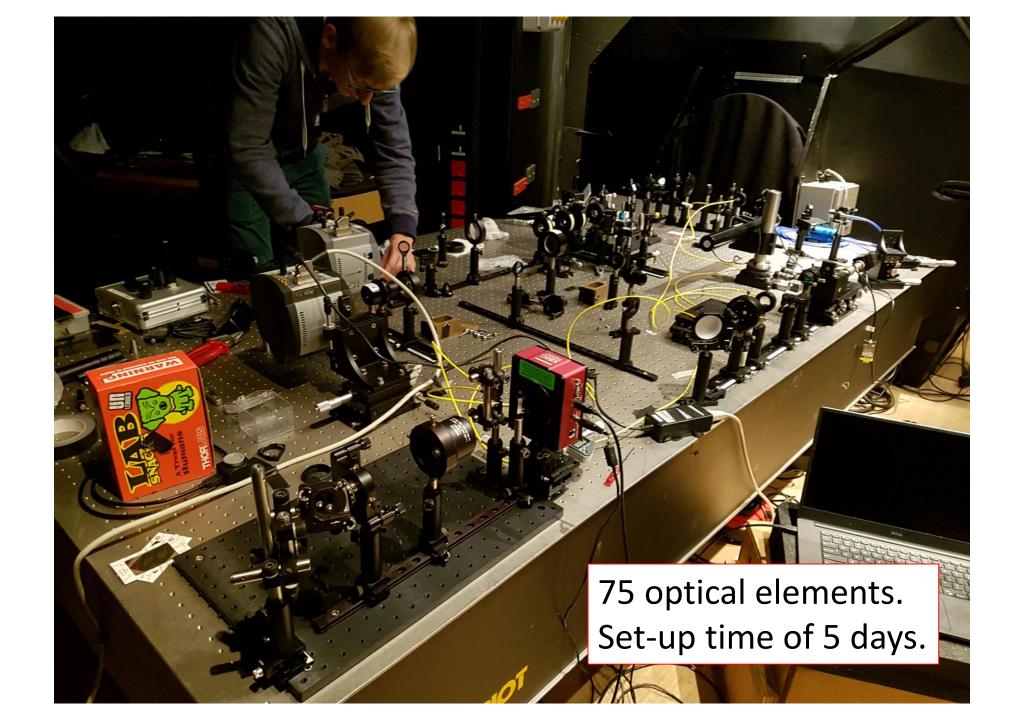


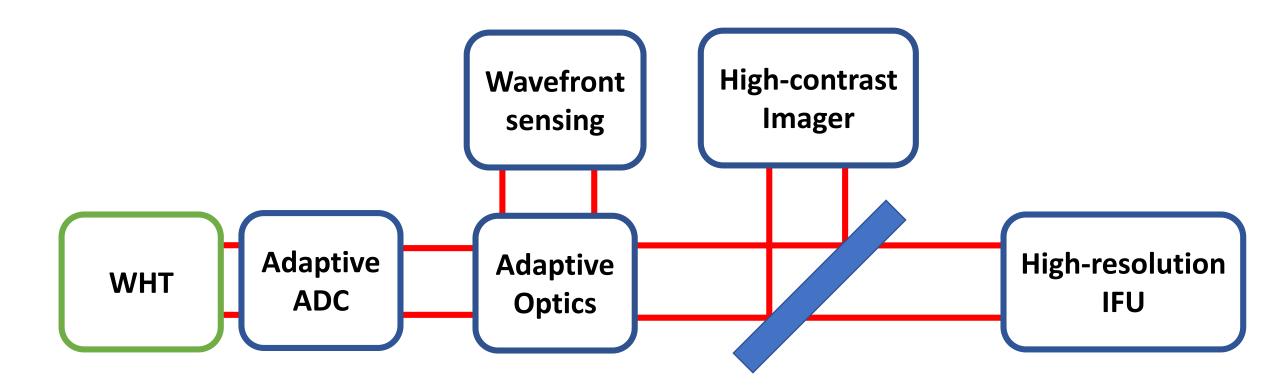




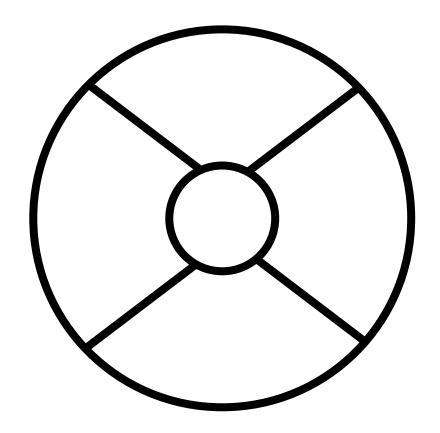






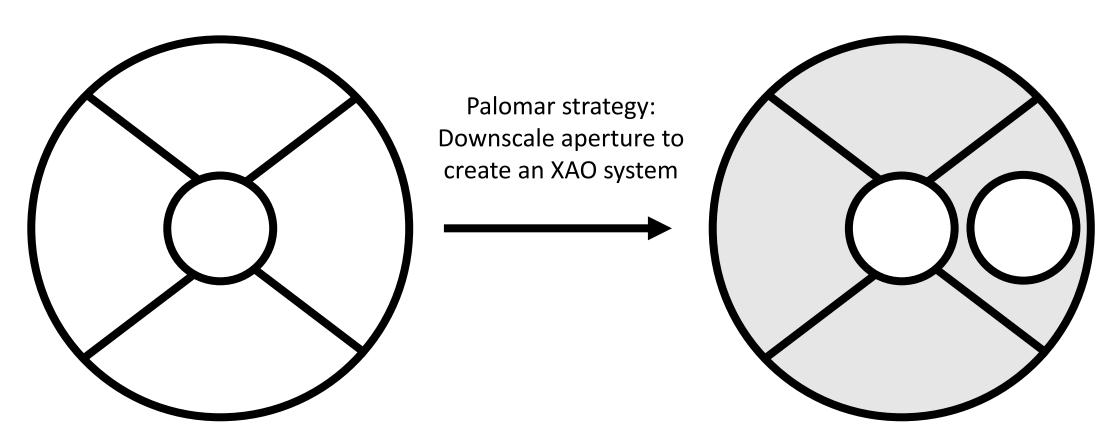


### Evolution of LEXI XAO

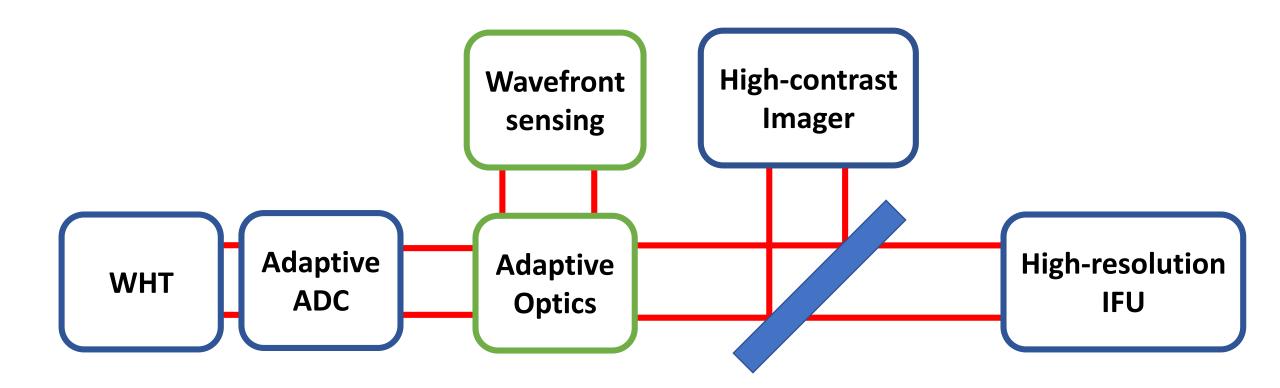


William Herschel Telescope 4.2m diameter LEXI run June 2016

#### Evolution of LEXI XAO



William Herschel Telescope 4.2m diameter LEXI run June 2016 William Herschel Telescope 1.2m off-axis segment LEXI run December 2017/2018



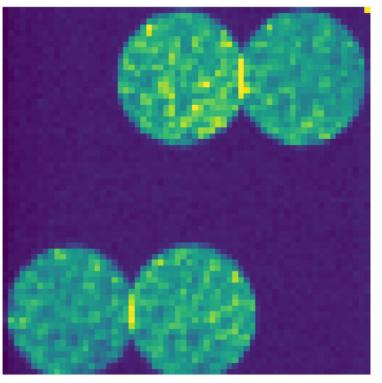
#### Shack-Hartmann wavefront sensor

AO speed: 500 Hz

Number of modes: 75-80

Alpao 97-15 DM

## generalised Optical Differentiation wavefront sensor

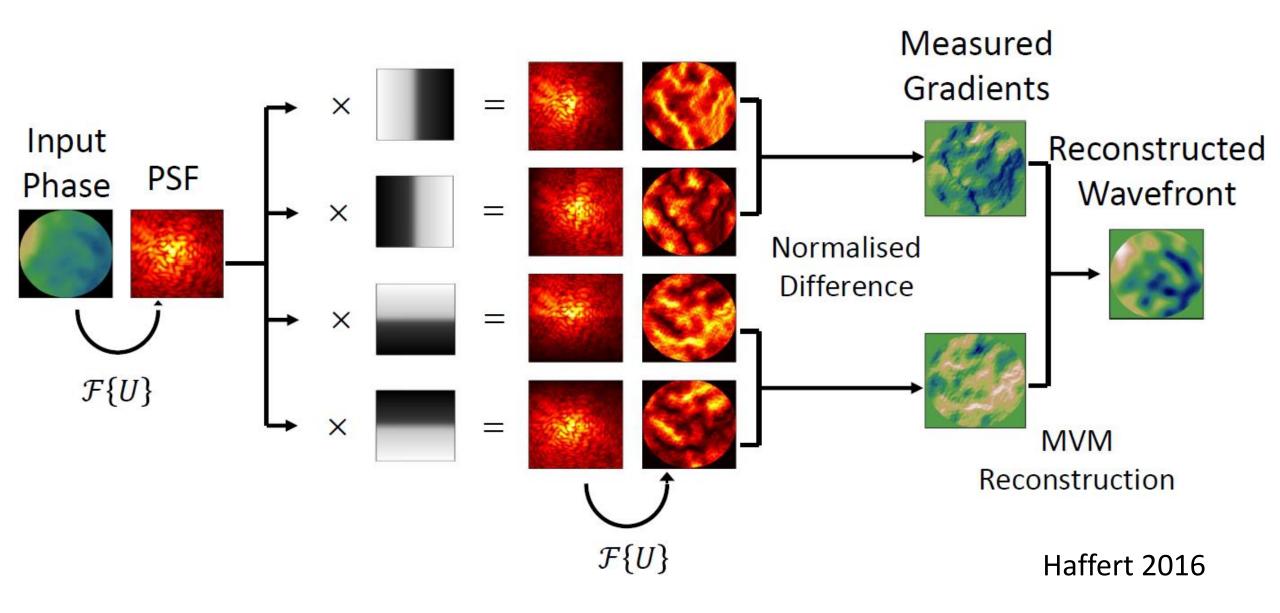


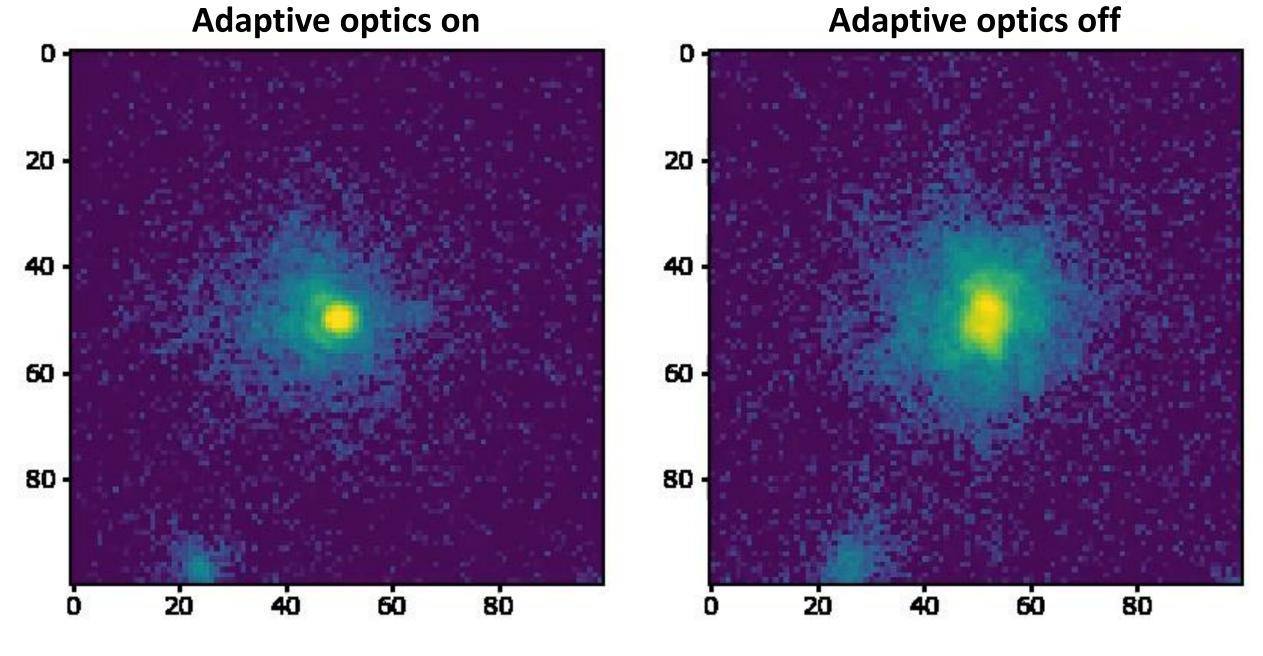
AO speed: 800 Hz

Number of modes: 96

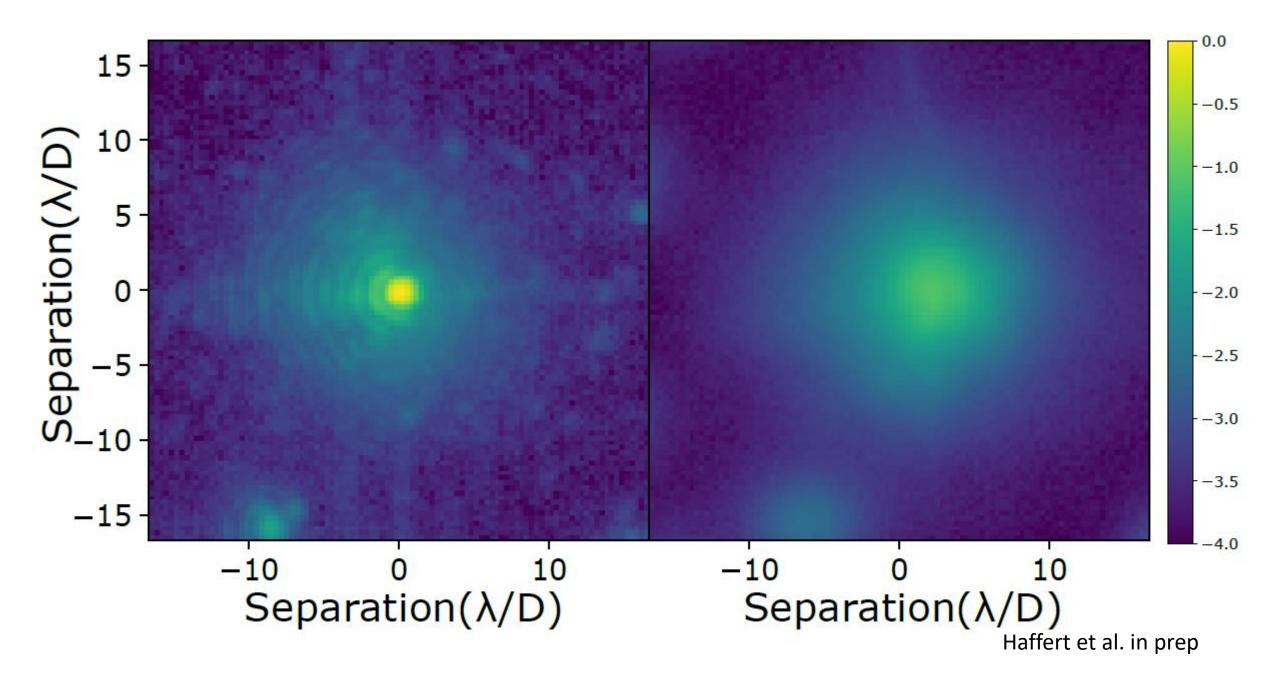
Alpao 97-15 DM

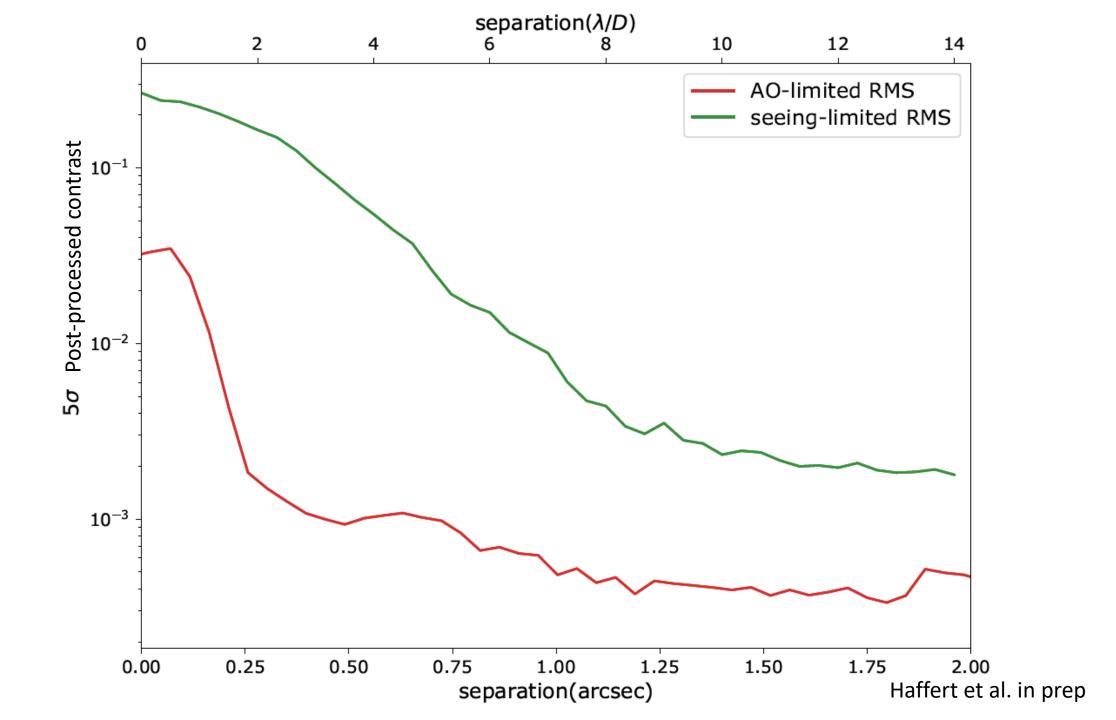
Haffert 2016, Haffert et al. 2019 (in prep)



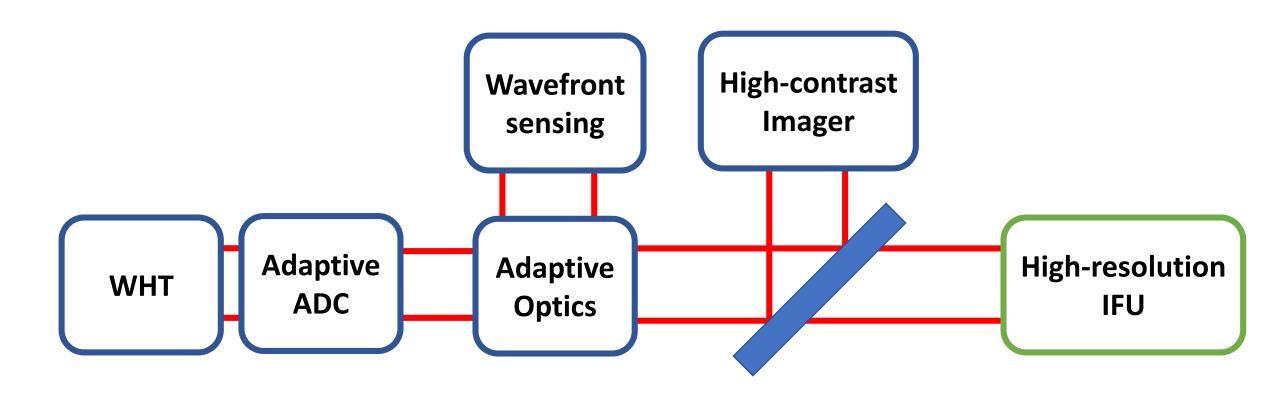


Haffert et al. in prep

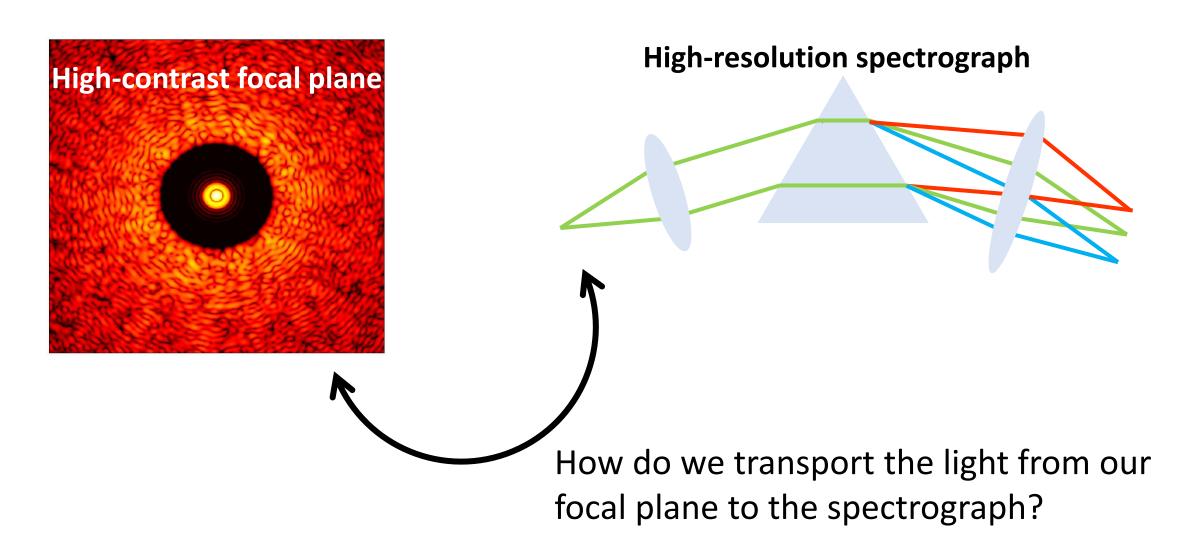




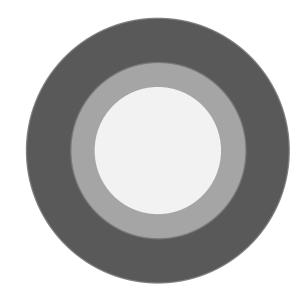
### LEXI



## How to couple HCI with HRS



### Fibers



Multi-mode ("seeing-limited") fiber



Single-mode ("diffraction-limited") fiber

### Fibers



Multi-mode ("seeing-limited") fiber

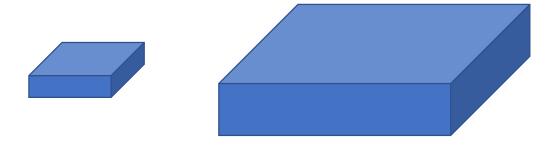


Single-mode ("diffraction-limited") fiber

## Advantaged of single-mode fibers



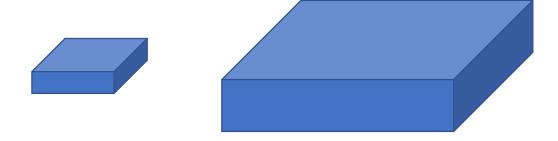
SMFs are diffraction-limited fibers And spectrographs scale with angular size



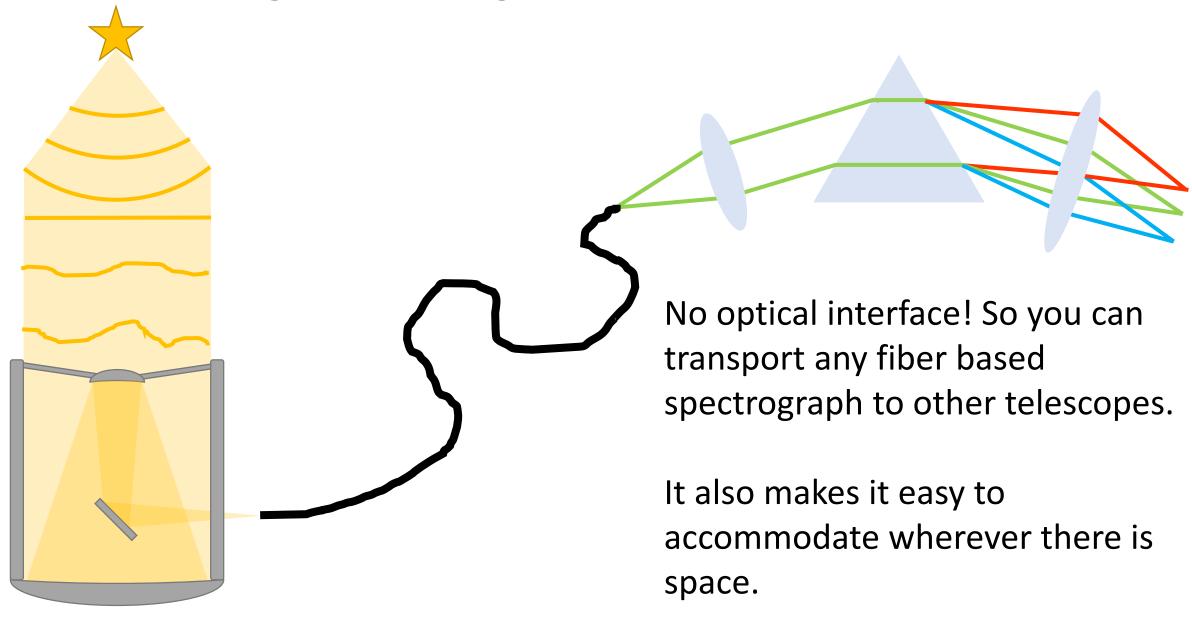
## Advantaged of single-mode fibers



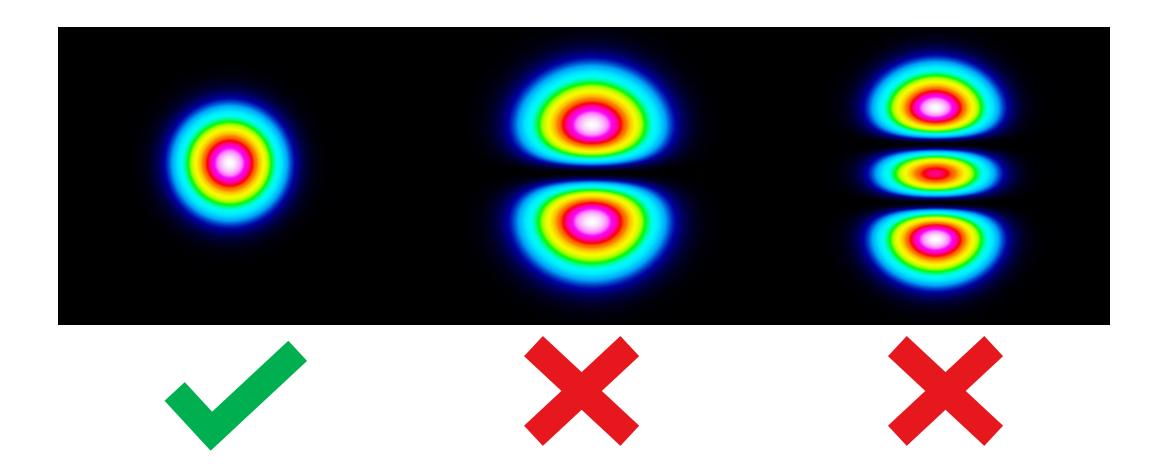
SMF spectrographs are small!



## Advantaged of single-mode fibers



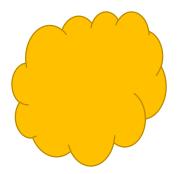
# Advantages of single-mode fibers



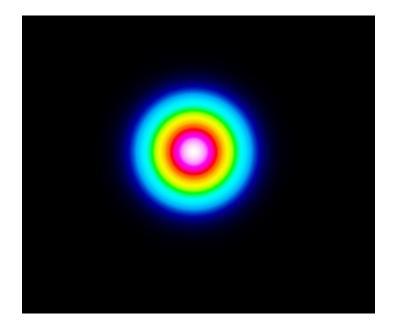
# Advantages of single-mode fibers



All inputs always transform into a gaussian.



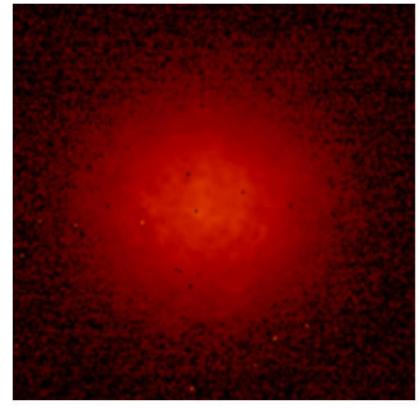




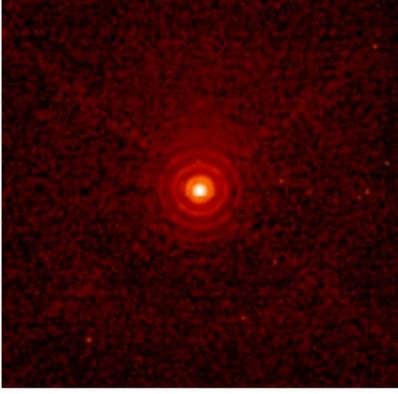


Creates very stable Line Spread Functions.

## Advantages of single-mode fibers



Seeing limited image  $5.2 \pm 2\%$  SR (a)



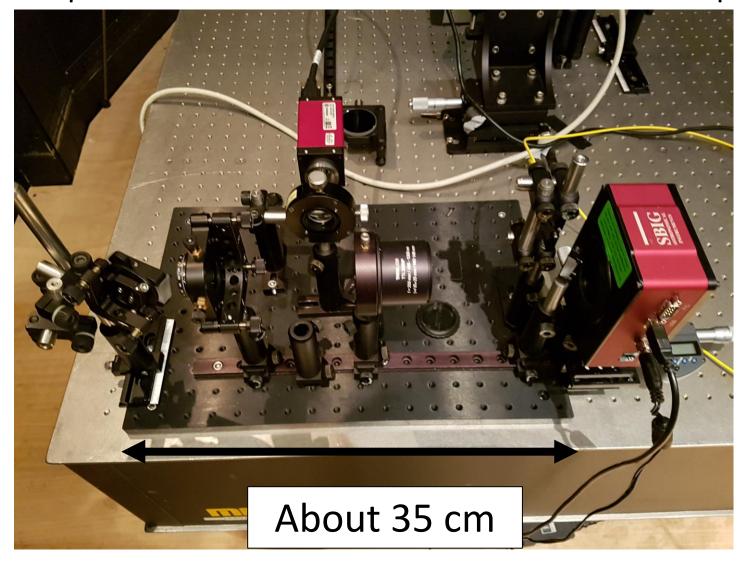
AO corrected image  $90.3 \pm 2\%$  SR (b)

We put in all the effort to make our instrument diffraction-limited!

Why convert it back to the non-diffraction limit?

Petit et al. 2016

#### Compact fiber-fed diffraction-limited spectrograph designs



#### **Specifications**

R of 100000

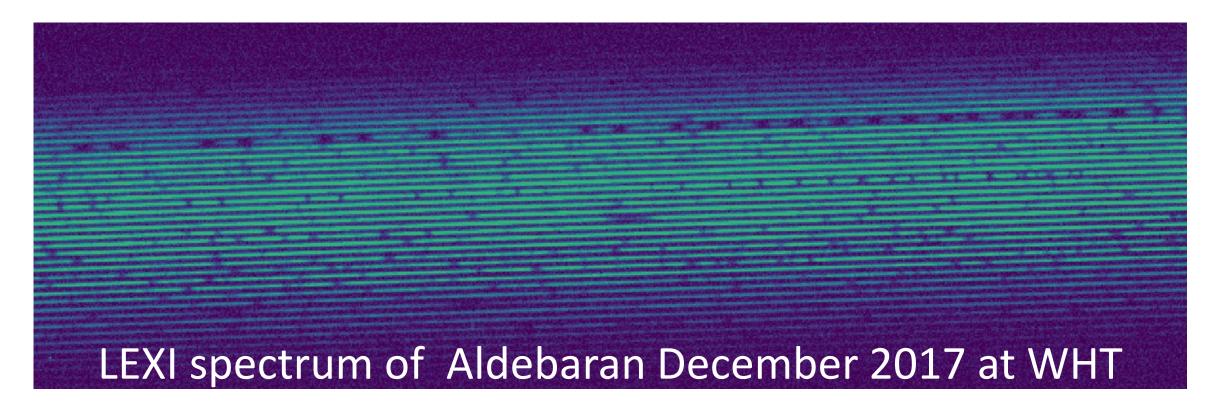
600 to 900 nm

19 fibers

Designed with SCAR

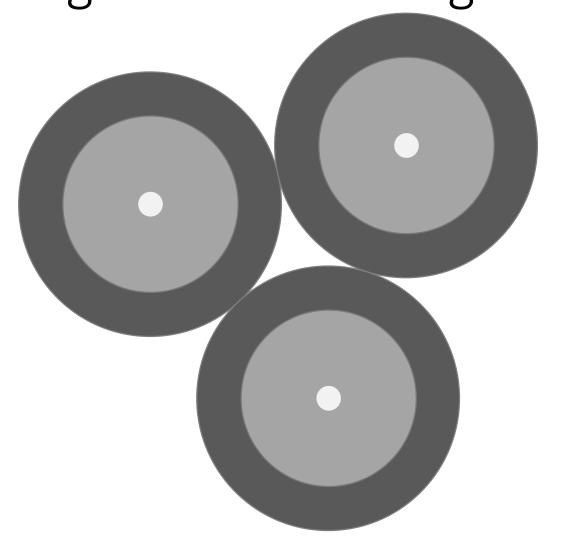
Off-the-shelf parts

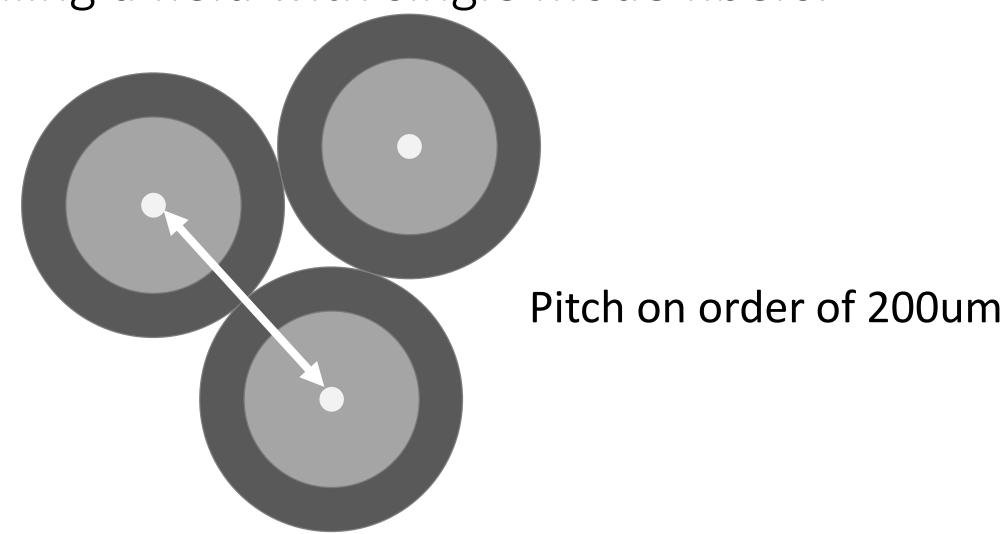
It is easy to design high-packing efficiencies due to the diffraction-limited spectrograph design.

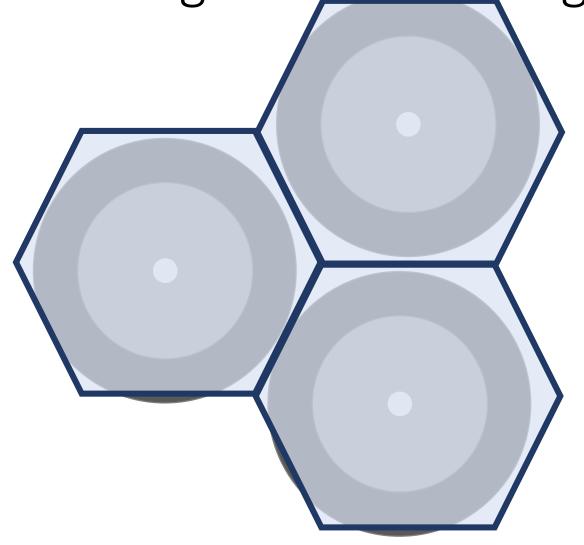


Now let's switch to an IFU instead







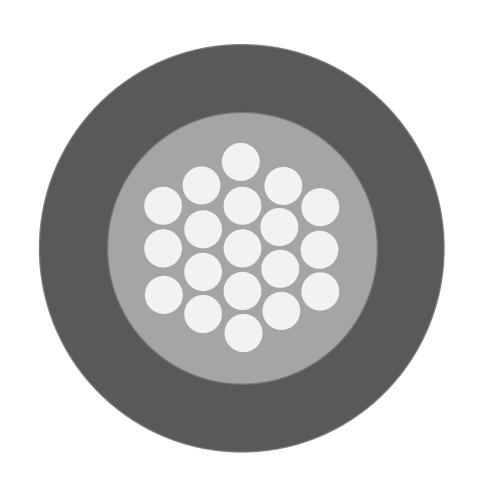


Micron alignment accuracy

A MLA is a piece of bulk optic

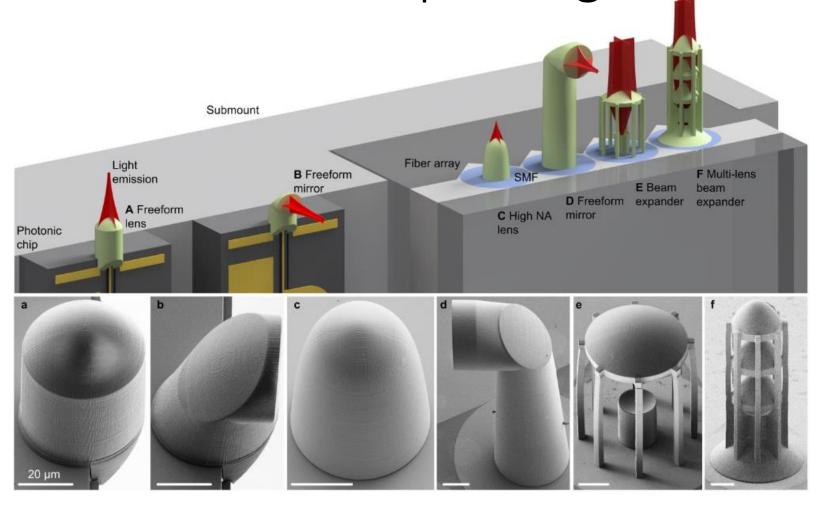
Fibers are not in a perfect grid.

### Use a **SINGLE** multi-core fibre



- Excellent core to pitch ratio
- Only a single fiber.
- Dense system.
- Small size is difficult with bulk optics

In-Situ 3D Nano-printing of freeform optics

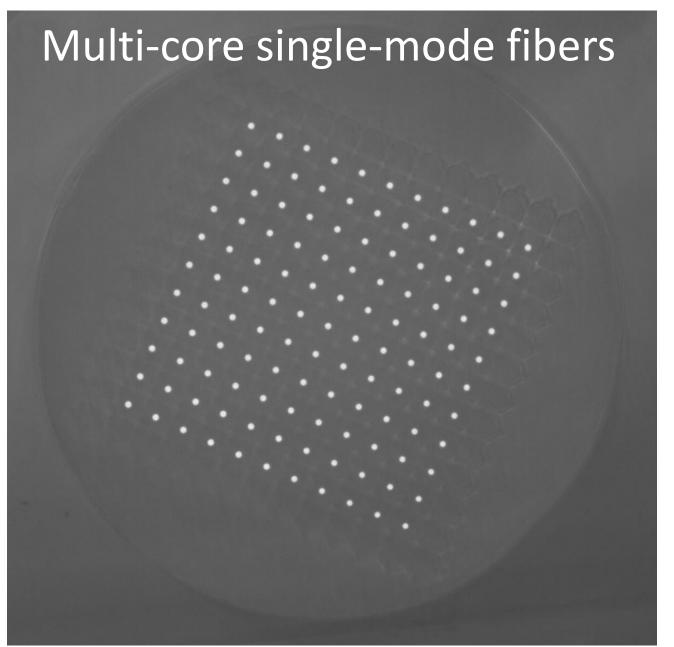


"In-Situ 3D Nano-Printing of Freeform Coupling Elements for Hybrid Photonic Integration" Dietrich et al. 2018

Pitch of 10.5 um

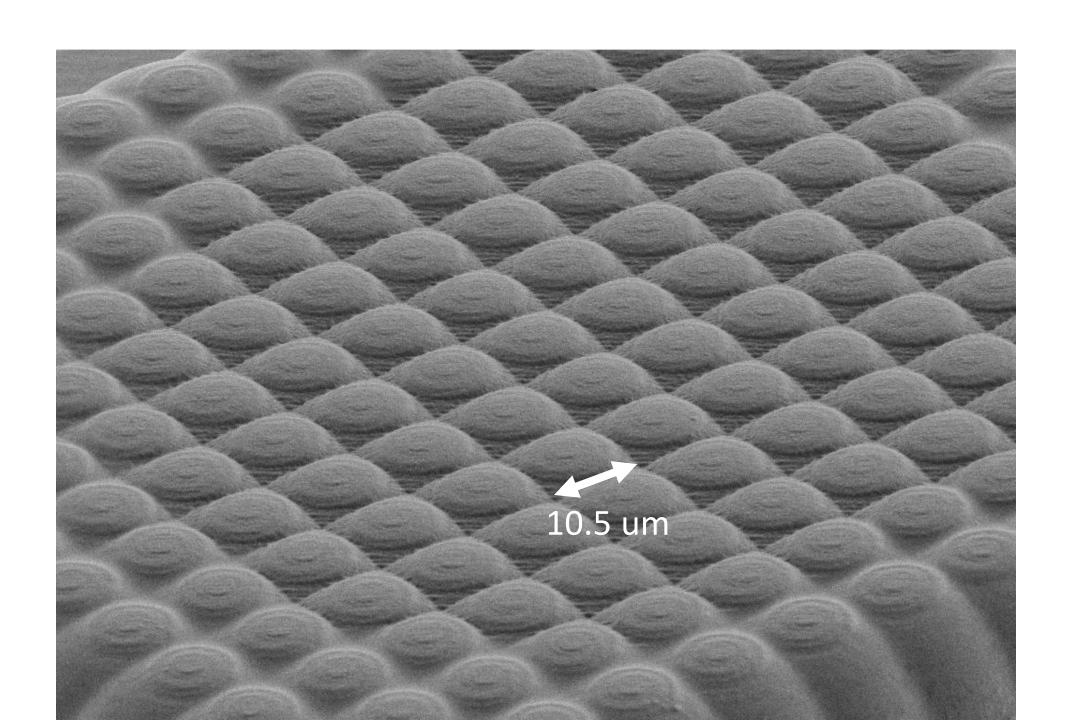
1.83um MFD

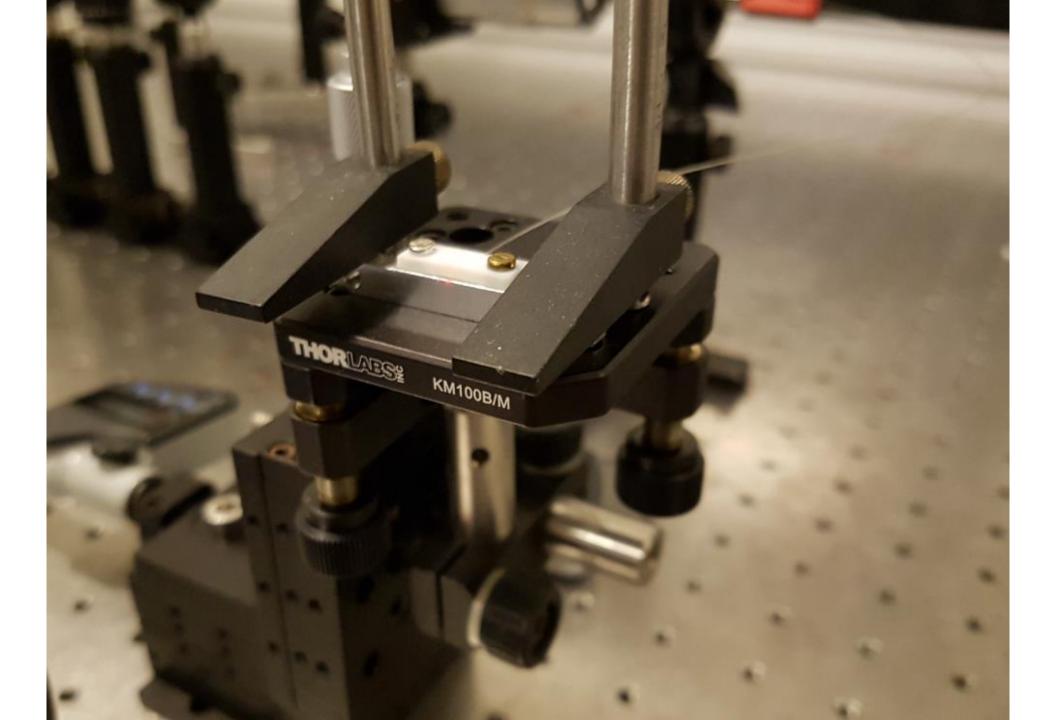
Strict requirements on microlens alignment from SCAR.



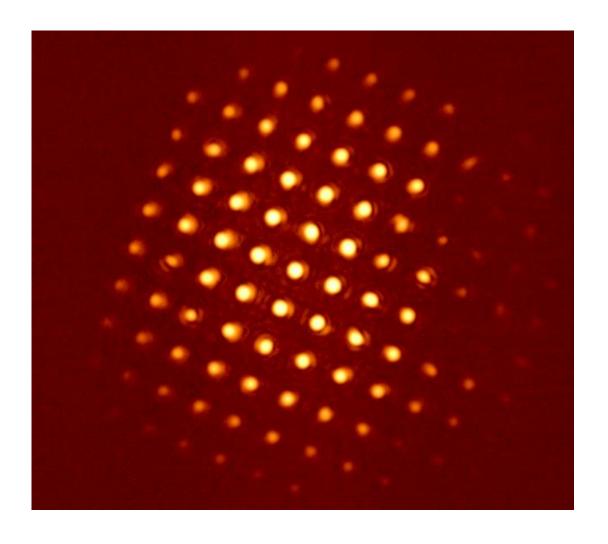
Chandrasekharan et al. 2017 (Bath)





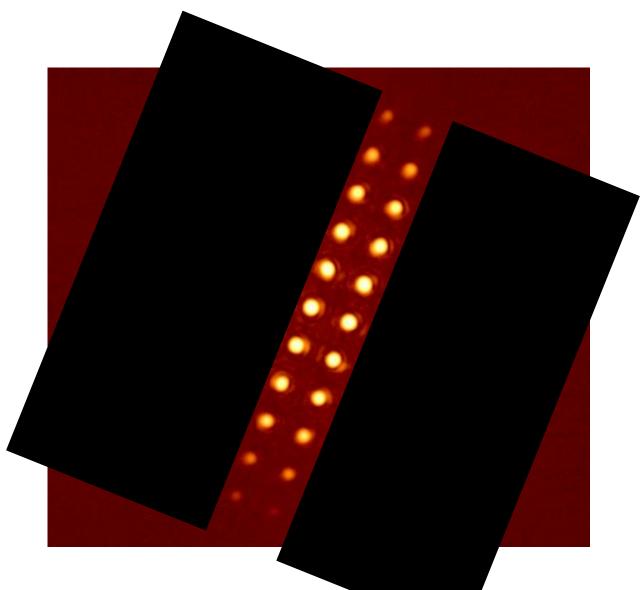


### Fiber core selection



- I added a slit because the 11x11 MCF has to many fibers.
- Due to the pitch to core ratio of this fiber I can only disperse two columns. There is no room for a third
- So we get an 2x11 area on-sky. This is roughly ~ 0.45x2.5 arcseconds projected on-sky.

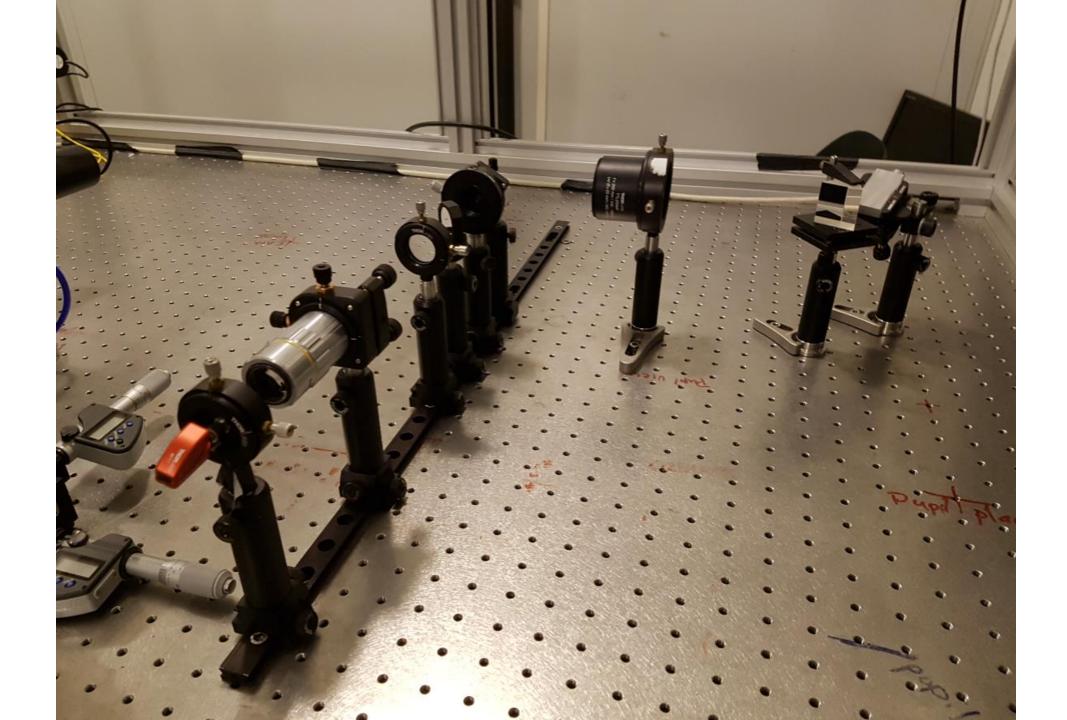
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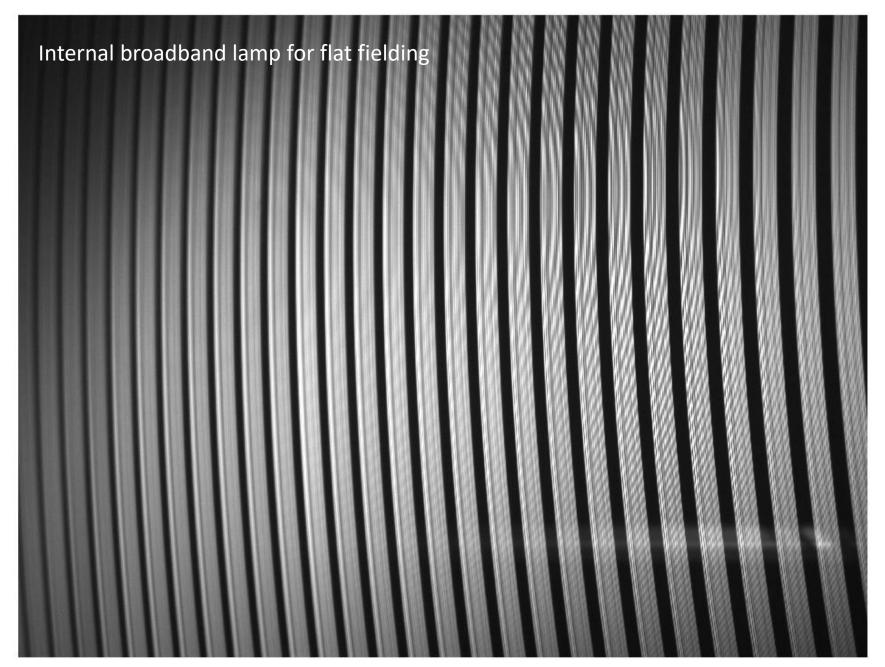
 So we get an 2x11 area on-sky. This is roughly ~ 0.45x2.5 arcseconds projected on-sky.



#### Latest LEXI run at WHT 2018

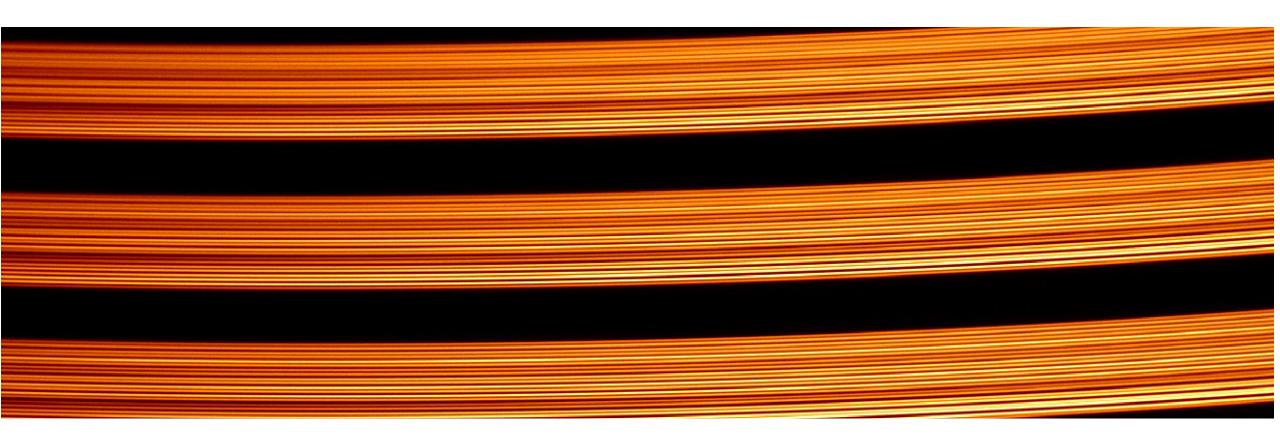
• Observing from 23<sup>rd</sup> – 31<sup>st</sup> of December 2018

• Telescope broke down during our run. We only had 2 half nights.

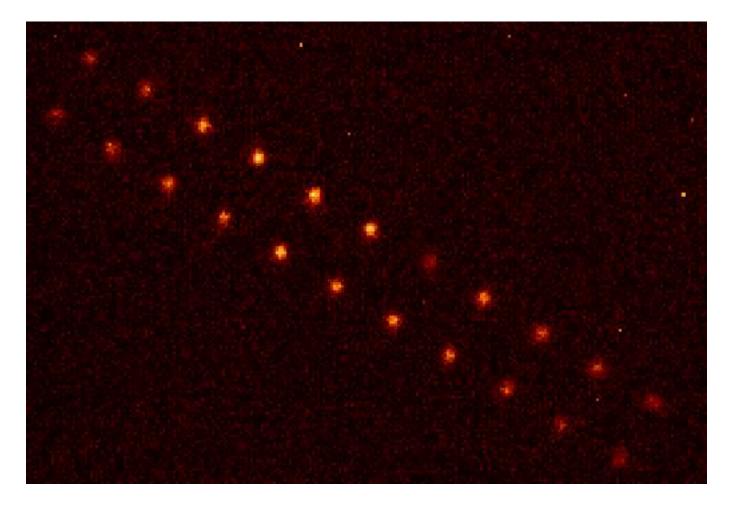


Haffert et al. in prep

# Zoom in of the echellogram

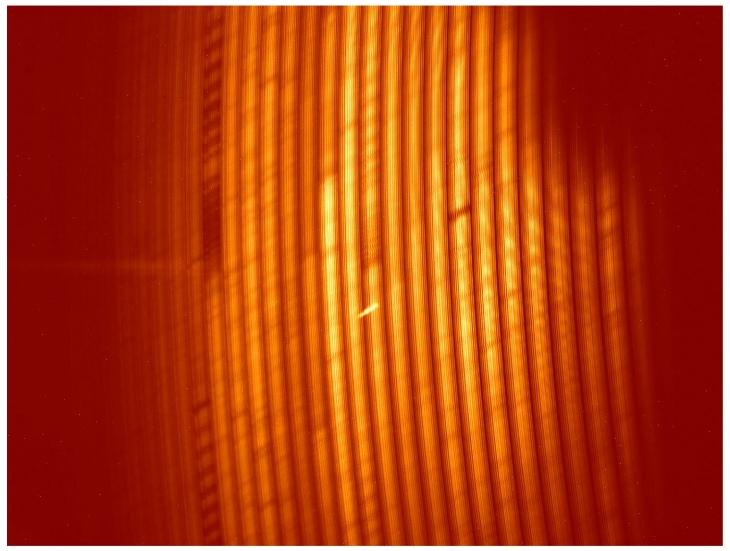


## Imaging through the fiber array on Regulus



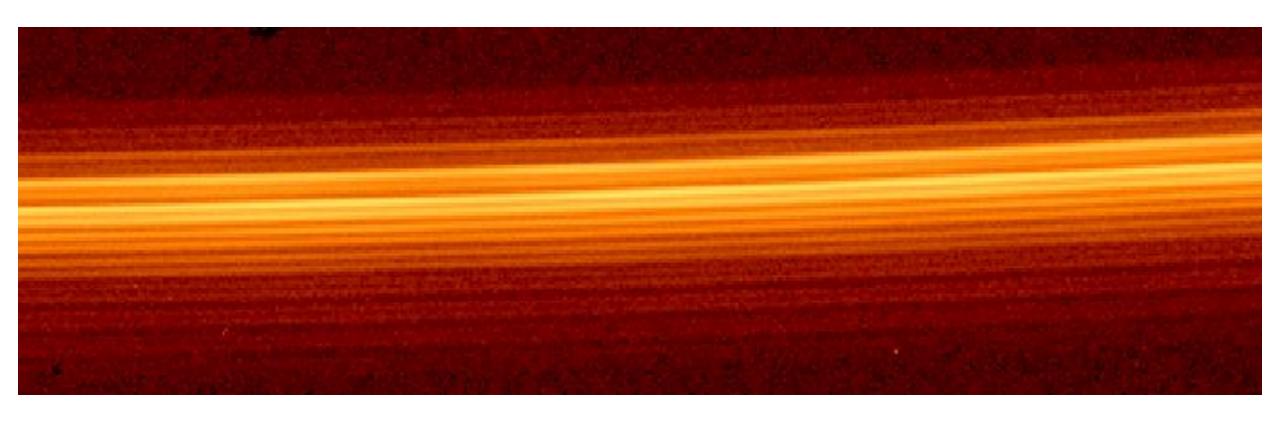
Haffert et al. in prep

# A spectrum of Betelgeuse



Haffert et al. in prep

# A spectrum of Betelgeuse zoom-in



### Conclusion

- LEXI has been developed over the past 3.5 years
- The new AO strategy works and delivers high quality PSFs

• Multi-core fibers with the 3D printing works very well for light injection

With the MCF we can make compact high-resolution IFUs